

Open strings in electric fields and time dependent backgrounds

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Introduction

- Much effort in string theory has been directed into searching for compactifications to **flat Minkowski space** which reproduce the Standard Model at low energies. Alas, $t_{LHC} > 2008$, and chances to observe strings directly are moderate.
- In contrast, **observational cosmology** is undergoing a fast revolution, from an order-of-magnitude Regime to a **high-precision** Era, posing a new challenge to string theory:

$$\omega_{\Lambda} = 71.0\% , \quad \omega_{baryon} = 4.7\% , \quad \omega_{dark} = 24.3\%$$

- While string-inspired cosmological scenarios have been much discussed in effective field theory, string theory in **time-dependent backgrounds** remains a mostly uncharted territory.

Strings in time-dependent backgrounds

Perturbative string theory is well-suited for **S-matrix** computations in **asymptotically flat** space.

Many questions arise in trying to generalize to (smooth) time-dependent backgrounds:

- **No (unitary) analytic continuation to Euclidean signature**, neither in **target space** nor on the **worldsheet**: amplitudes are superficially divergent, modular group acts ergodically...
- **Many different choices of vacuum** are possible, how can one implement **Bogolioubov transformations** from one to another ? Is worldsheet locality sacred ?
- **Observables are unclear**, especially in the case of **closed** universes, or with **pathological asymptotic regions** like such as the Cheshire's Cat Universe and its whiskers.

String Field Theory seems a crying need in order to address these issues.

Strings at cosmological singularities

More questions arise in relation with spacelike singularities, which a purported theory of quantum gravity had better address:

- Can perturbative string theory still hold, despite the **infinite blueshift** towards the singularity ?
- Can extra degrees of freedom of string theory **resolve spacelike singularities**, or rather **prevent their appearance** ? How can one evade the no-bounce theorem ?
- If instead spacelike singularities signify the End or Beginning of time, how can one specify **boundary conditions** there ?
- Is the **BKL oscillatory behaviour** generic also in string theory ? As different bits of the string fall outside of causal contact at the spacelike singularity, does the string reduce to a **Matrix model** ?

Cosmological Singularities: a Toy Story

Various toy models have been proposed recently to study time-dependence and cosmological singularities in string theory:

- The **Lorentzian orbifold**, quotient of $R^{1,1}$ by a boost J : this gives a free-field realization of the **Milne Universe**

$$ds^2 = -dt^2 + t^2 dx^2, \quad x \equiv x + 2\pi$$

together with two **whiskers** with CTC,

$$ds^2 = -r^2 dt^2 + dr^2, \quad t \equiv t + 2\pi$$

Horowitz Steif; Seiberg; Nekrasov

- The **Parabolic orbifold**, quotient of $R^{1,2}$ by the product of a boost J_{01} and a rotation R_{12} ,

$$ds^2 = -2dy^+ dy^- + (y^+)^2 dy^2, \quad y \equiv y + 2\pi$$

which is better thought of as a singular **gravitational wave**.

Simon; Liu Moore Seiberg

- Flux branes and null branes, where the boost is combined with a **translation** on an extra coordinate, hence lifting any fixed point; **WZW models** such as the Nappi Witten cosmology, which reduces to the Lorentzian orbifold at the singularity.

*Cornalba Costa; LMS; Craps Kutasov Rajesh
Elitzur Giveon Kutasov Rabinovici*

Toys are broken

- These models all seem to be plagued with **perturbative divergences**, related to a **large backreaction** at the singularity. Divergences may be avoided by fine-tuning initial conditions.

*Liu Moore Seiberg
Berkooz Craps Kutasov Rajesh*

- In addition, due to **high blue-shift**, the images of the particles on the covering space may non-perturbatively form a **large black hole**, that eats up the space. Combining with a translation does not cure this instability except in high dimension.

Horowitz Polchinski

- These models are also highly **non-generic** trajectories on the **cosmological billiard**: can one study more general Kasner singularities ? find the BKL behaviour ?

