

Experimental Constraints to High Energy Hadronic Interaction Models using the Pierre Auger Observatory Part III

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QCD at Cosmic Energies VII, Chalkida, Greece
May the 17th 2016

Outline

- **Monte-carlo for Cosmic Ray analysis**

 - ➔ Sibyll 2.3

- **Muon Production Depth and hadronic interactions**

 - ➔ baryons

- **Nuclear Interactions**

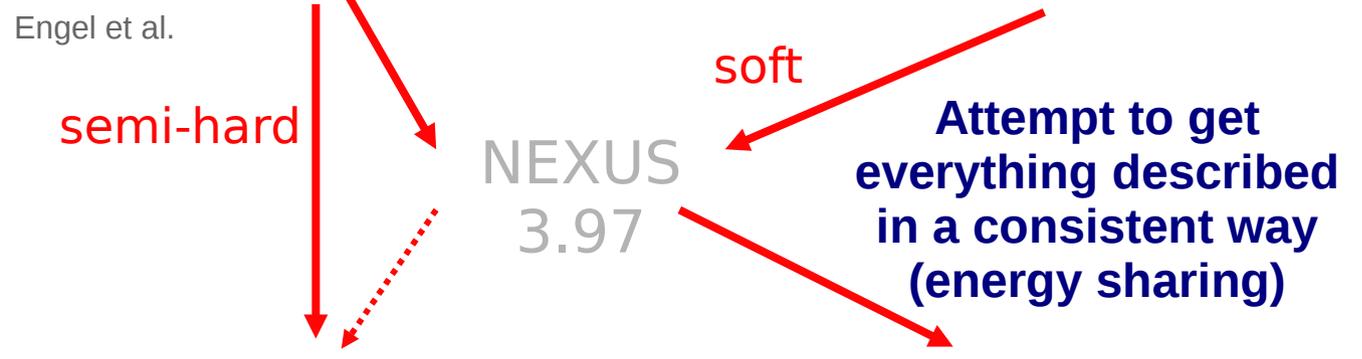
 - ➔ model differences

Hadronic Interaction Models in CORSIKA

(HDPM)

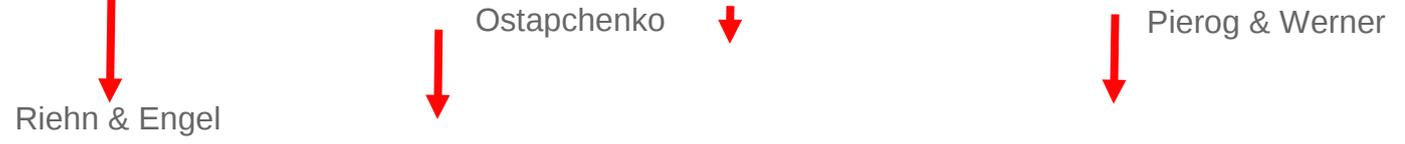
Old generation : SIBYLL 2.1 (QGSJET01 DPMJET 2.55 VENUS) (<2001)

All Glauber based
But differences in hard, remnants, diffraction ...



New generation : (QGSJET II-03)(DPMJET III) (EPOS 1.99) (2005-2012)

LHC tuned : QGSJET II-04 (2013-) EPOS LHC (2013-)



LHC inspired : SIBYLL 2.3 QGSJET III (?) EPOS 3 (2016-)

Motivation :
- update with latest LHC results in simple model

Motivation :
- Hard Pomeron-Pomeron connexion

Motivation :
- binary scaling in hard probes

Cosmic Ray Hadronic Interaction Models

● Theoretical basis :

- pQCD (large p_t)
- Gribov-Regge (cross section with multiple scattering)
- energy conservation

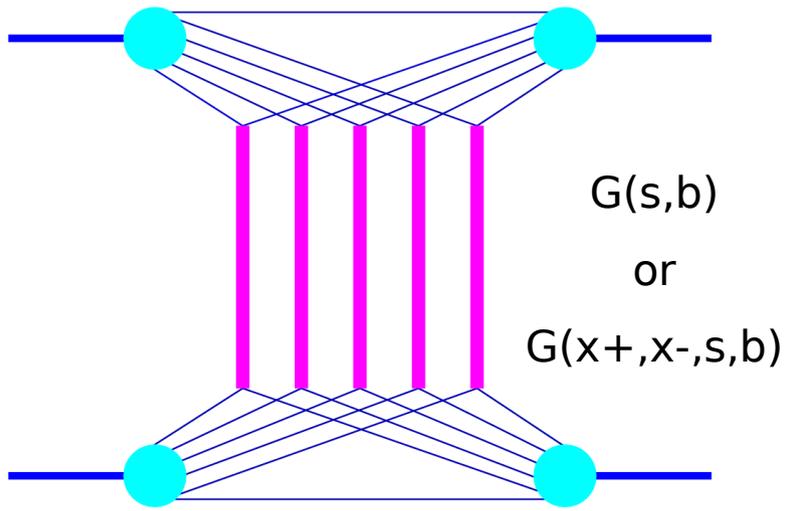
● Phenomenology (models) :

- hadronization
 - string fragmentation
 - EPOS : high density effects (statistical hadronization and flow)
- diffraction (Good-Walker, ...)
- higher order effects (multi-Pomeron interactions)
- remnants

● Comparison with data to fix parameters

Better predictive power than HEP models thanks to link between total cross section and particle production (GRT) tested on a broad energy range (including EAS)

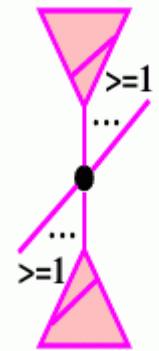
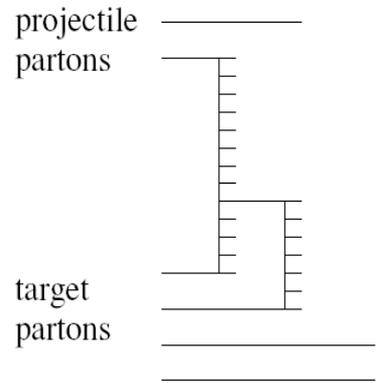
Cross Section and Multiplicity in Models



- **Gribov-Regge and optical theorem**
 - ➔ Basis of all models (multiple scattering) but
 - Classical approach for QGSJET and SIBYLL (no energy conservation for cross section calculation)
 - ◆ Parton based Gribov-Regge theory for EPOS (**energy conservation at amplitude level**)

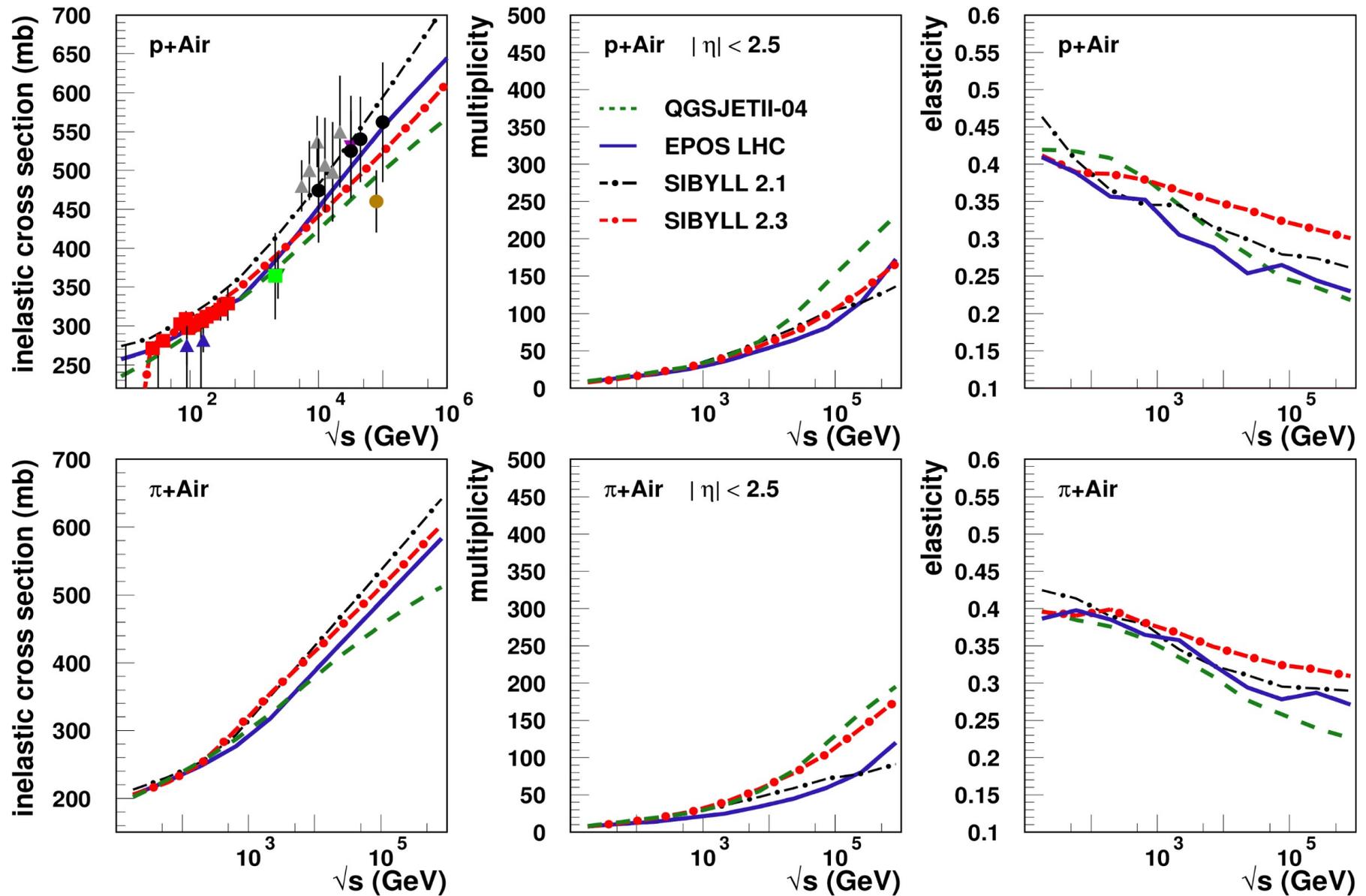
EPOS

QGSJET II

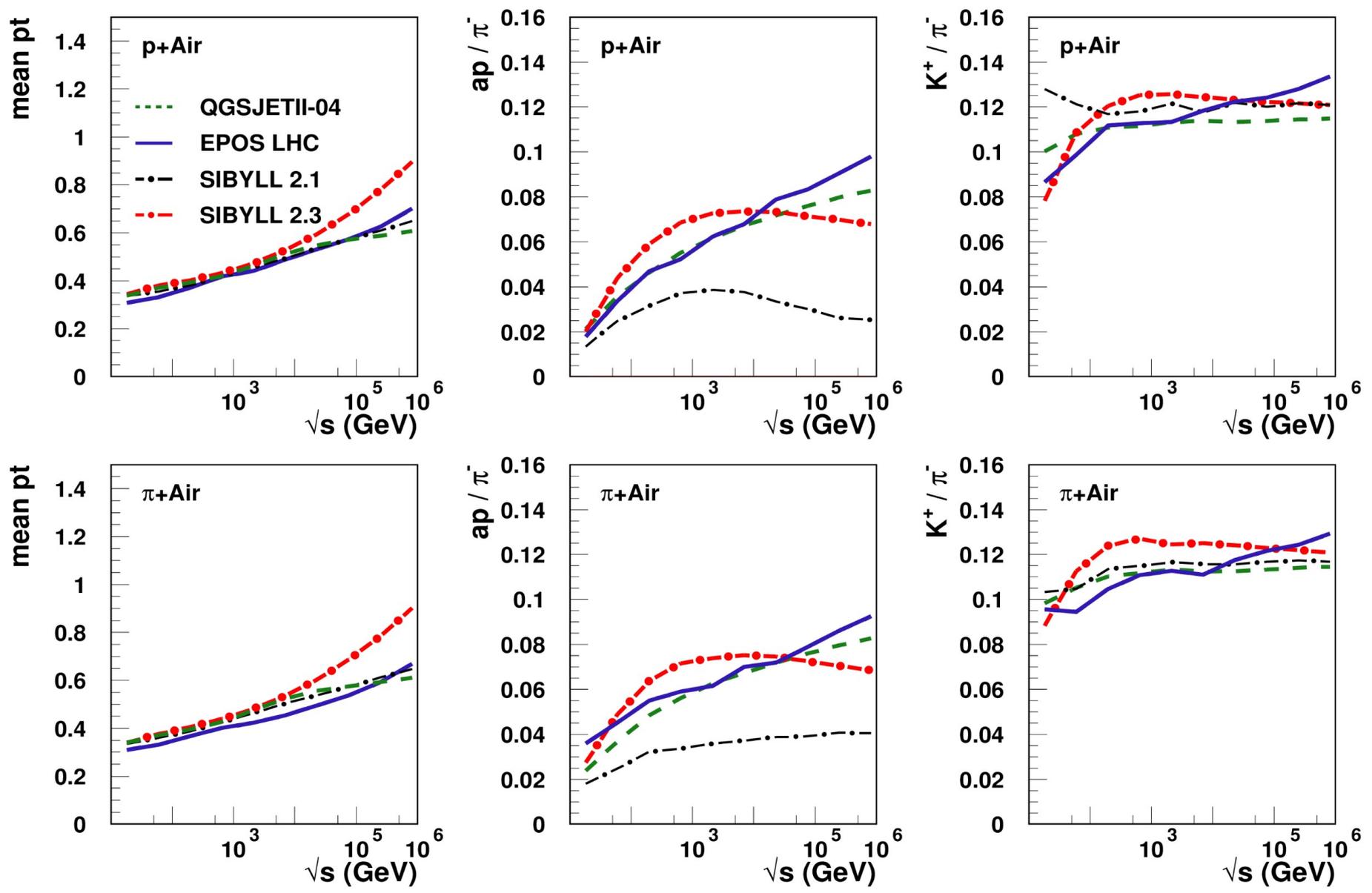


- **pQCD**
 - ➔ Minijets with cutoff in SIBYLL
 - ➔ Same hard Pomeron (DGLAP convoluted with soft part : no cutoff) in QGSJET and EPOS but
 - ◆ Generalized enhanced diagram in QGSJET-II
 - ◆ Simplified non linear effect in EPOS
 - Phenomenological approach

Model Predictions (1)



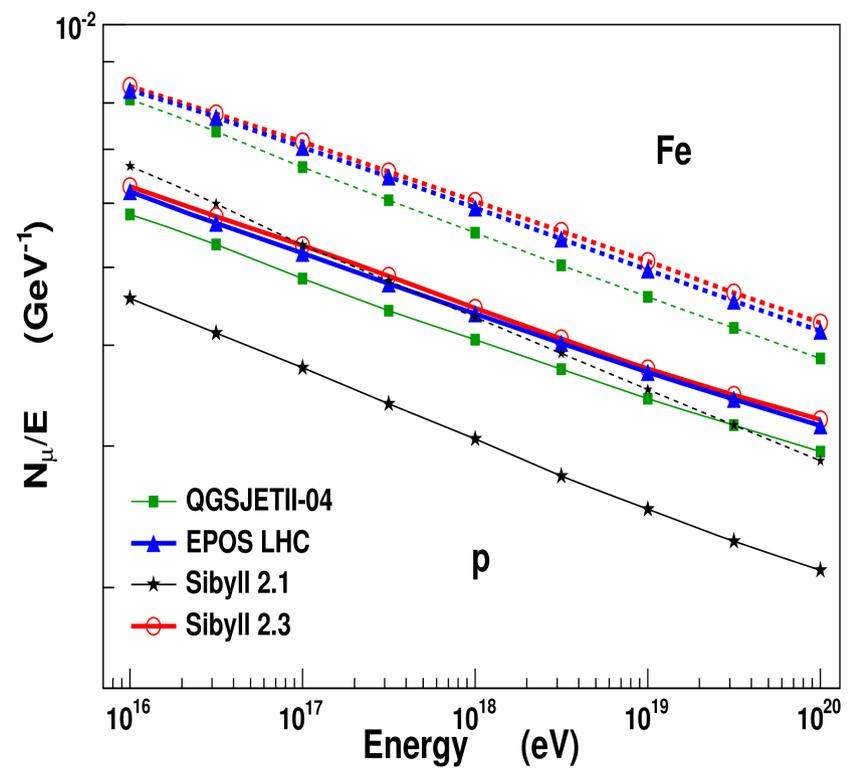
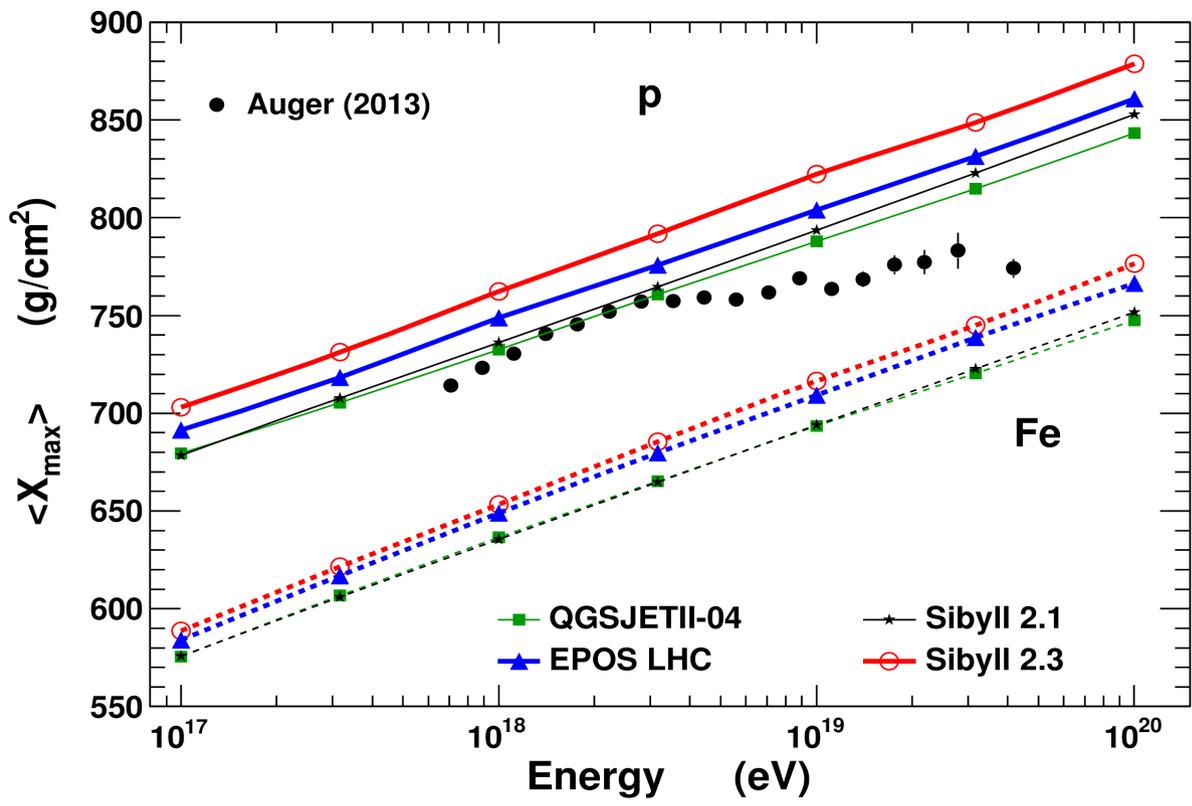
Model Predictions (2)



Air Shower Observables

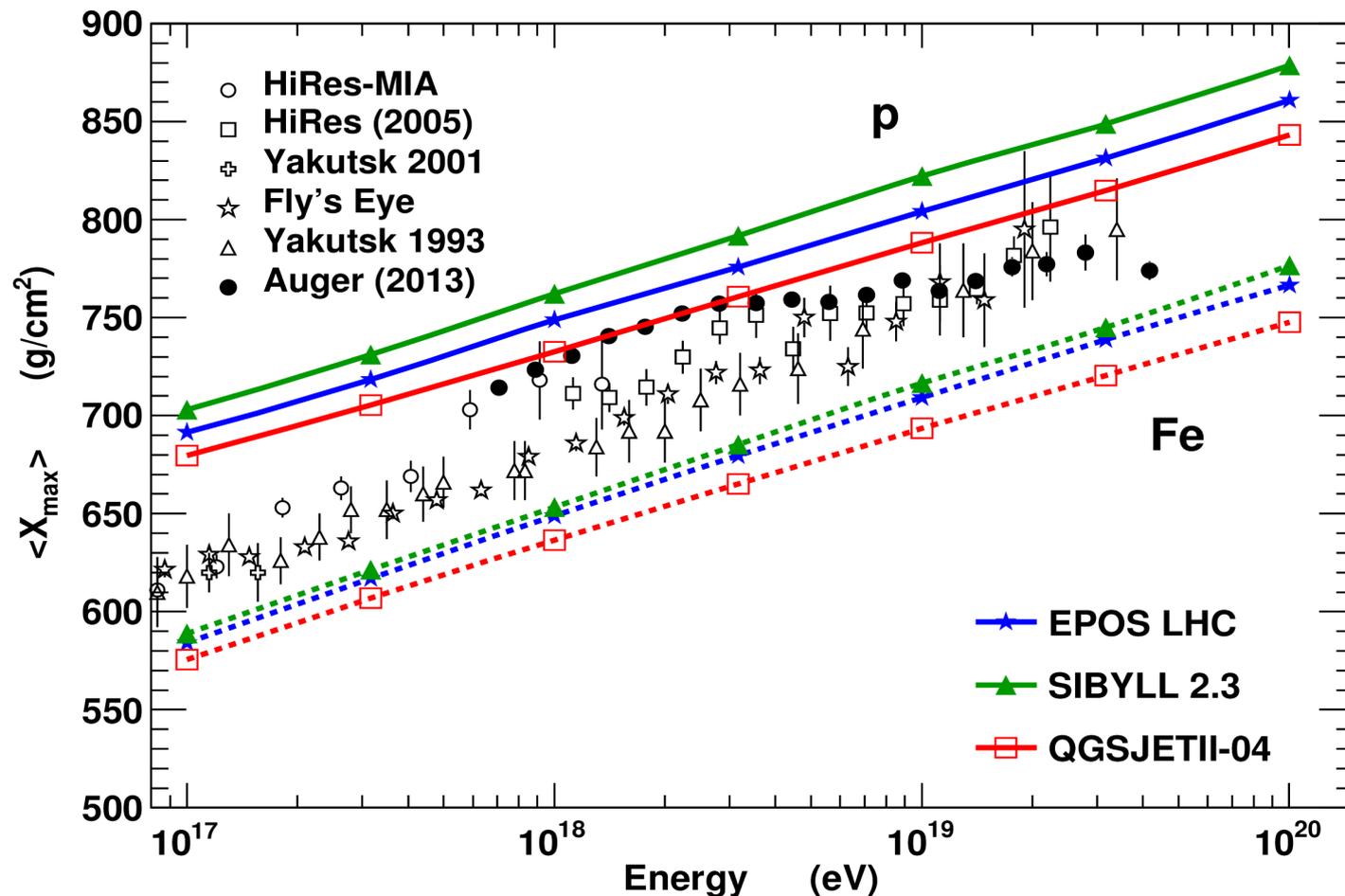
Post-LHC models have very similar energy evolution for X_{\max} and N_{μ} and small difference in absolute value but

- ➔ Sibyll 2.3 have quite large X_{\max} for proton
- ➔ different muon spectra between models



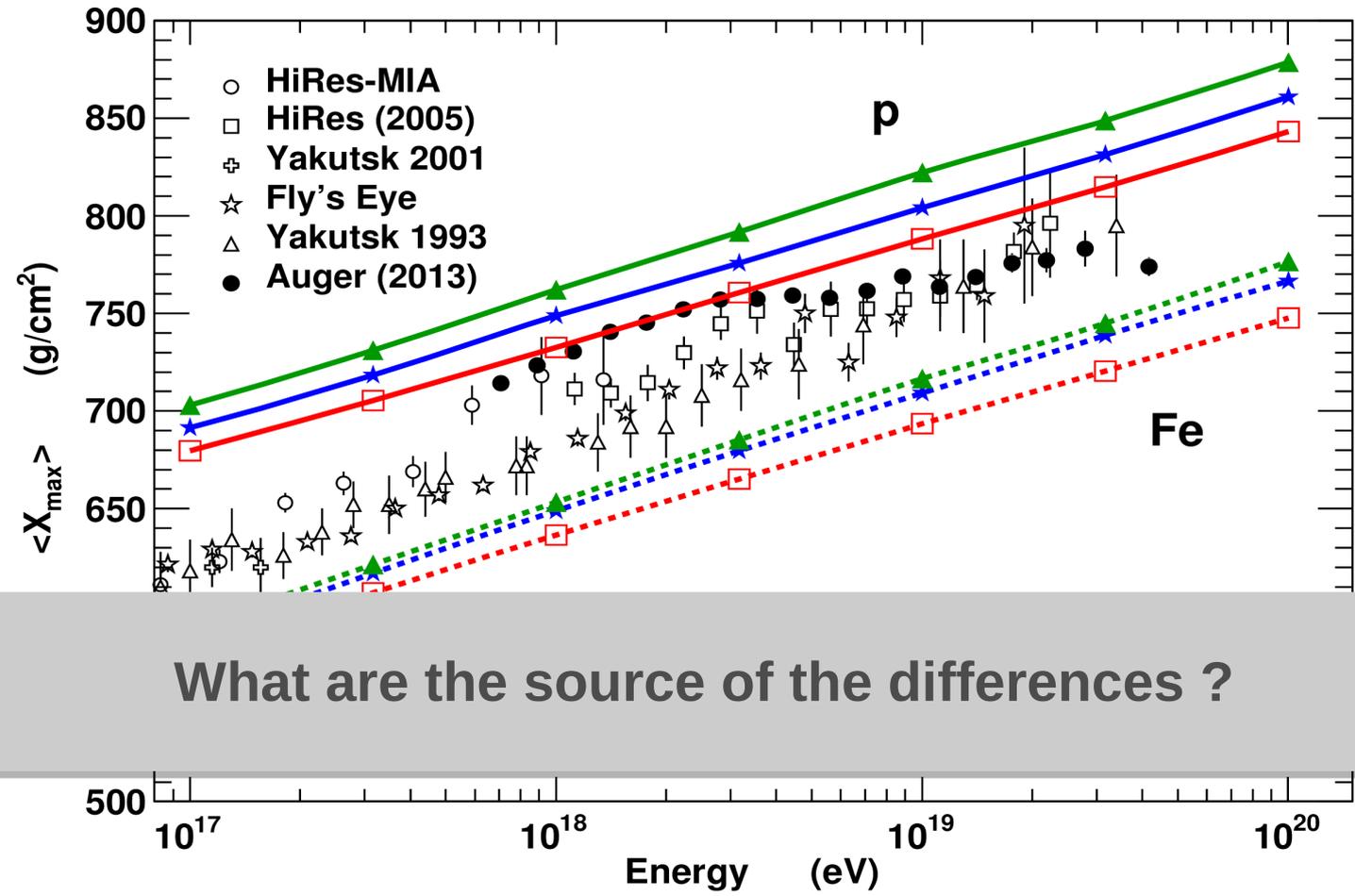
X_{\max} Model Uncertainties

After LHC still about 20g/cm^2 (40g/cm^2) difference between EPOS LHC (Sibyll 2.3) and QGSJETII-04 while only $\sim 10\text{g/cm}^2$ by changing p-Air within LHC uncertainties (see S. Ostapchenko, Phys. Rev. D 89, 074009 (2014))



X_{max} Model Uncertainties

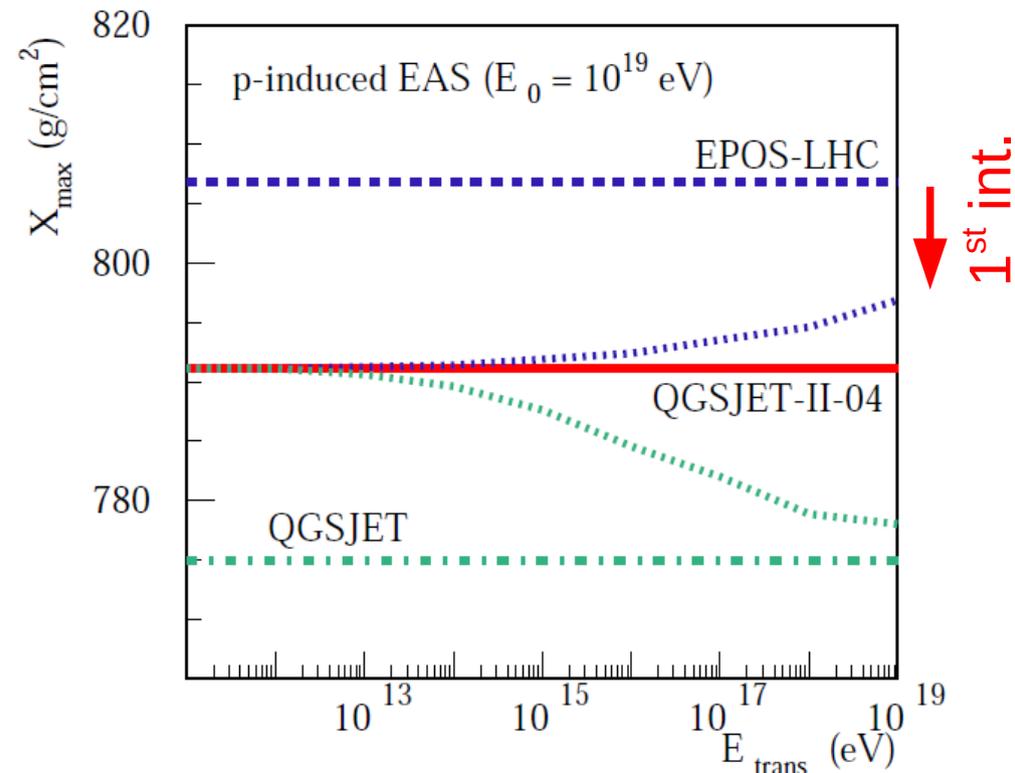
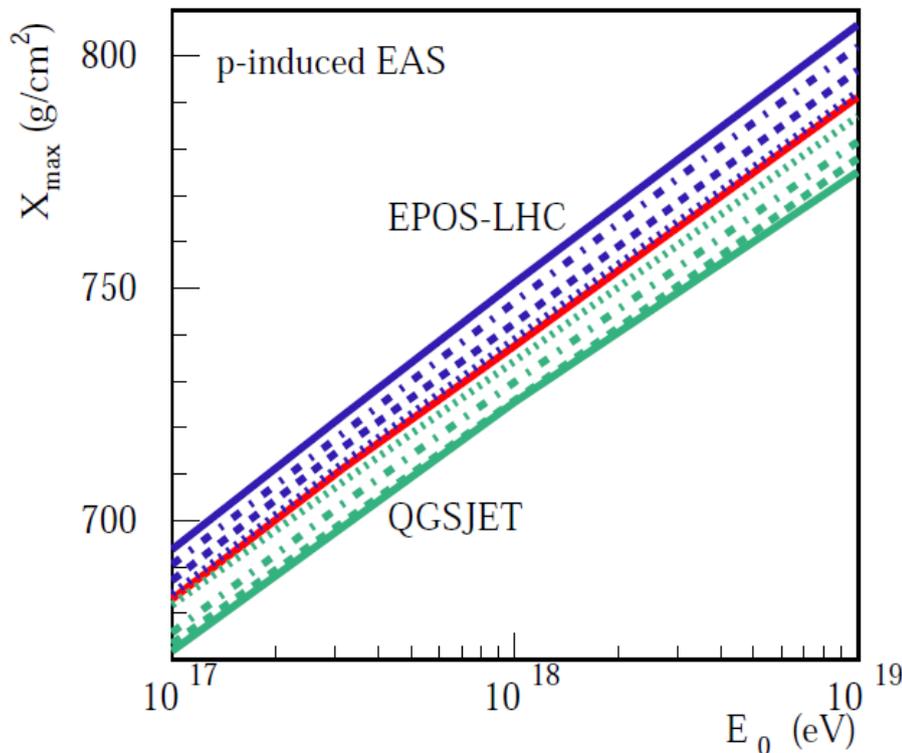
After LHC still about 20g/cm² (40g/cm²) difference between EPOS LHC (Sibyll 2.3) and QGSJETII-04 while only ~10g/cm² by changing p-Air within LHC uncertainties (see S. Ostapchenko, Phys. Rev. D 89, 074009 (2014))



What are the source of the differences ?