Damour Henneaux

More toys: open strings in electric fields

For the purpose of studying time-dependence in string theory, it may be simpler to consider **time-dependent D-brane** configurations, or equivalently **open strings in electric fields**:

- Backreaction in the **closed** string sector may be neglected as $g_s \rightarrow 0$. Yet production of **open** strings is retained. Backreaction in the open string sector is analogous to **D-brane recoil**.
- Powerful techniques are available: boundary states, string field theory ... Classical configurations can often be found explicitly due to the fact that the worldsheet theory is **free in the bulk**.
- Analogues of spacelike singularities are **D-brane collisions**, or (in the simplest case) a **constant electric field**. Analogues of null singularities are **null scissor configurations**, or a **constant null field**.

Bachas Hull

- The analogy is very precise: charged open strings in an electric field have (half) the same **mode structure** as twisted closed strings in a Lorentzian orbifold. Issues of **physical states** can be dealt in the same way.

Outline

1. Introduction
2. Open strings in constant electromagnetic field
3. Open strings in electromagnetic waves
4. Open strings in a constant electric field, revisited
5. Remarks on Milne universe

Open strings in a constant electromag field

- Open strings couple to an electromagnetic field through their **boundary** only. The embedding coordinates are therefore **free bosons** in the bulk of the Minkowskian strip $0 < \sigma < \pi$, $\tau \in R$,

$$X^\mu(\tau, \sigma) = f^\mu(\tau + \sigma) + g^\mu(\tau - \sigma)$$

- The electric field may be different on each of the two D-branes. The boundary conditions at $\sigma = \sigma_a \in \{0, \pi\}$

$$\partial_\sigma X^\mu + (2\pi\alpha') F^\mu{}_{\nu;a}(X) \partial_\tau X^\nu = 0$$

- For a constant F , this is a linear system of **non-local ODEs**,

$$\begin{aligned} \dot{f}^\mu - \dot{g}^\mu + (2\pi\alpha') F^\mu{}_{\nu;0} (\dot{f}^\nu + \dot{g}^\nu) &= 0 \\ T \dot{f}^\mu - T^{-1} \dot{g}^\mu + (2\pi\alpha') F^\mu{}_{\nu;1} (T \dot{f}^\nu + T^{-1} \dot{g}^\nu) &= 0 \end{aligned}$$

where $\cdot = d/d\tau$ and $Tf(\tau) = f(\tau + \pi)$.

- This can be solved in Fourier space, $T = e^{-i\pi\omega}$, $\partial/\partial\tau = -i\omega$. Eigenmodes satisfy, assuming $[F_0, F_1] = 0$,

$$e^{-2\pi i\omega_n} = \frac{1 + F_0}{1 - F_0} \cdot \frac{1 - F_1}{1 + F_1}$$

Open strings in a constant electromag field

The dispersion relation again:

$$T^2 = e^{-2\pi i \omega_n} = \frac{1 + F_0}{1 - F_0} \cdot \frac{1 - F_1}{1 + F_1}$$

- Magnetic field: $F = b \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \rightarrow \{ib, -ib\}$ hence $|T| = 1$ and frequencies are **real**:

$$\omega_n = n \pm \nu, \quad \pi\nu = \text{ArcTan } b_1 - \text{ArcTan } b_0$$

The string is stable and follows **Landau orbits**.

- Electric field: $F = e \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \rightarrow \{e, -e\}$ hence $|T| \neq 1$ and frequencies have an **imaginary part**:

$$\omega_n = n \pm i\nu, \quad \pi\nu = \text{ArcTanh } e_1 - \text{ArcTanh } e_0$$

This instability is due to **Schwinger pair production**:

$$A = (e_0 + e_1) \int_0^\infty \frac{dt}{t^{13}} e^{-\pi\nu^2 t} \frac{1}{\eta^{21}(it)\theta_1(\nu t | it)}$$

$$\Im(A) \sim \sum_{k=1}^{\infty} a(k) \exp\left(-\frac{\pi k}{\nu}\right)$$

Bachas Porrati

Born-Infeld critical electric field

- At the **critical electric field** $e_a = 1/\alpha'$, the electric force pulling the two ends of the string apart overwhelms the string tension, leading to the production of **stretched macroscopic strings**, that discharge the condensator at infinity.
- By scaling $e_a\alpha' \rightarrow 1$ and $\alpha' \rightarrow 0$ while keeping the effective tension of charged strings fixed, one obtains **NCOS**, a theory of interacting non-commutative open strings, decoupled from closed strings, propagating in a fixed open string metric.