Mix models to identify source of differences

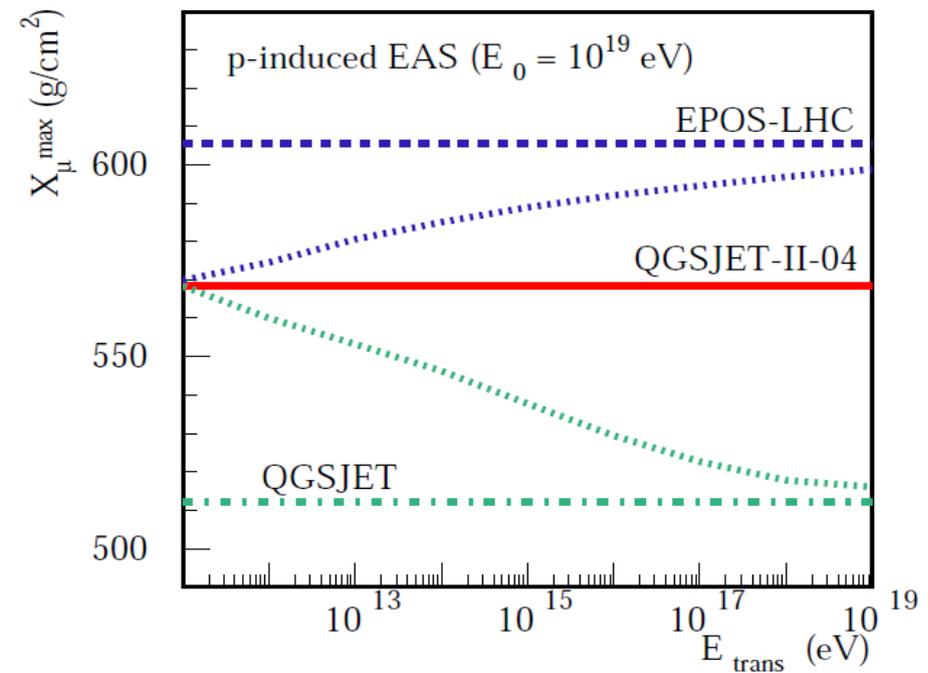
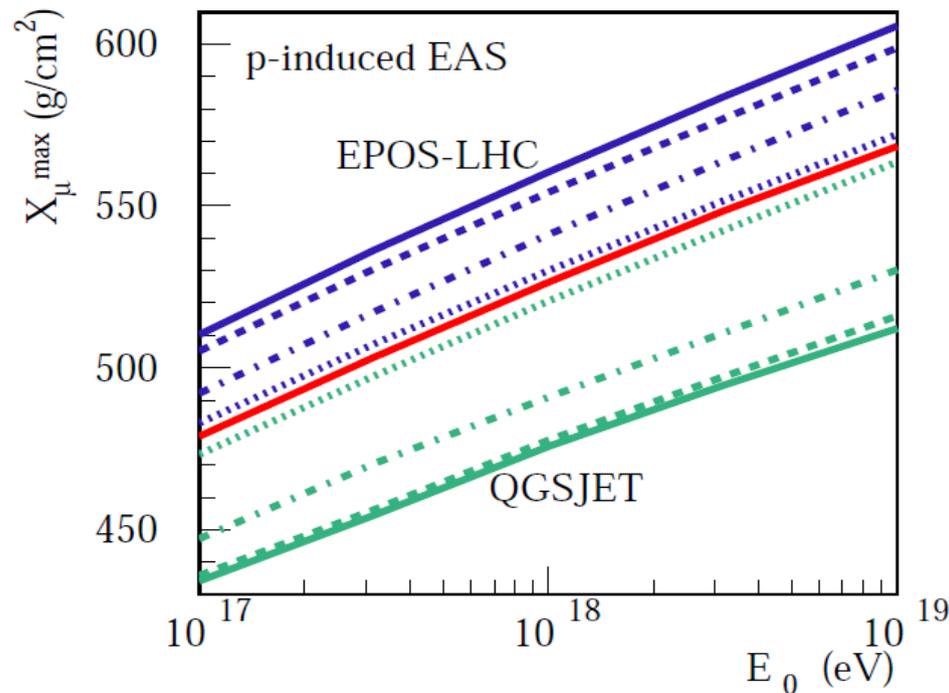
- Different mixing to extract useful information on X_{\max}
 - ➔ QII only for cross-section and nucleon spectra of 1st int. : dot-dashed
 - ➔ QII complete 1st int : dashed
 - ➔ QII complete 1st int and all nucleon prod. in the shower: dotted
 - ➔ For energy dependence, QII for $E > E_{\text{trans}}$, other model below



From arXiv:1601.06567 by S. Ostapchenko and M. Bleicher

Mix models to identify source of differences

- Different mixing to extract useful information on X_{μ}^{\max}
 - ➔ QII complete 1st int. : dashed
 - ➔ QII complete 1st int. and all nucleon prod. in the shower: dot-dashed
 - ➔ QII complete 1st int. and hadron spectra in pion and kaon int.: dotted
 - ➔ For energy dependence, QII for $E > E_{\text{trans}}$, other model below



From arXiv:1601.06567 by S. Ostapchenko and M. Bleicher

Summary of arXiv:1601.06567

Modifications	X_{\max}	X_{\max}^{μ}
cross-section and nucleon spectra of 1 st interaction	5 g/cm ²	
rest of 1 st interaction	5 g/cm ²	5 g/cm ²
nucleon spectra in all int.	5 g/cm ²	15 g/cm ²
all pion and kaon interactions		15 g/cm ²
Model difference fractions		
1 st interaction	70%	10%
pion interactions	30%	90%

Conclusions on Hadronic Interactions

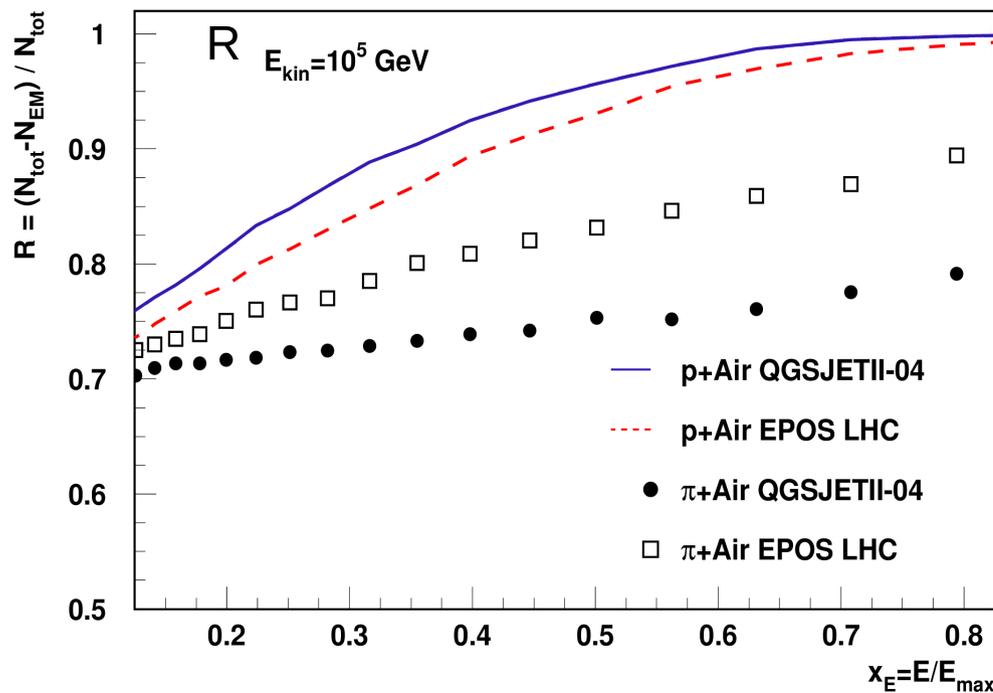
- Differences in first interaction dominates X_{\max} uncertainties
 - ➔ from where ? results at LHC are very similar ...
- Remaining uncertainties in X_{\max} due to different results for pion-air interactions at high energy
 - ➔ Problem : no data for pion interactions at high energy
- X_{\max}^{μ} very sensitive to pion interactions at all energies (incl. high energies) so MPD can be use to probe pion interactions and limit uncertainties on X_{\max}
 - ➔ Role of baryons
 - ➔ pion spectra ?
- Test using EPOS LHC and Sibyll 2.3

Muon Number

- From Heitler

$$N_{\mu} = \left(\frac{E_0}{E_{dec}} \right)^{\alpha}, \quad \alpha = \frac{\ln N_{had}}{\ln (N_{had} + N_{em})}$$

→ In real shower, not only pions : Kaons, (anti)Baryons and resonances



$$\alpha = \frac{\ln (N_{had})}{\ln (N_{tot})} = 1 + \frac{\ln (R)}{\ln (N_{tot})}$$

$$R = \frac{N_{had}}{N_{tot}} \approx \frac{N_{\pi^{ch}} + N_B}{N_{\pi^{ch}} + N_B + N_{\pi^0}} < 1$$

Very important :
in (a)Baryon-Air interactions, no
leading neutral pion !
 $R \sim 1$

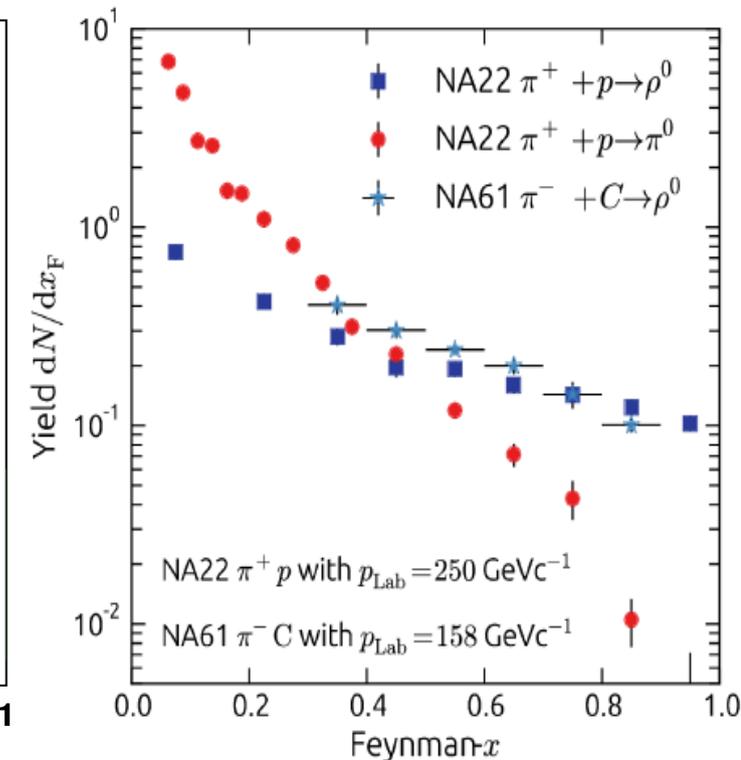
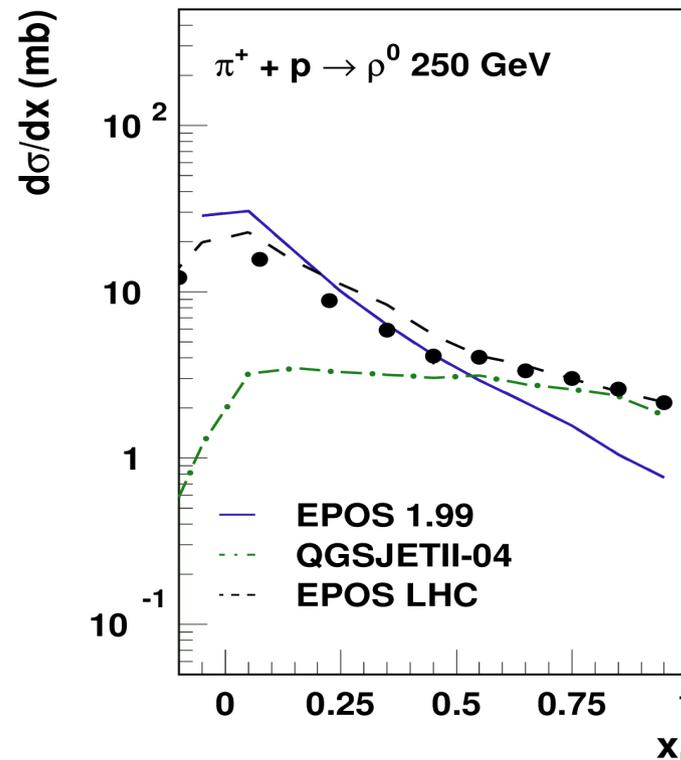
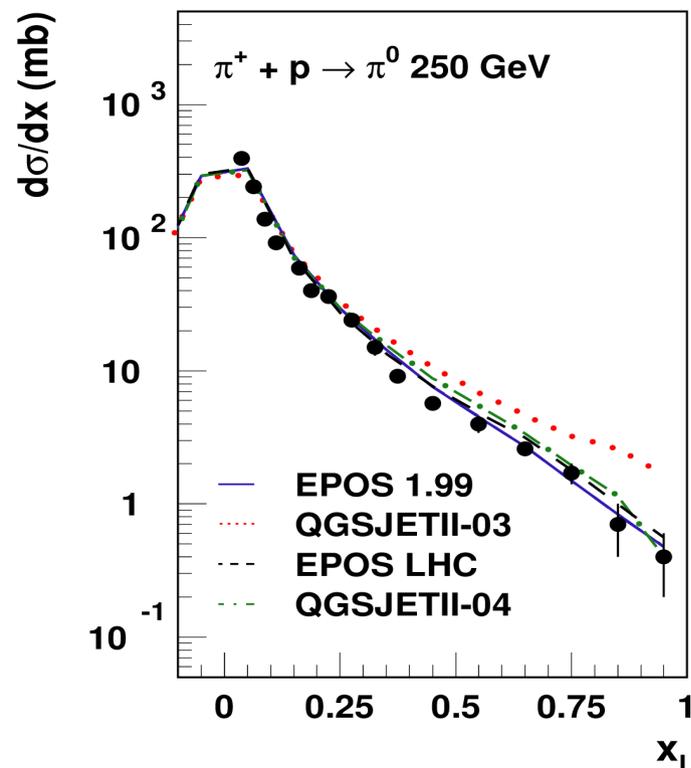
R depends on the number of (anti)B and ρ^0 in p- or π -Air interactions

More fast (anti)baryons or ρ^0 or larger $N_{tot} = \alpha \rightarrow 1$ = more muons

Pion Leading Particle Effect

Rho meson production added in QGSJETII (and Sibyll 2.3) to take into account leading particle effect in pion-Air interaction

- ➔ same effect as baryon production : forward π^0 replaced by charged pions (reduced leading π^0)
- ➔ increase muon production
- ➔ higher minimum muon energy (less generations) compared to baryons

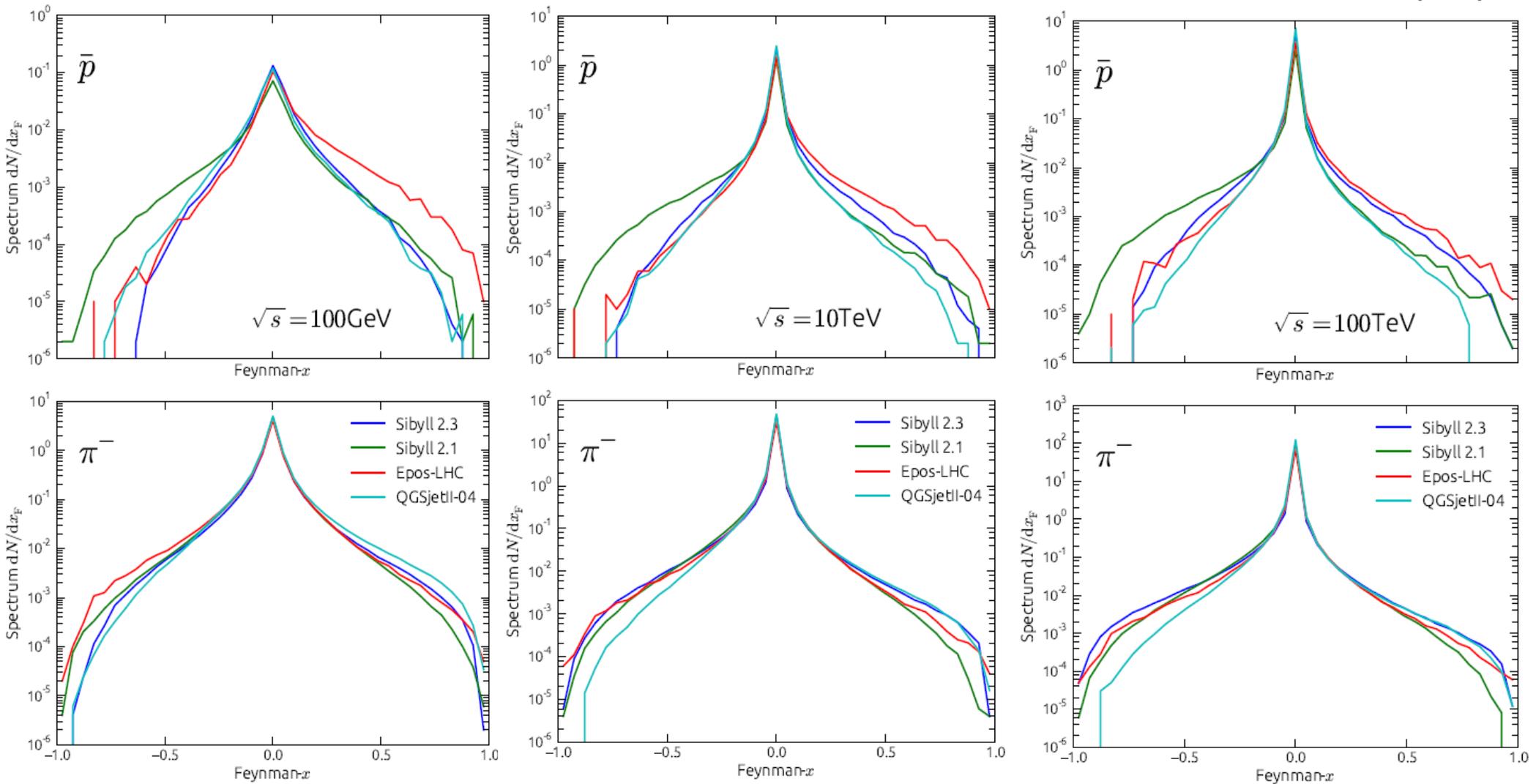


Plot from F. Riehn (KIT)

Pion(+)-Carbon Interactions

Different model predictions as a function of energy:

Plots from F. Riehn (KIT)

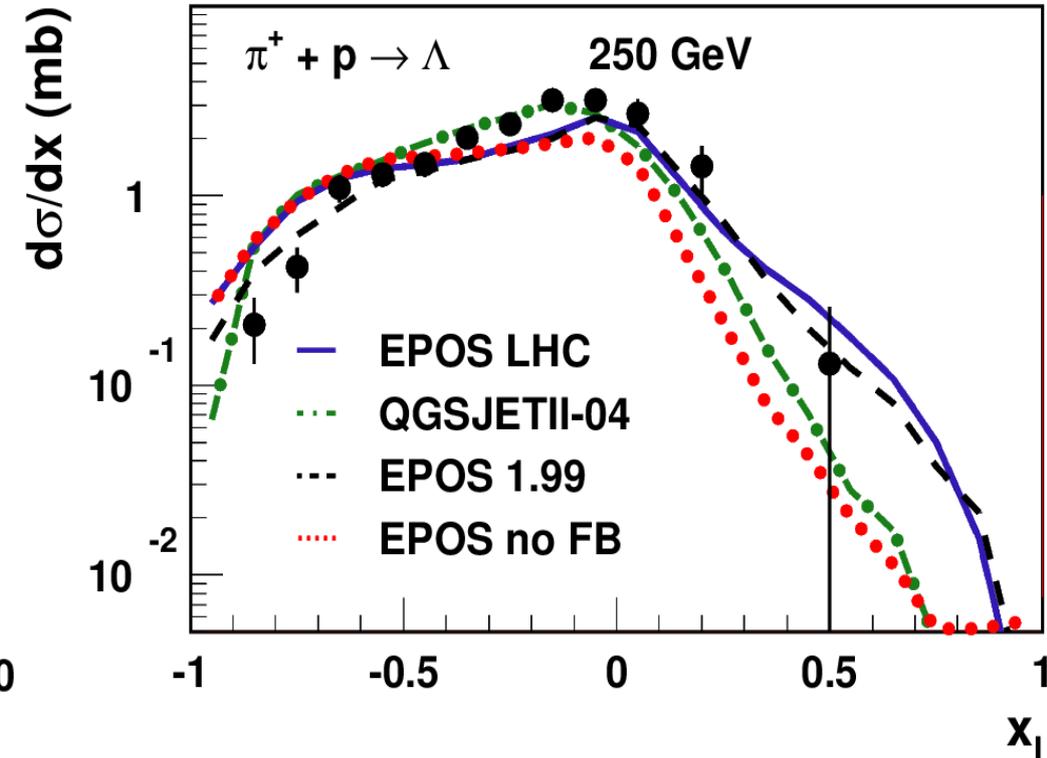
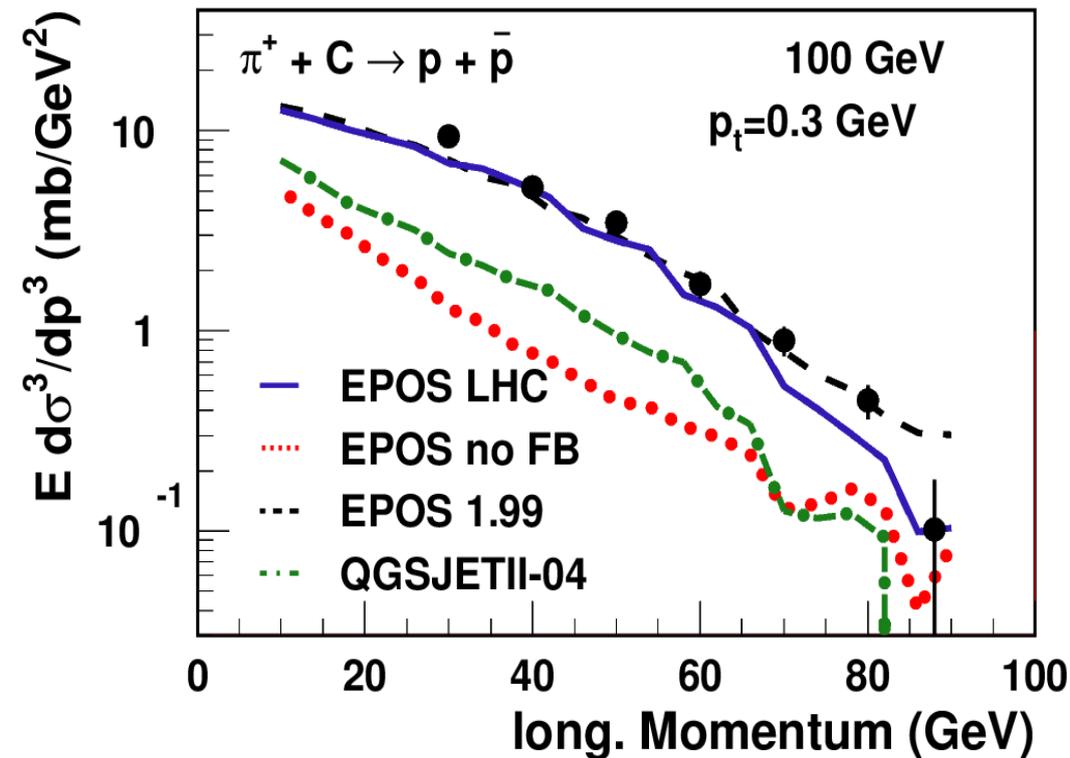


Baryons in Pion-Carbon

Very few data for baryon production from meson projectile, but for all :

- ➔ strong baryon acceleration (probability $\sim 20\%$ per string end)
- ➔ proton/antiproton asymmetry (valence quark effect)
- ➔ target mass dependence