*Gopakumar Maldacena Minwalla Strominger
Seiberg Sussking Toumbas*

- This classical instability occurs even for neutral dipoles. In contrast, the non-perturbative Schwinger instability requires charged particles.

Open strings in a null electric field

- A generic $F_{\mu\nu}$ can always be brought to the electric or magnetic form depending on $\text{sgn } F_{\mu\nu}F^{\mu\nu}$. However there is a non-generic possibility,

$$F = \Phi dx \wedge dx^+, \quad x^\pm = (x^0 \pm x^1)/\sqrt{2}, \quad x = x^2$$

which satisfies $F_{\mu\nu}F^{\mu\nu} = 0$. In 4D, it amounts to a configuration with **crossed** fields $\vec{E} \perp \vec{B}$ of **equal** magnitude $|\vec{E}| = |\vec{B}|$.

- The matrix F^μ_ν now has a non-trivial **Jordan form**, (the only non-trivial form for $SO(1, d-1)$)

$$F = \Phi \begin{pmatrix} 0 & 1 & \\ & 0 & 1 \\ & & 0 \end{pmatrix} \rightarrow \{0, 0, 0\}$$

hence the spectrum is **unaffected**, $\omega_n = n \in \mathbb{Z}$. Precise eigenmodes do depend on Φ however.

- This agrees with the fact that there is **no polarization** in a configuration with null electric field. In fact, this configuration preserves **half SUSY**, namely the generators such that $\Gamma^+ \epsilon = 0$.
- After T-duality on x , this describes a **null scissor** configuration, i.e. two intersecting straight D-branes whose intersection point moves with the speed of light.

Bachas Hull

Relativistic string in a pulse

- More generally, one may allow an arbitrary dependence in the **light-cone time** x^+ :

$$A = \Phi(x^+)x dx^+ , \quad F = \Phi'(x^+)dx \wedge dx^+$$

All contractions of F_{x^+} and ∂_+ vanish, hence this is an **exact supersymmetric solution** of the open strings eom to all orders in α' .

- In light-cone gauge $X^+ = x_0^+ + p^+\tau$, the boundary conditions receive a **time-dependent source term**,

$$\partial_\sigma X + (2\pi\alpha')p^+\Phi'_a(X^+) = 0 , \quad \sigma = \sigma^a \in \{0, \pi\}$$

while X is still free in the bulk, $(\partial_\tau^2 - \partial_\sigma^2)X = 0$. Classical solutions can be computed by **linear response**.

- Assuming that the electric field vanishes at $x^+ = \pm\infty$, it is now straightforward to compute the quantum mechanical **transition amplitudes**. An incoming string in its ground state will in general emerge in an excited state, depending on the profile $a(x^+)$.
- After T-duality along x , the bc becomes

$$X(\tau) = -(2\pi\alpha')\Phi^{(a)}(X^+(\tau)) + b^{(a)}$$

It describes open strings stretched between two D-branes with a null intersection: **null scissors**

Bachas

Exact travelling waves

- As a matter of fact, the class of electromagnetic waves which are exact solutions of open string theory is much larger:

$$A = \Phi(x^+, x^i) dx^+ , \quad F = \partial_i \Phi(x^+, x^i) dx^i \wedge dx^+$$

Maxwell's equations require that Φ be an **harmonic** function in transverse space.

- The corresponding open string metric is a **gravitational wave** in Brinkmann coordinates:

$$ds^2 = 2dx^+ dx^- + G_{ij} dx^i dx^j + (2\pi\alpha')^2 |\partial_i \Phi(x^+, x^i)|^2 (dx^+)^2$$

- In light-cone gauge $X^+ = x_0^+ + p^+ \tau$, the boundary conditions read

$$\partial_\sigma X^i + (2\pi\alpha') p^+ \partial_i \Phi(X) = 0$$

Just like closed strings in pp-waves, conformal invariance is broken in the light-cone gauge, but only through **boundary effects**.

- A **constant** magnetic field B_{ij} can be added, at the cost of using the open string metric in the harmonicity equation.

Solvable travelling waves

- The simplest harmonic solution is a **linear** potential $\Phi(x^+, x^i) = \phi_i(x^+)x^i$, leading to the uniform null field already discussed.

- The next simplest case is a **quadratic** potential

$$\Phi(x^+, x^i) = h_{ij}(x^+)x^i x^j / 2$$

leading to a **massive** linear boundary condition:

$$\partial_\sigma X^i + p^+(h_a)_{ij}(x^+)X^j, \quad \sigma = \sigma^a$$

- This is very reminiscent of studies of **open string tachyon condensation** in BSFT. However,
(i) due to the tracelessness of h , the boundary deformation $p^+ \oint \Phi(x^+, x^i) dX^+$ is **unbounded** from below or above.
(ii) the worldsheet is a **Lorentzian strip**, instead of an Euclidean cylinder or annulus. *Can tachyon dynamics be derived from Born-Infeld ?*

*Witten, Shatashvili; Arutyunov Pankiewicz Stefanski,
Bardakci Konechny*

- As for gravitational waves, supersymmetric **non-conformal** boundary deformations, in particular integrable, can be used to construct **on-shell** exact backgrounds.

Maldacena Maoz

A word on T-duality

- In terms of the T-dual coordinate

$$\tilde{X}^i = f^i(\tau + \sigma) - g^i(\tau - \sigma)$$

the bc become, after differentiating once,

$$\partial_\tau^2 \tilde{X}^i + p^+ (h_a)_{ij} \partial_\sigma \tilde{X}^j = 0,$$

This is an open string with two beads of mass h_a^{-1}/p^+ at its ends.

- This corresponds to a boundary deformation $(h^{-1})_{ij} \oint X^i \partial_\tau^2 X^j / p^+$ by an excited state. Deformations by more general excited states $X \partial^n X$ are also solvable.
- When $h = 0$, this is a Dirichlet bc. However, at finite coupling, D0-branes have finite mass $1/g_s$, hence $h \sim g_s$: D-brane recoil can be taken into account by going off-conformality.
- We will momentarily predict an instability of the T-dual system, at a critical line $m_0 - m_1 = \alpha' p^+$: fast elastic rotator ?
- A T-dual-like but inequivalent bc would be to take

$$\partial_\tau X^i + p^+ \partial_i \Phi(X) = 0$$

The ends of the open string follow the gradient lines of Φ : we are back to null scissors of arbitrary shape.

Point particles in electromagnetic waves

- The action for a charged particle is

$$S = \int \left[\frac{1}{2e} (\partial_\tau X^\mu)^2 - e m^2 \right] d\tau + A_\mu dX^\mu$$

- After choosing the gauge $e = 1$, the eom read

1. $(d^2/d\tau^2)X^+ = 0$
2. $(d^2/d\tau^2)X^i + \partial_i \Phi \partial_\tau X^+ + B_{ij} \partial_\tau X^j = 0$
3. $(d^2/d\tau^2)X^- - \partial_i \Phi \partial_\tau X^i = 0$

- 1. can be integrated to $X^+(\tau) = x_0^+ + p^+ \tau$. 2. and 3. imply that

$$H = \frac{1}{2}(p_i)^2 + p^+ p^- + p^+ \Phi(X^+, X^i) + m^2$$

is a constant, where $p_i = \partial_\tau X^i - B_{ij} X^j$ and $p^- = dX^-/d\tau - \Phi$ are the canonical momenta conjugate to X^i and to X^+ .