➔ **NA61 Data to check !**

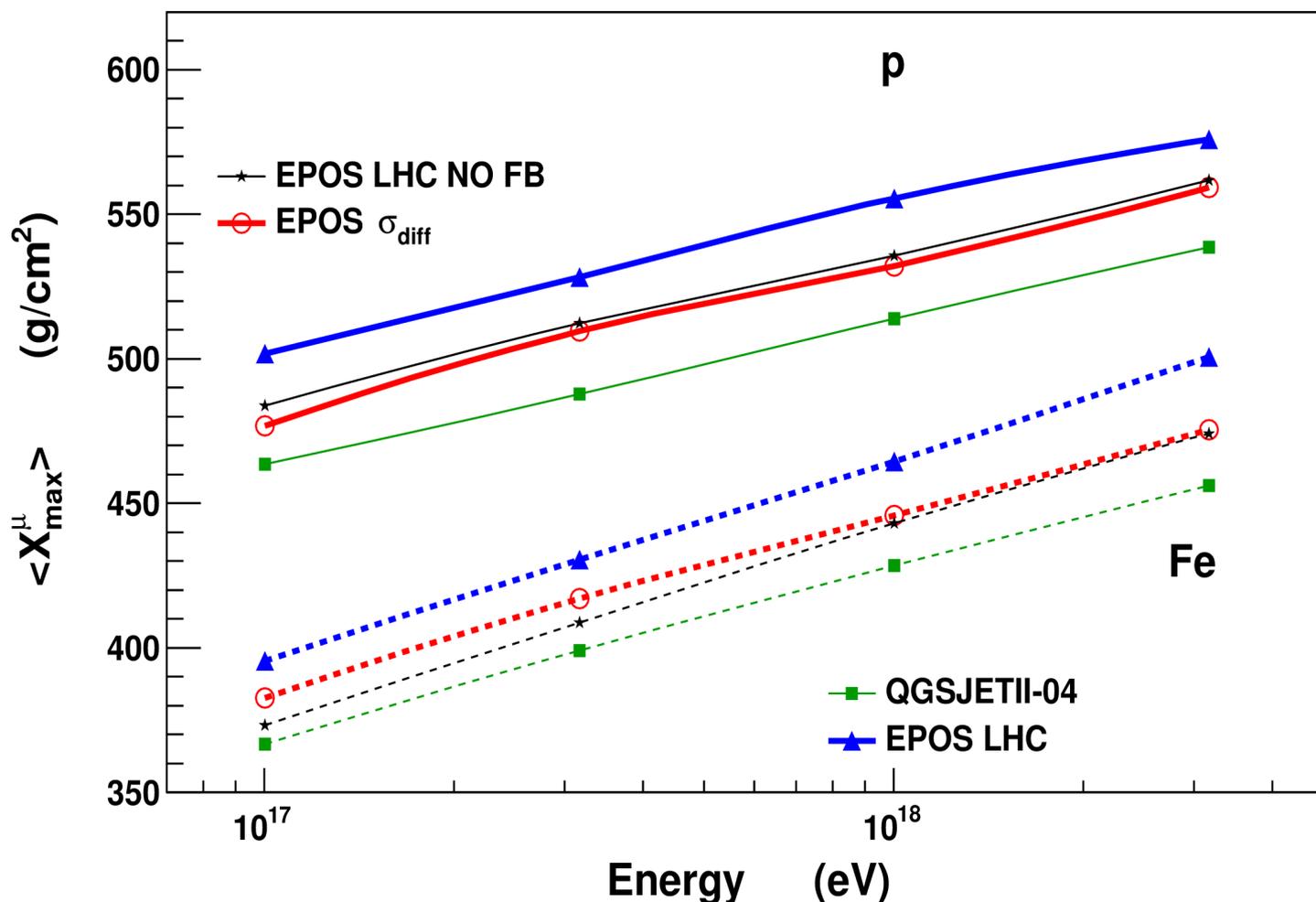


$\langle X_{\max}^{\mu} \rangle$ with modified EPOS LHC

Same than in mixed models

→ softer meson spectra (lower elasticity) : lower X_{\max}^{μ}

→ less forward baryons: lower X_{\max}^{μ}



-25 g/cm² for diff

-20 g/cm² for baryons

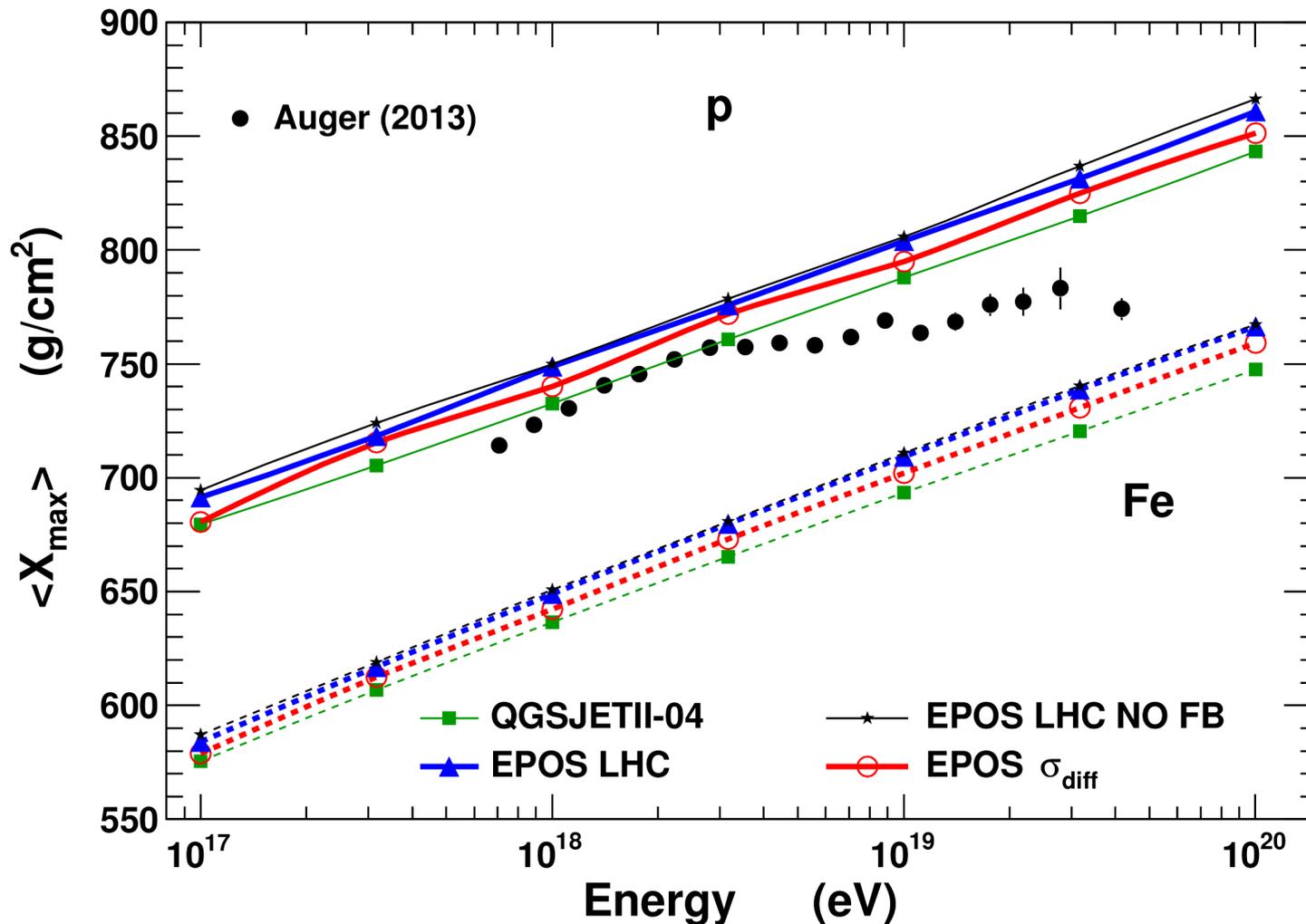
MPDs sensitive to baryon (less generation) and meson spectra in pion interactions

$\langle X_{\max} \rangle$ with Modified EPOS

Same than in mixed models

→ softer meson spectra: lower X_{\max}

→ forward baryons: small effect



-10 g/cm² for diff

~0 g/cm² for baryons

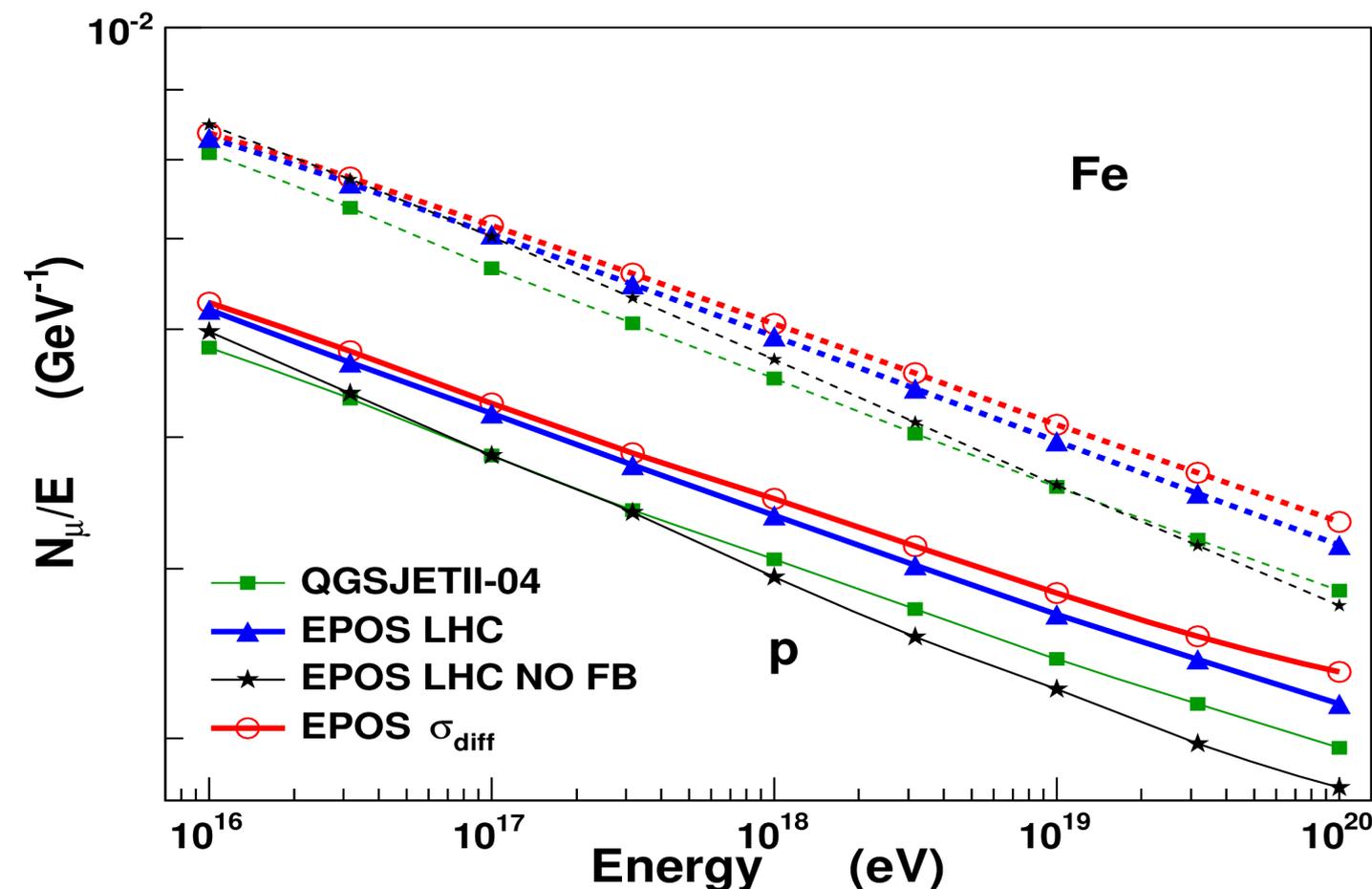
X_{\max} less sensitive to baryon spectra than to pion spectra in pion interactions

In Sergey's model, energy is not conserved (baryons not replaced by mesons)

N_μ with Modified EPOS

Number of muons depends on the same parameters

- softer meson spectra: larger N_μ
- forward baryons: lower N_μ but could be compensated by ρ^0 (keep energy to produce muons but doesn't change the number of generations: lower MPD)



N_μ sensitive to baryon (less generation) and meson spectra in pion interactions

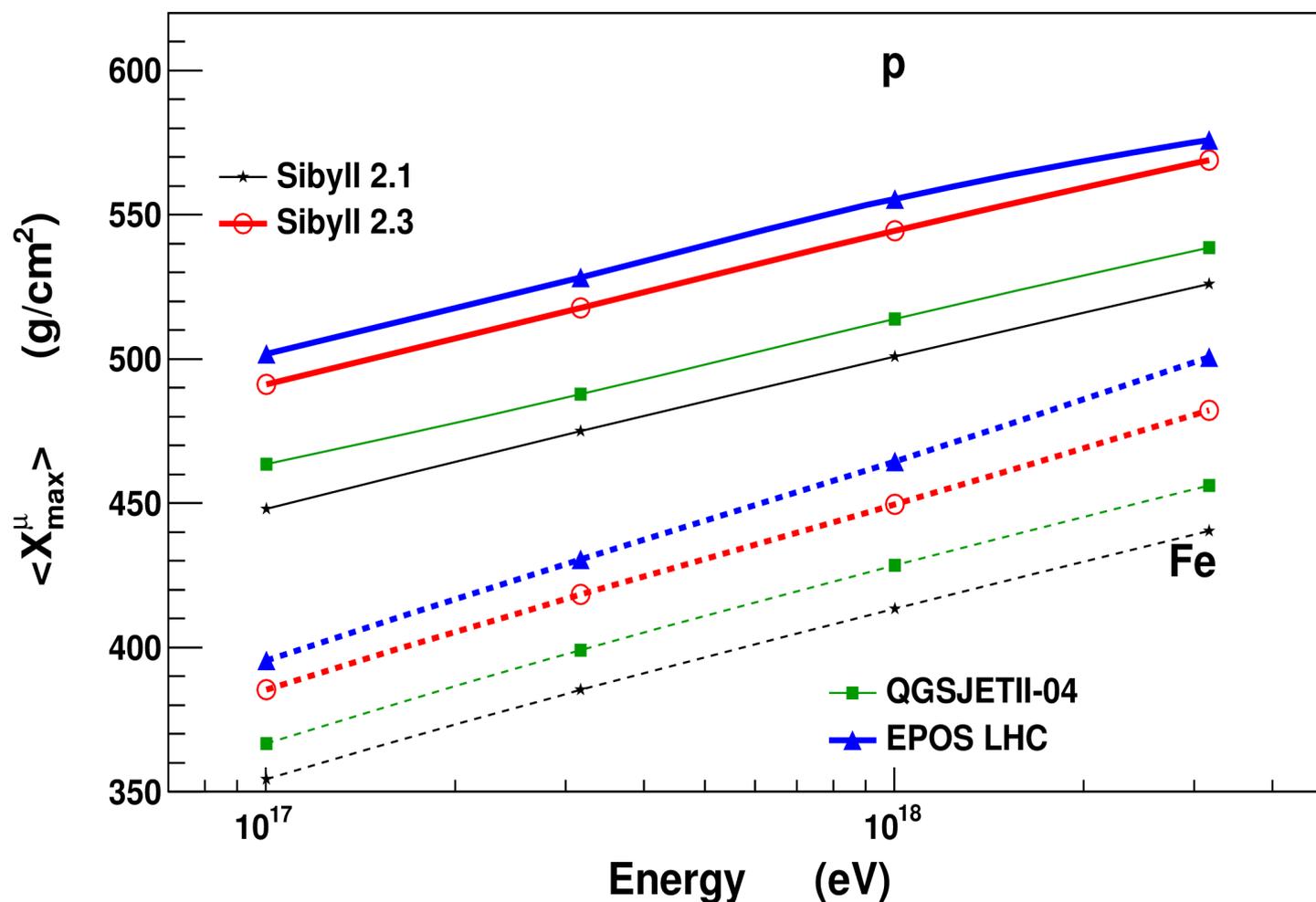
+5% for diff

-15% without forward baryons

$\langle X_{\max}^{\mu} \rangle$ with new Sibyll 2.3

● Same than for EPOS LHC

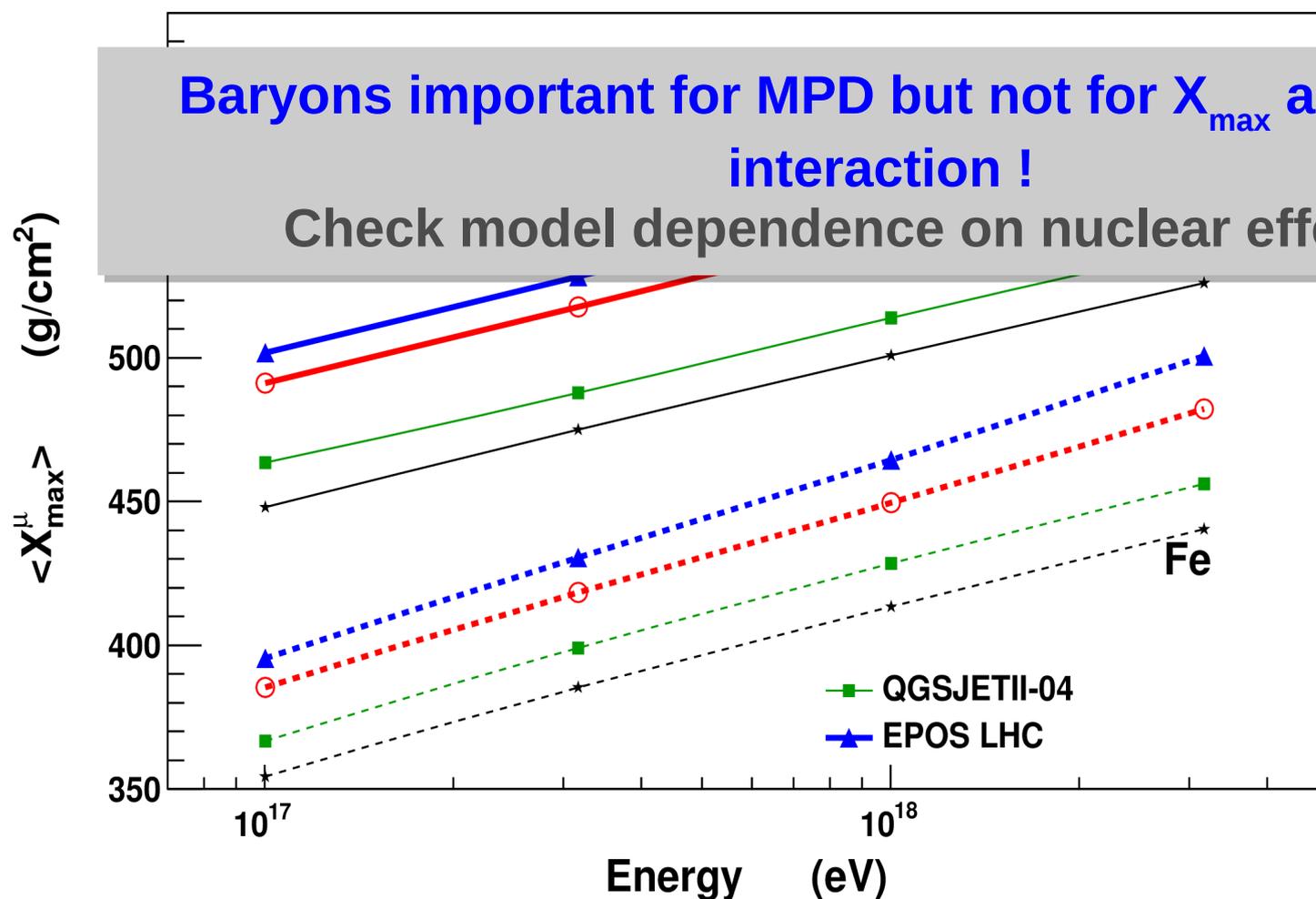
- ➔ low pion-air elasticity: higher X_{\max}^{μ}
- ➔ more forward baryons: higher X_{\max}^{μ}