- The motion in transverse coordinates is therefore that of a **non-relativistic** particle in an electrostatic potential $V = \Phi(X^+, X^i)$.
- Similarly, a relativistic string in an electromag wave behaves as a non-relativistic **elastic dipole** (possibly with overall charge)

Non-relativistic dipole and critical gradient

- We have seen that on the light-cone, a relativistic string behaves like a non-relativistic dipole. This implies that its tensile energy is proportional to the **square** of its length:

$$V_t = \frac{1}{\alpha' p^+} (x_L - x_R)^2$$

- For a quadrupolar wave, the electrostatic energy scales also like the square of the distance,

$$V_e = p^+ \left(h_{ij}^0 x_L^i x_L^j - h_{ij}^1 x_R^i x_R^j \right)$$

- There exists therefore a line of **critical electric gradients**

$$h_1 - h_0 - \pi (p^+)^2 \alpha' h_0 h_1 = 0$$

at which the two forces balance against each other, leading to **stretched macroscopic strings**. This is the non-relativistic analogue of the Born-Infeld critical field. *What is the analogue of the open string metric ?*

- One may wonder if a decoupled theory of **non-relativistic interacting open strings** exists at that point, analogue to NCOS. This theory would exhibit **light-like non-commutativity**.

Aharony Gomis Mehem

First quantization

- Since the bulk theory is still free, one may separate X into left and right movers,

$$X^i = f^i(\tau + \sigma) + g^i(\tau - \sigma)$$

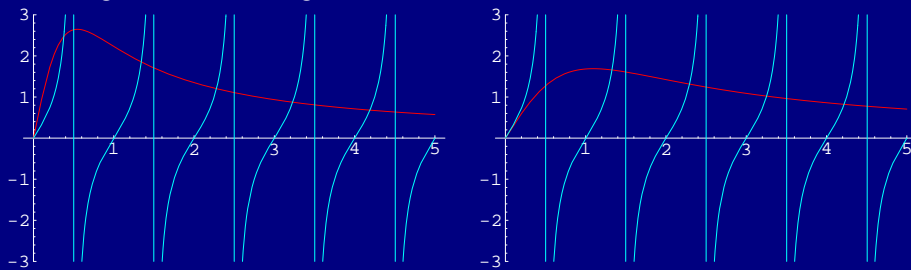
which satisfy boundary conditions:

$$\begin{aligned} \dot{f}(\tau) - \dot{g}(\tau) + p^+ h_0 (f(\tau) + g(\tau)) \\ T^2 \dot{f}(\tau) - \dot{g}(\tau) + p^+ h_1 (T^2 f(\tau) + g(\tau)) = 0 \end{aligned}$$

- Again, we can work in Fourier space, and find the dispersion relation ($e_i = \pi p^+ h_i$)

$$\tan(\pi\omega) = \frac{(e_1 - e_0)\pi\omega}{(\pi\omega)^2 + e_0 e_1}$$

Indeed, a pair of real roots disappear at the **critical line** $e_0 - e_1 - e_0 e_1 = 0$:



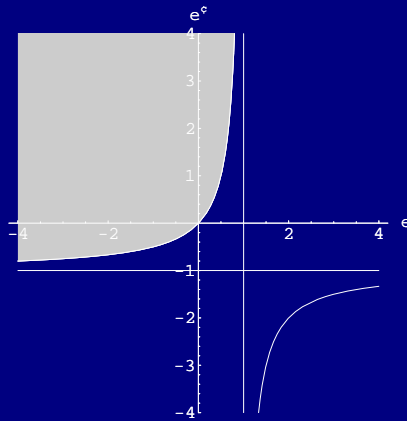
- The partition function for open strings in a quadrupolar field is then simply

$$Z_{op}(t, e_0, e_1) = q^{E_X} \prod_{n=1}^{\infty} (1 - q^{\omega_n})^{-1} \times \begin{cases} (1 - q^{\omega_0})^{-1} & \text{if } D > 0 \\ (1 - q^{ik_0})^{-1} & \text{if } D < 0 \end{cases}$$

with $E_X = \sum_{n=0}^{\infty} \omega_n$ the zero-point energy.