MPDs sensitive to baryon (less generation) and meson spectra in pion interactions

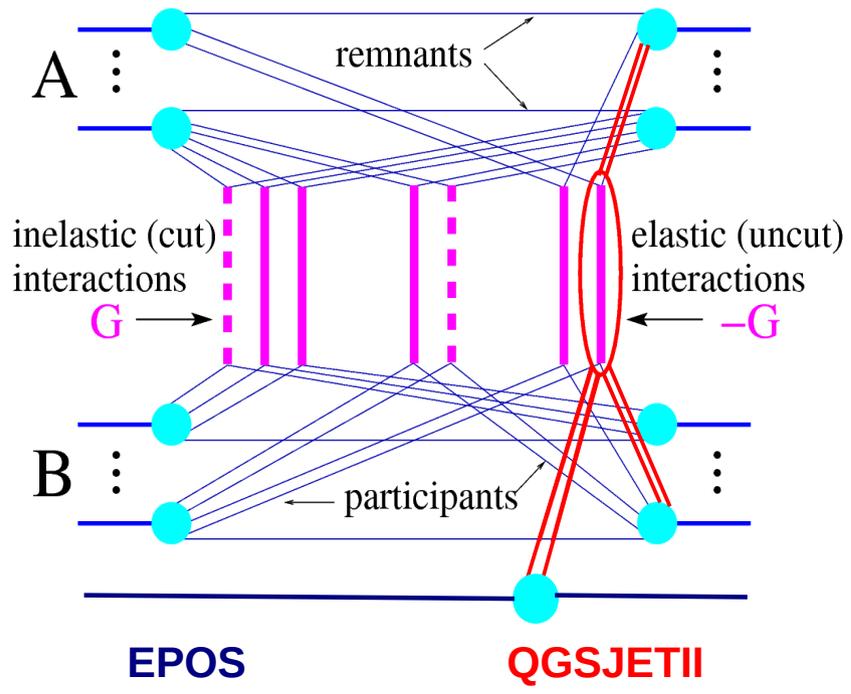
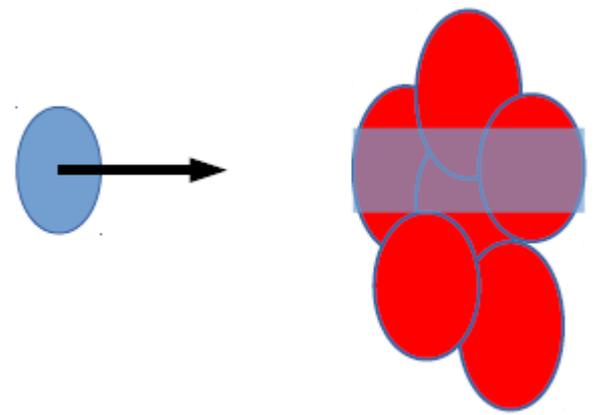
$\langle X_{\max}^{\mu} \rangle$ with new Sibyll 2.3

- Same than for EPOS LHC
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MPDs sensitive to baryon (less generation) and meson spectra in pion interactions

Nuclear Interactions

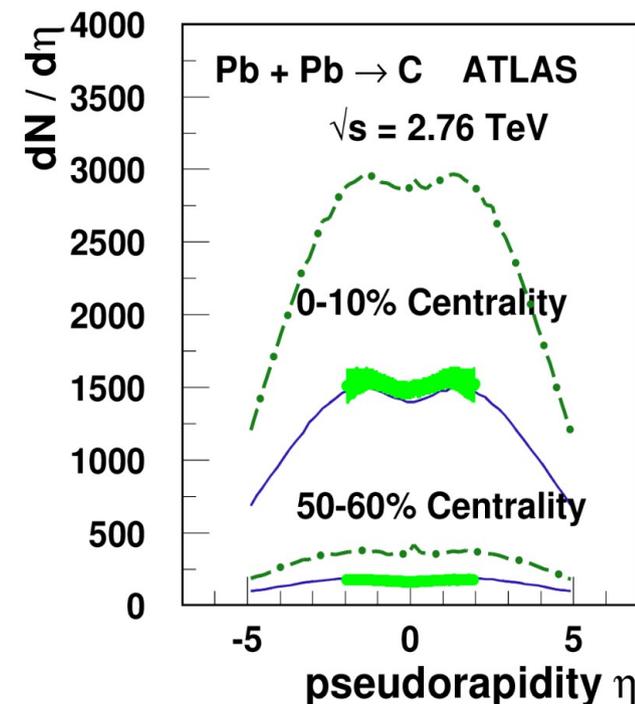
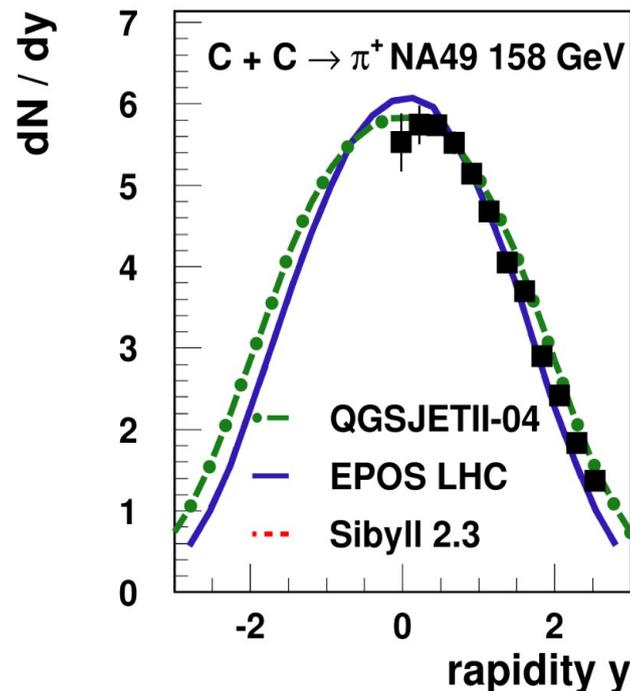
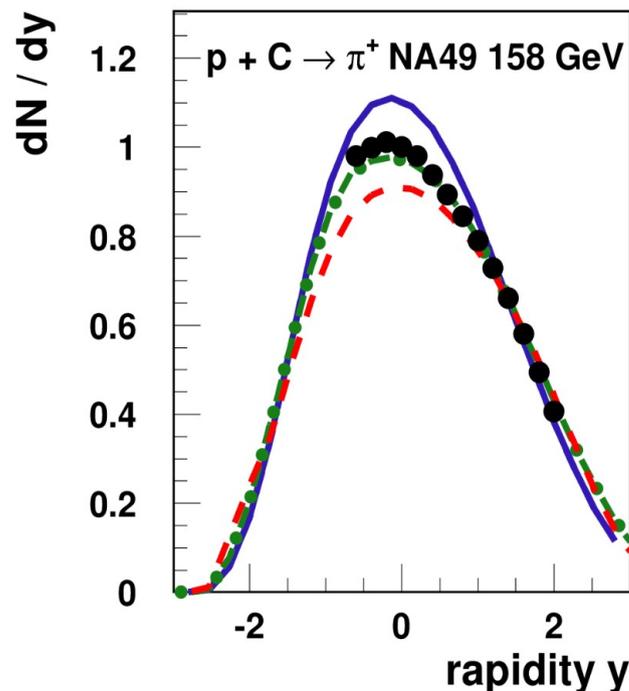


- **Sibyll**
 - ➔ Glauber for pA
 - with inelastic screening for diffraction in new Sibyll 2.3 (only nuclear effect)
 - ➔ superposition model for AA ($A \times pA$)
- **QGSJETII**
 - ➔ Pomeron configuration based on A projectiles and A targets
 - ➔ Nuclear effect due to multi-leg Pomerons
- **EPOS**
 - ➔ Pomeron configuration based on A projectiles and A targets
 - ➔ screening corrections depend on nuclei
 - ➔ final state interactions (core-corona approach and collective hadronization with flow for core)

Light Ion Data

Very few data to compare with all CR models :

- ➔ strong limitations in Sibyll (projectile up to Fe only and target up to O !)
- ➔ no final state interactions exclude heavy nuclei for QGSJETII
- ➔ no light ion at high energy

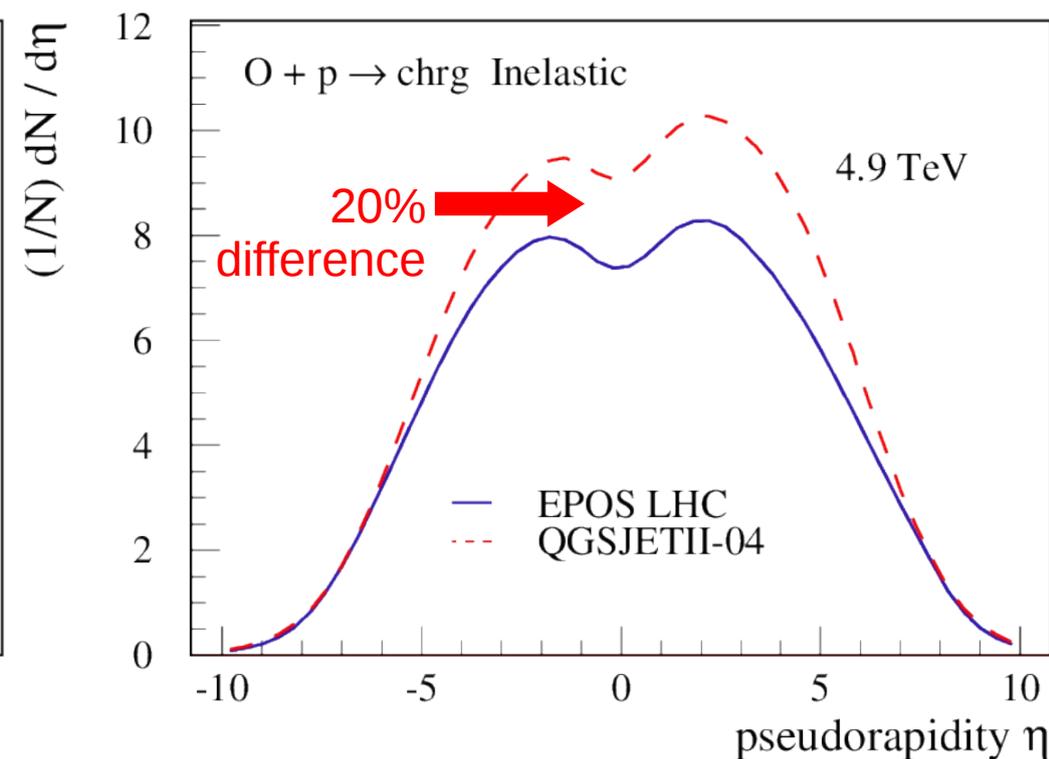
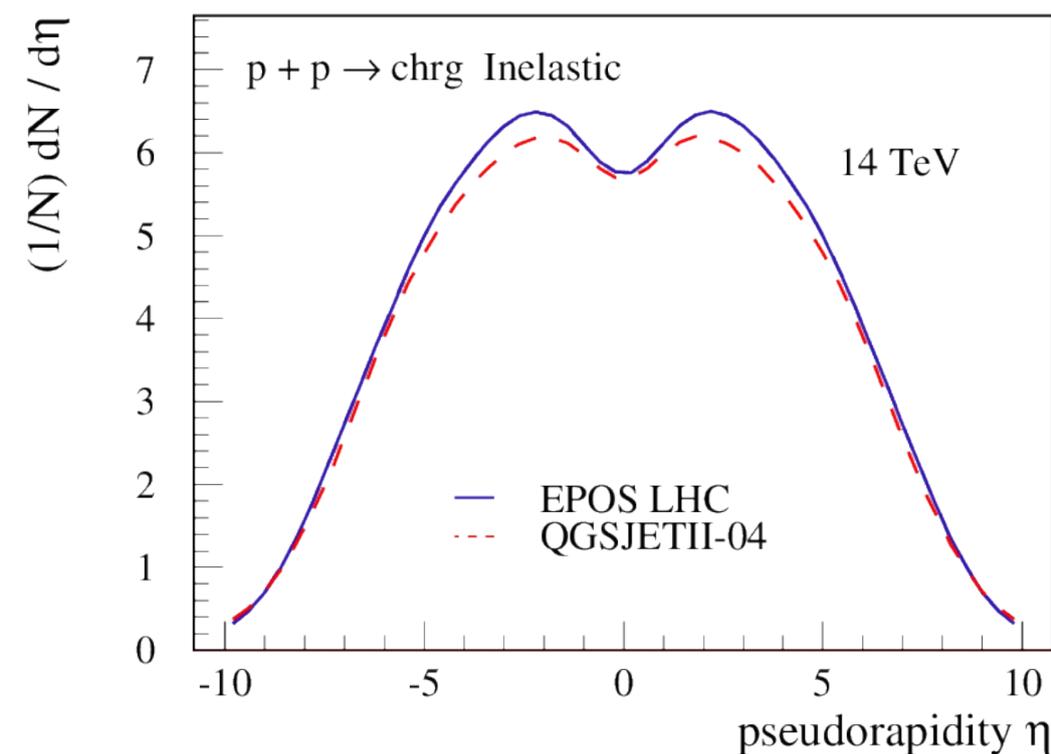