Dynamical instability

- For a quadratic potential depending on a **single** direction, the motion is stable in the shaded region, extending slightly outside the domain of stability of a dipole with vanishing tension:



- For a **traceless** quadratic potential h_{ij} , the motion is always unstable, due to the convexity of the stability domain. However, this is a **kinematical instability** of the string probes, not of the background itself: much like the divergence of geodesics in purely gravitational plane waves,

$$ds^2 = 2dx^+ dx^- + dx^2 + dy^2 - \mu^2(x^2 - y^2)(dx^+)^2$$

Marolf Zayas; Brecher Gregory Saffin

- (Former) atomic physicists know how to deal with these instabilities...

Strings in quadrupolar ion traps

Several ways to make a stable electromagnetic trap:

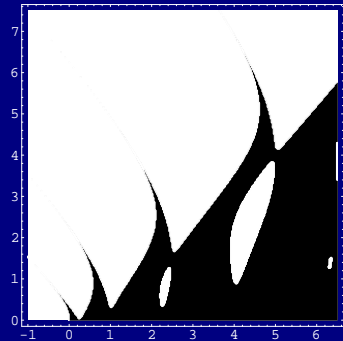
- a. **The Penning trap:** use a static magnetic field to confine charged particles in the transverse unstable plane:

$$V(x) = -\frac{e}{2}(x^2 + y^2 - 2z^2) , \quad B = b dx \wedge dy$$

is stable if $b^2 > e$ and $e > 0$.

- b. **The RF or Paul trap:** no magnetic field, but modulate the electric field at a frequency such that the particle experiences a restoring force on average: parametric resonance

$$V = (\omega^2 + \alpha^2 \cos t)(x^2 - y^2)$$



- c. **The quadrupolar trap:** a static quadrupolar potential confines neutral particles with a negative polarizability, by drawing them to regions of low electric field strength: $W = -\alpha E^2$. Degenerate excited states usually have negative polarizability

Mechanisms a. and b. carry over to the string case easily.

Closed string channel and boundary state

- In the closed string channel the boundary states satisfies

$$\partial_\tau X + \frac{\hat{e}}{\pi} X |B(\hat{e})\rangle = 0$$

This is solved by the usual **coherent state** techniques,

$$|B(\hat{e})\rangle = \mathcal{N}(\hat{e}) e^{i\frac{\pi p_0^2}{2\hat{e}}} \exp\left(\sum_{n=1}^{\infty} -\frac{1}{n} \frac{i\pi n + \hat{e}}{i\pi n - \hat{e}} \alpha_{-n} \tilde{\alpha}_{-n}\right) |0, \tilde{0}\rangle$$

- The partition function is therefore given by the **overlap** of the two boundary states,

$$Z_{cl}(\hat{t}, \hat{e}_0, \hat{e}_1) = \mathcal{N}(\hat{e}_0) \mathcal{N}(\hat{e}_1) \sqrt{\frac{2}{\hat{t} + i\left(\frac{1}{\hat{e}_1} - \frac{1}{\hat{e}_0}\right)}} e^{\pi\hat{t}/12}$$

$$\prod_{n=1}^{\infty} \left(1 - \frac{i\pi n + \hat{e}_0}{i\pi n - \hat{e}_0} \frac{i\pi n - \hat{e}_1}{i\pi n + \hat{e}_1} e^{-2\pi n\hat{t}}\right)^{-1}$$

Arutyunov Pankiewicz Stefanski, Bardakci Konechny

Open-closed duality

- Equality with the open string channel can be formally seen by representing the sum by a **residue integral**

$$\log Z_{op} = \frac{1}{2\pi} \int_C (\log \Phi_{cl}) \frac{d \log \Phi_{op}}{dz} dz$$

with

$$\Phi_{op}(z) = 1 - e^{-2\pi iz} \frac{i\pi z + e_0}{i\pi z - e_0} \frac{i\pi z - e_1}{i\pi z + e_1}, \quad \Phi_{cl}(z) = 1 - e^{-2\pi tz}$$

Integrating by parts shows that

$$Z_{op}(t, e_0, e_1) = Z_{cl}(\hat{t}, \hat{e}_0, \hat{e}_1)$$

where the deformation parameters are related by

$$\hat{e}_a = e_a t$$

in full agreement with open/closed duality of the one-loop amplitude (after compactification of the light-cone).

- A careful proof takes much more effort, but can be made along the lines of a similar computation in the context of open strings in gravitational waves.

Bergman Gaberdiel Green

Strings in time-dependent quadrupolar fields

- We now take $h_a(x^+)$ with finite support in x^+ . At $\tau \rightarrow \pm\infty$ we have **free field** mode expansions,

$$X = x_0 + p_0\tau + i \sum_{n \neq 0} \frac{2}{n} a_n \cos(n\sigma) e^{-in\tau}$$

and a similar expansion with primes at $\tau \rightarrow \infty$.

- The two sets of modes are related by a symplectic matrix, the **Bogoliubov transformation**:

$$\begin{pmatrix} x'_0 \\ p'_0 \\ a'_m \end{pmatrix} = \begin{pmatrix} \alpha & \beta & A_n \\ \gamma & \delta & B_n \\ \tilde{A}_m & \tilde{B}_m & B_{mn} \end{pmatrix} \begin{pmatrix} x_0 \\ p_0 \\ a_n \end{pmatrix}$$

- In the **Born approximation** ($h \ll 1$), the incoming state is a source for the outgoing perturbation, and one finds easily e.g.

$$B_{mn} = \delta_{mn} + \frac{i}{\pi^2 n} \int_{-\infty}^{\infty} (e_0 - (Te_1))(p^+\tau) e^{-i(n-m)\tau} d\tau$$

- The final excitation number of the mode n is

$$\langle 0_{in} | a'_{-m} a_m | 0_{in} \rangle = \sum_{n \neq 0} |B_{m,-n}|^2 + \dots$$

hence the total energy diverges if $h_a(x^+)$ has a delta function singularity.

Two zero-modes in one

- In fact, the open string zero-mode has an **ambiguity** which corresponds to the splitting between left- and right-movers:

$$x_0 = f_0 + g_0, \quad a = f_0 - g_0$$

a is the **position** of the T-dual D-brane, hence the value of the worldvolume $U(1)$ **gauge field** A_x on the original D-brane.

- In flat space, a can be changed by a gauge transformation hence has no physical meaning.
- In a time-dependent situation, this is no more the case: the difference $a(x^+ = +\infty) - a(x^+ = -\infty)$ is the electric field F_{+x} . In the Born approximation,

$$\delta f_0 - \delta g_0 = -\frac{1}{\pi} \int_{-\infty}^{\infty} e_0(p^+ \tau) X(\sigma = 0, \tau) d\tau$$

This is possibly the simplest computation of the **backreaction** of an open string on an electric background.

- Similar computations can be made in the **adiabatic** approximation, but keeping $h^a(\pm\infty)$ finite, as the limit $h \rightarrow 0$ is non adiabatic.

Electric field and Lorentzian orbifold

- Closed strings in the w -th **twisted sector** of the Lorentzian orbifold satisfy

$$X^\pm(\sigma + 2\pi, \tau) = e^{\pm\nu} X^\pm(\sigma, \tau) , \quad \nu = w\beta$$

Expanding in left and right movers, we have the normal mode expansion:

$$X_R^\pm(\tau - \sigma) = \frac{i}{2} \sum_{n=-\infty}^{\infty} (n \pm i\nu)^{-1/2} \alpha_n^\pm e^{-i(n \pm i\nu)(\tau - \sigma)}$$

$$X_L^\pm(\tau + \sigma) = -\frac{i}{2} \sum_{n=-\infty}^{\infty} (-n \mp i\nu)^{-1/2} \tilde{\alpha}_n^\pm e^{-i(-n \mp i\nu)(\tau + \sigma)}$$