Light Ion Data

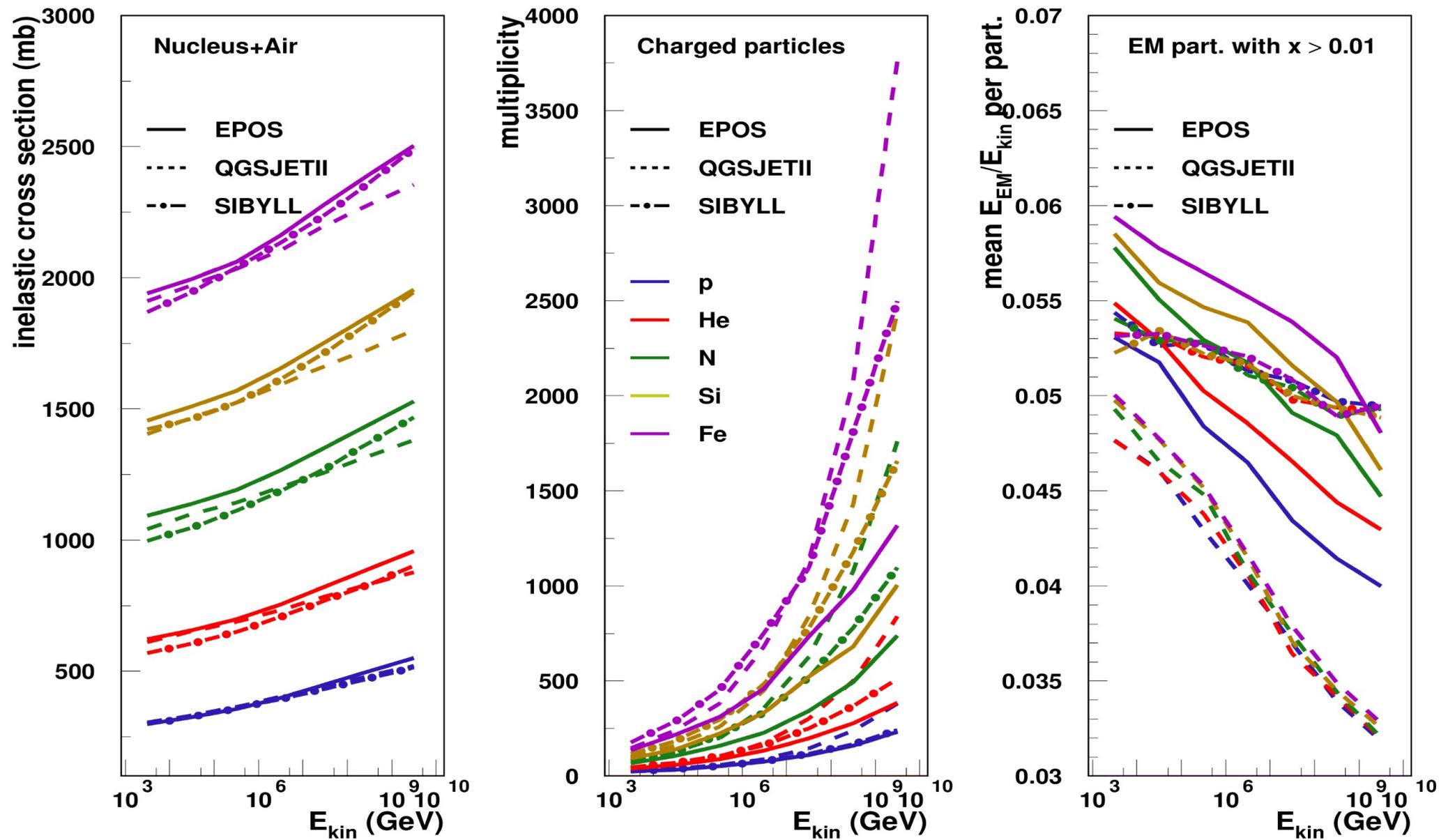
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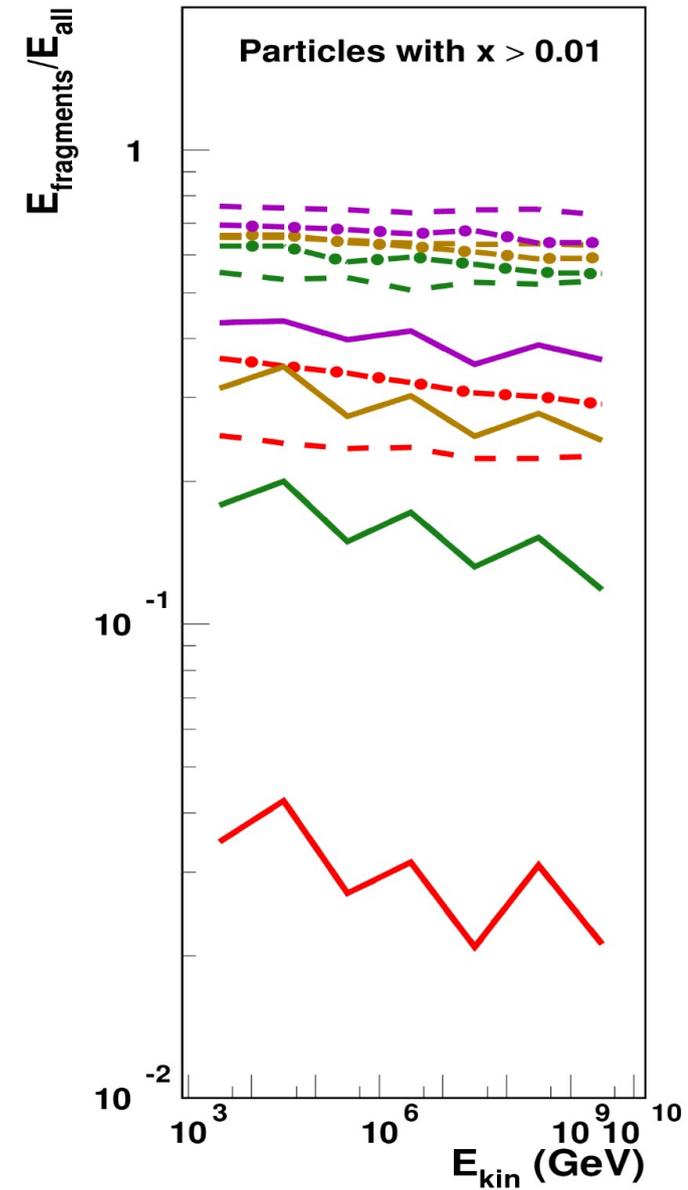
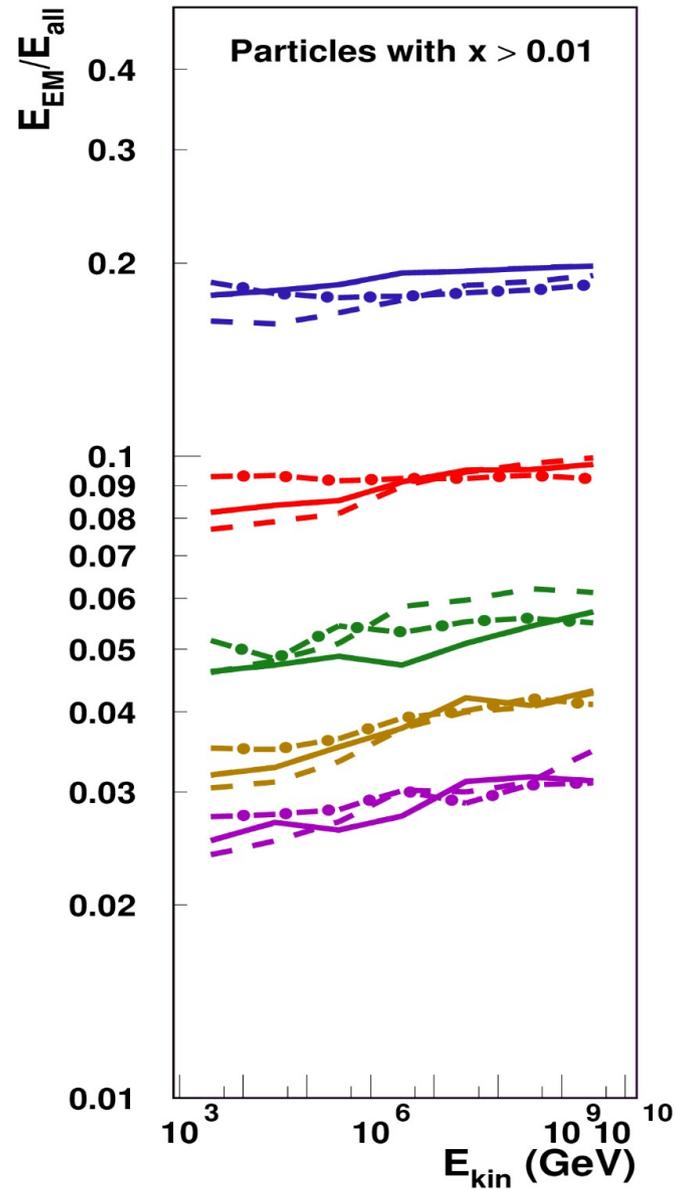
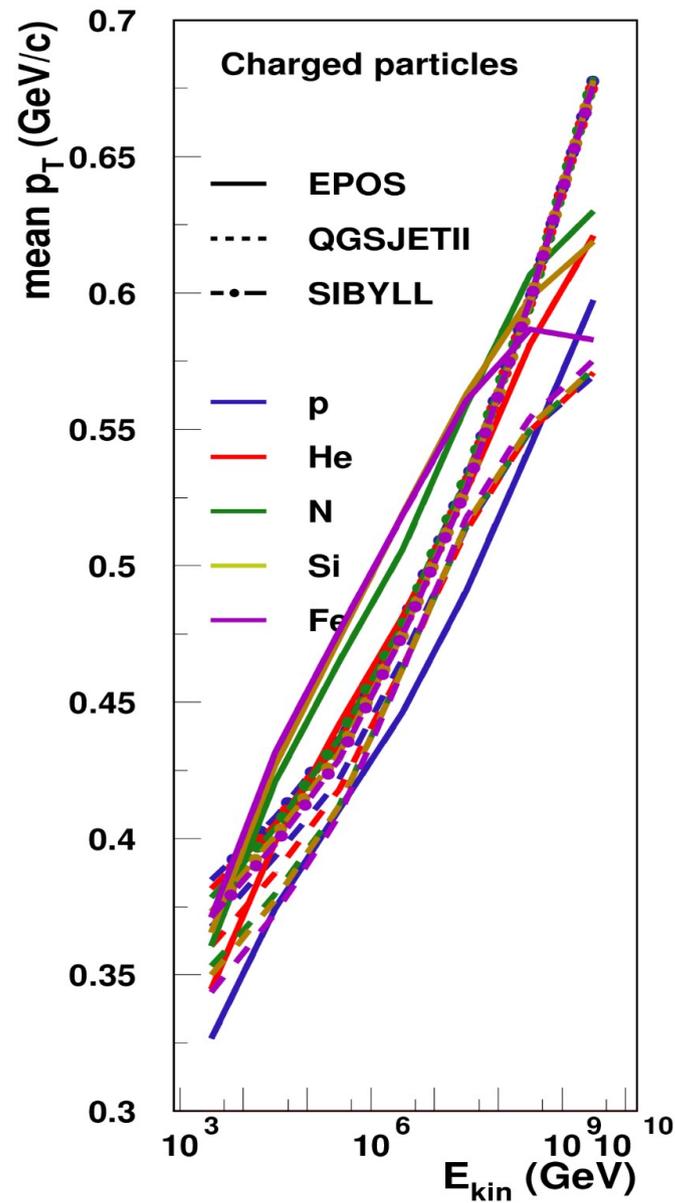
➔ pO@LHC to check models at high energy



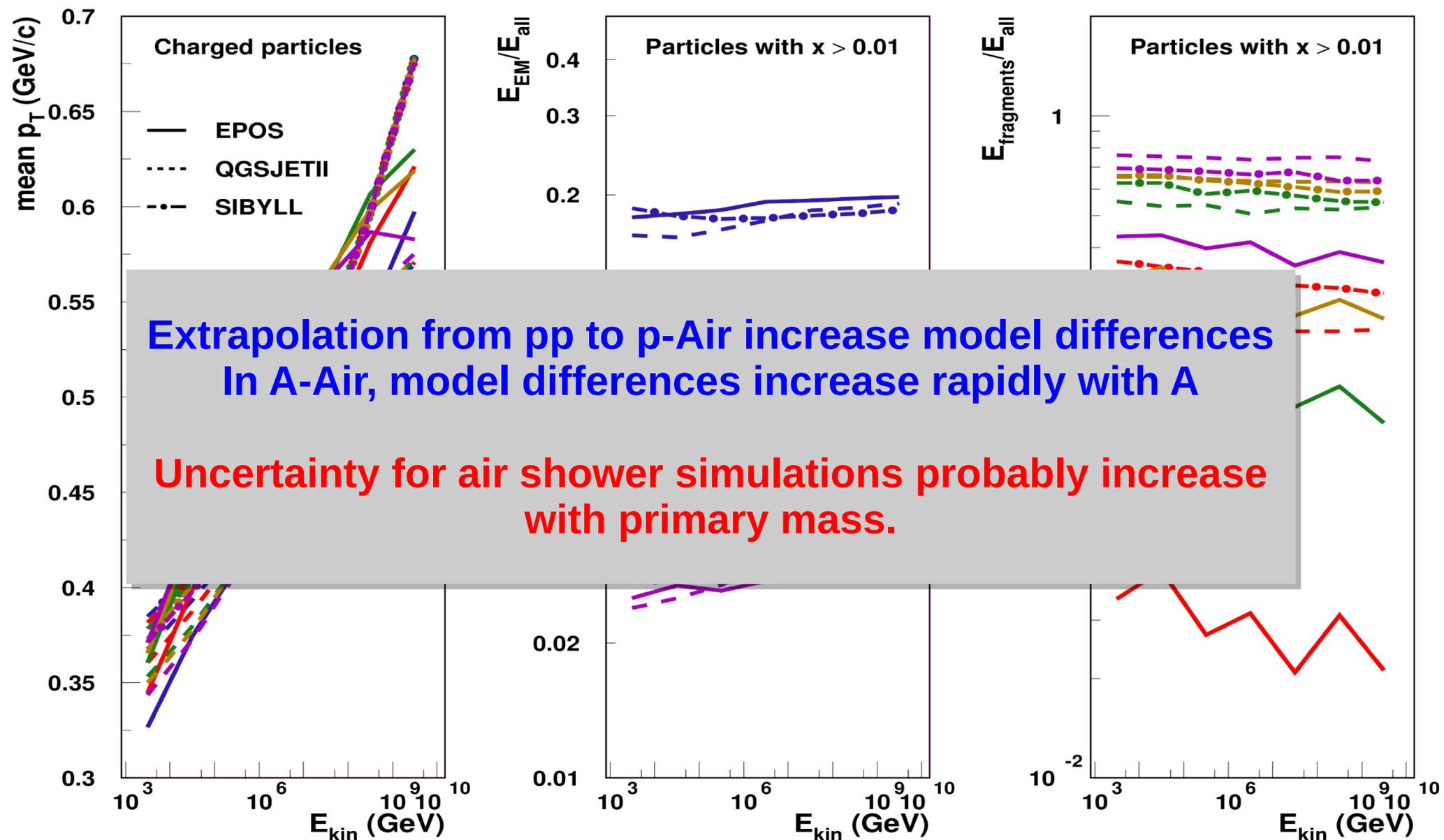
Model Comparison (1)



Model Comparison (2)



Model Comparison (2)