- Upon identifying the oscillators

$$\alpha_n^\pm = a_n^\pm , \quad \tilde{\alpha}_{-n}^\pm = (a_n^\pm)^*$$

and setting

$$X_{open}^\pm = x^\pm + X_R^\pm(\tau - \sigma) + X_L^\pm(-\tau - \sigma) ,$$

this reduces to the open string mode expansion

$$X^\pm = x^\pm + ia_0^\pm \phi_0^\pm(\sigma, \tau) + i \sum_{n=1}^{\infty} [a_n^\pm \phi_n^\pm(\sigma, \tau) - h.c.]$$

where $\phi_n^\pm(\sigma, \tau) = (n \pm i\nu)^{-\frac{1}{2}} e^{-i(n \pm i\nu)\tau} \cos[(n \pm i\nu)\sigma]$

- Canonical commutation relations are

$$[\alpha_n^+, (\alpha_n^-)^*] = [\alpha_n^-, (\alpha_n^+)^*] = [\tilde{\alpha}_n^-, (\alpha_n^+)^*] = [\tilde{\alpha}_n^-, (\tilde{\alpha}_n^+)^*] = -1$$

No-physical state no-theorem

- The worldsheet Hamiltonian for open strings reads, after **normal ordering** ($a^\pm := a_0^\pm$)

$$L_0 = - \sum_{n=1}^{\infty} (n-i\nu) (a_n^+)^* a_n^- - \sum_{n=0}^{\infty} (n+i\nu) (a_n^-)^* a_n^+ + \frac{1}{2} i\nu(1-i\nu) + L_{int}$$

- Representing in a Fock space with vacuum annihilated by all $a_{n>0}^-$ and $a_{n>0}^+$, eigenstates have **imaginary** energy. This does not contradict hermiticity, since they also have zero norm ! Hence the physical state condition L_0 has no solutions.
- For Milne Universe it is the same story with tildas. The no-physical state statement is warranted by modular invariance of the one-loop amplitude.

Nekrasov

Scattering states and tunnelling

- Alas, this vacuum is the one obtained by **analytic continuation from Euclidean**, i.e. for strings in a **magnetic field** (disregarding time direction). There physical states are **Landau states**, corresponding to **discrete normalizable** eigenmodes of the harmonic oscillator (times a continuous degeneracy label).
- In the Minkowskian (electric) case, the harmonic oscillator becomes **inverted**, and the continued Landau states now have **imaginary** energy. However there is now a **continuum of delta-normalizable physical scattering states** with **real** energy.

Das Jevicki; Moore; Alexandrov Kazakov Kostov

- Physically, these represent electrons and positrons being deflected by the electric field. **Tunneling** through the barrier is just **induced Schwinger emission**, $e^- \rightarrow \mu e^- + (1 + \mu)e^+$.

Brezin Itzykson

- **Scattering amplitudes** may be computed in the Euclidean (magnetic) theory, after expanding the scattering states on the basis of (analytically continued) “Landau states” – and proper regularization.

Berkooz P., in progress

Rindler, Liouville, ...

- A constant electric field $F = e dx^+ \wedge dx^-$ preserves symmetries under boost. One may consider states of fixed boost momentum J , and use adapted Rindler coordinates $x^\pm = \pm e^{y \pm \eta}$ in the R region.
- Neutral massive particles have a rectilinear motion in Minkowski, but in Rindler coordinates experience a Liouville wall:

$$(dy/d\eta)^2 + 4m^2 e^{2y} = 0$$

Charged massive particles have an hyperbolic motion, but in Rindler coordinates experience a Liouville type potential barrier

$$(dy/d\eta)^2 + 4m^2 e^{2y} - (J + e e^{2y})^2 = 0$$

Tunnelling corresponds again to induced pair production or teleportation.

Narozhny Mur Fedotov

... Unruh and Milne ...

- **Winding closed strings** in the Lorentzian orbifold behave exactly as **massive charged particles** in Rindler space, with boost momentum fixed by the matching condition $wJ = N_L - N_R$.
- Fixing the boundary conditions in the whiskers region should lead to **Unruh-type** vacua, with **thermal** particle production. *A stringy Hawking radiation computation ?*
- Winding strings will be pair-produced and should **backreact** on the geometry so as to “discharge the condensator”: is there enough time for the cosmological singularity to take place ?
- Electric fields are full of promises for the study of **Time and String Theory...**