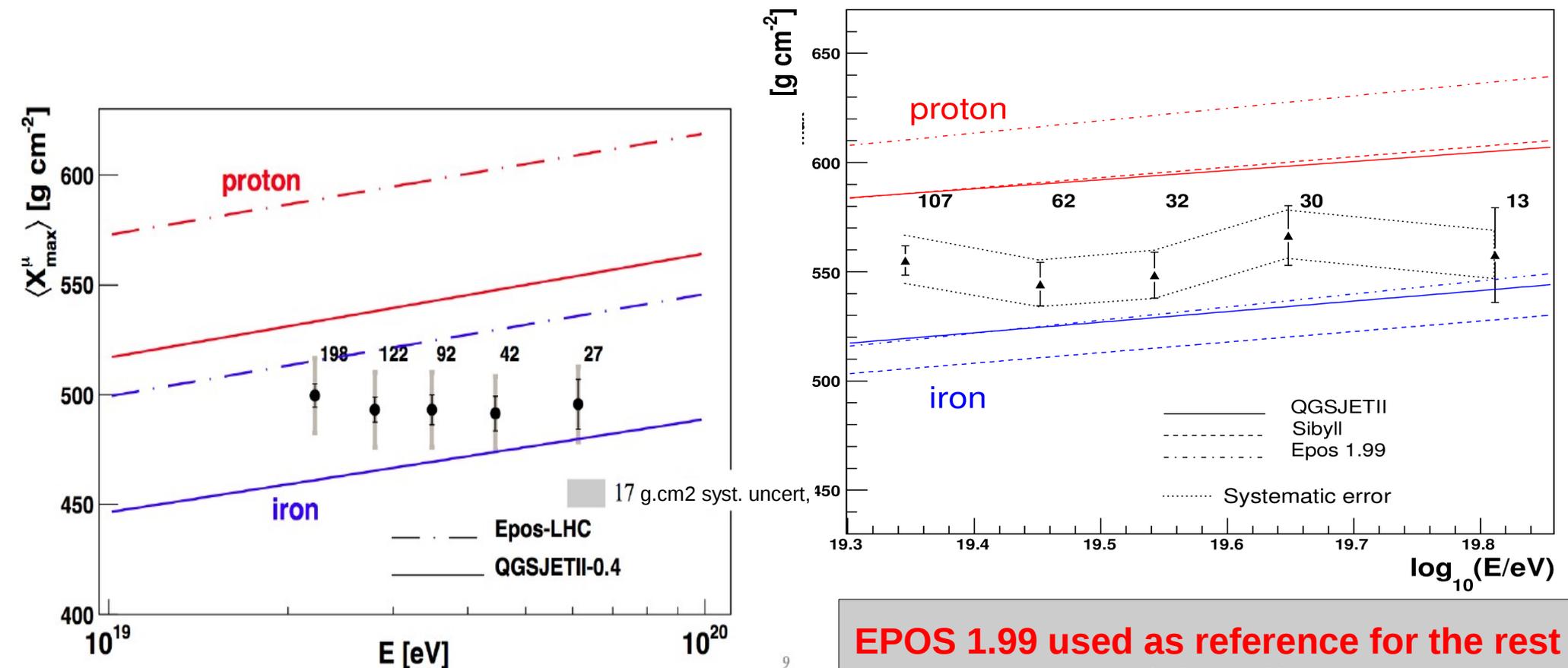
Summary

- **Very strong sensitivity of MPD on pion interactions which is badly measured**
 - ◆ MPD can be used to constrain models
 - ◆ then MPD can not be used for mass composition (X_{\max} less sensitive to details) unless more accelerator data can constrain the models
- **Better MPD = better X_{\max} ?**
 - ➔ YES
 - ◆ meson spectra influence both MPD and X_{\max}
 - ➔ NO
 - ◆ forward baryons change MPD but X_{\max} only if meson spectra is not changed
 - ◆ in EPOS LHC if forward baryons are suppressed, we get harder meson spectra and X_{\max} do not change
- **Remaining main source of uncertainty in X_{\max} probably related to extrapolations due to nuclear interactions** (lack of data at high energy and forward). **See David D'enterria talk for more hint on that.**

MPD and EPOS

● 2 independent mass composition measurements

- ➔ both results should be between p and Fe
- ➔ both results should give the same mean logarithmic mass for the same model
- ➔ problem with EPOS appears after corrections motivated by LHC data

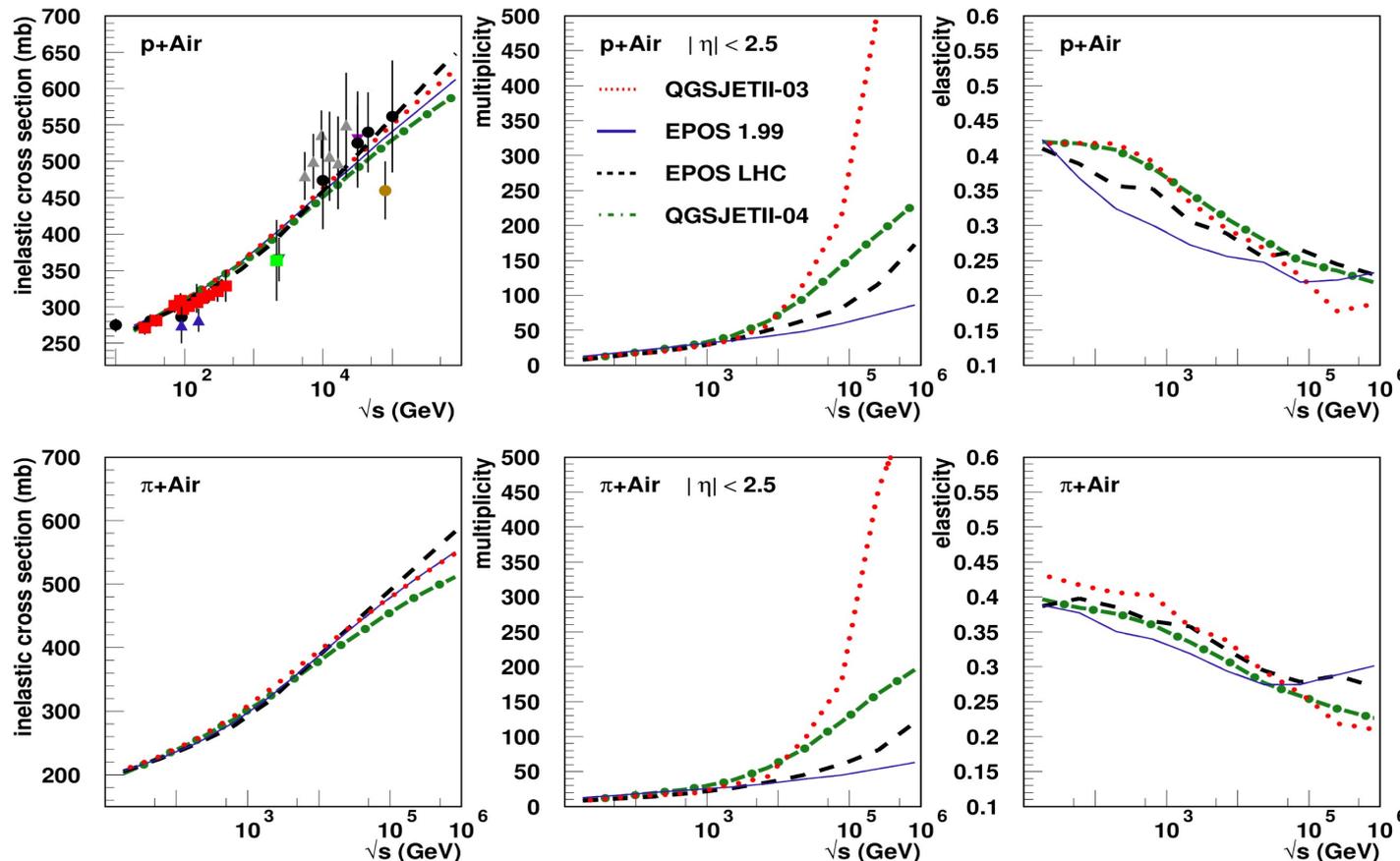


EPOS 1.99 used as reference for the rest of the talk

Difference EPOS 1.99/EPOS LHC

● EPOS 1.99 to EPOS LHC

- ➔ tune cross section to TOTEM value
- ➔ change old flow calculation to a more realistic one
- ➔ introduce central diffraction and improve rapidity gap distributions



- shallower MPD
- ➔ larger multiplicity
- ➔ larger cross-section (small)
- deeper MPD
- ➔ larger elasticity

Elasticity should be the source of the difference

(In)elasticity

● Difficult to measure : larger uncertainty

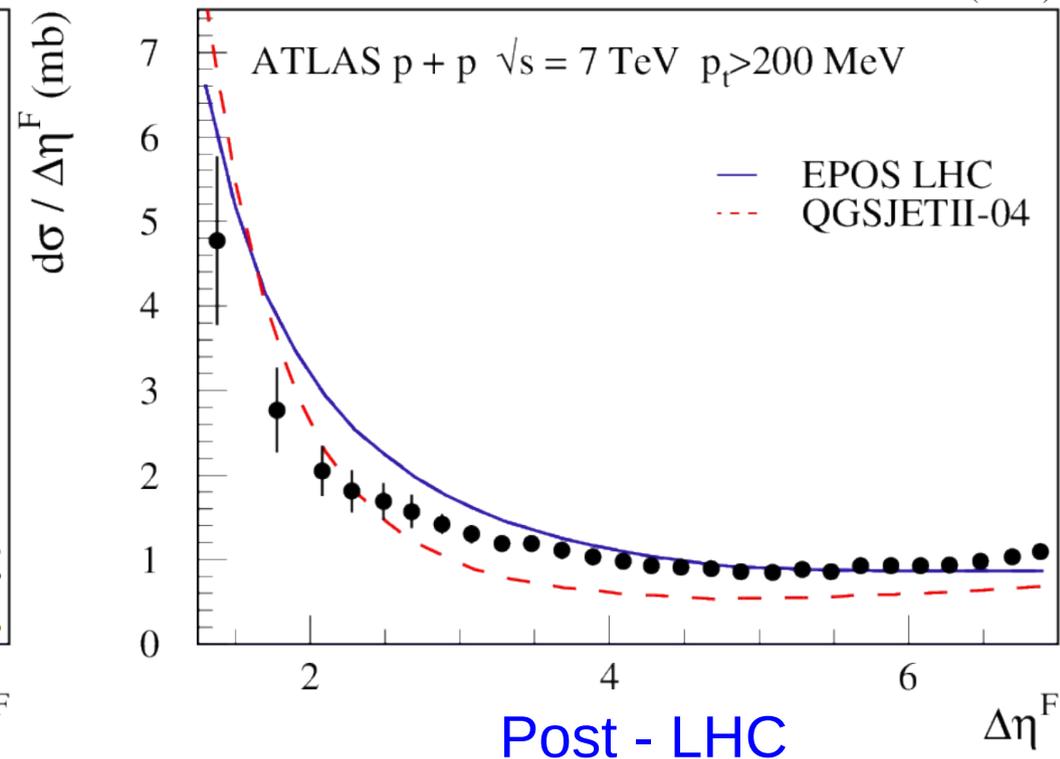
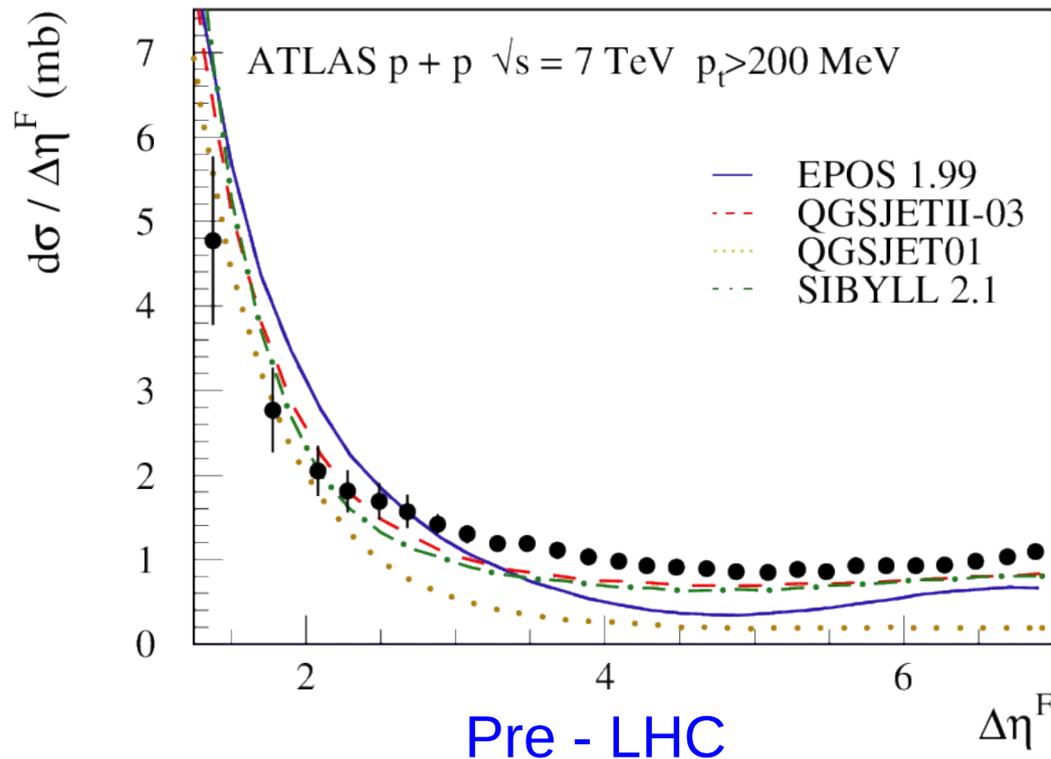
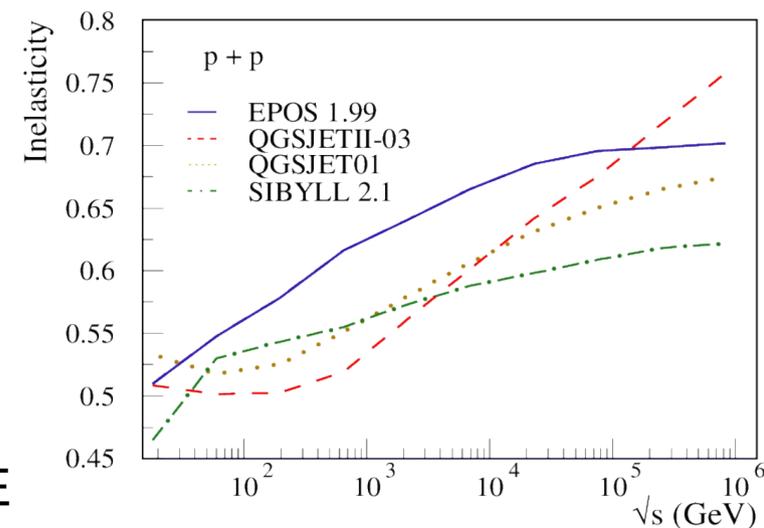
➔ Difference in diffraction

■ low mass / high mass / central diffraction

➔ difference for pions/Kaons/nucleons

■ very few data (and at low energy)

➔ Rapidity gap : first precise measurement at high E

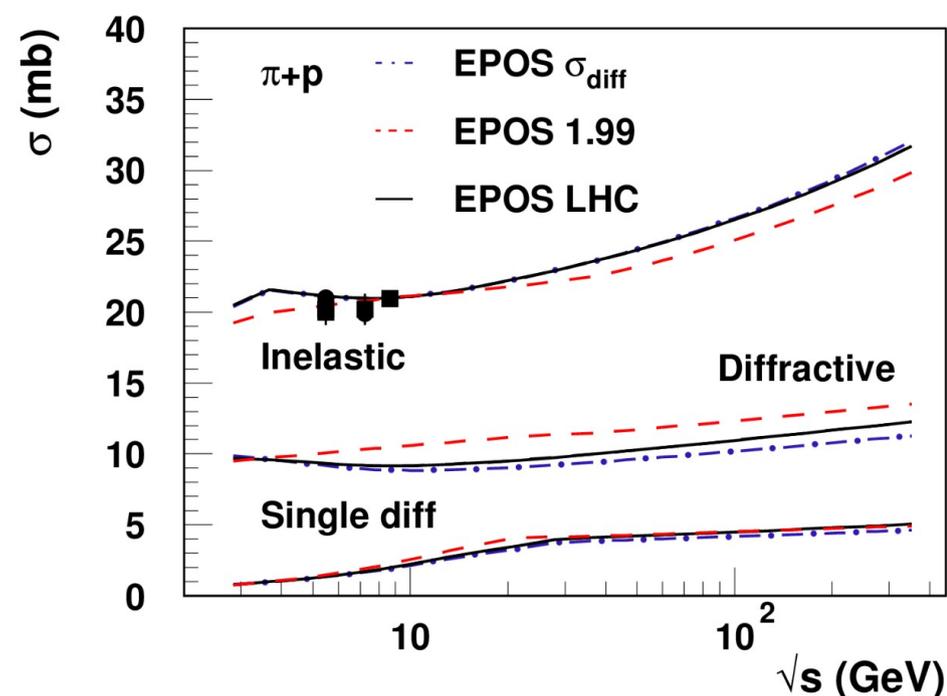


Pion Diffraction and MPD

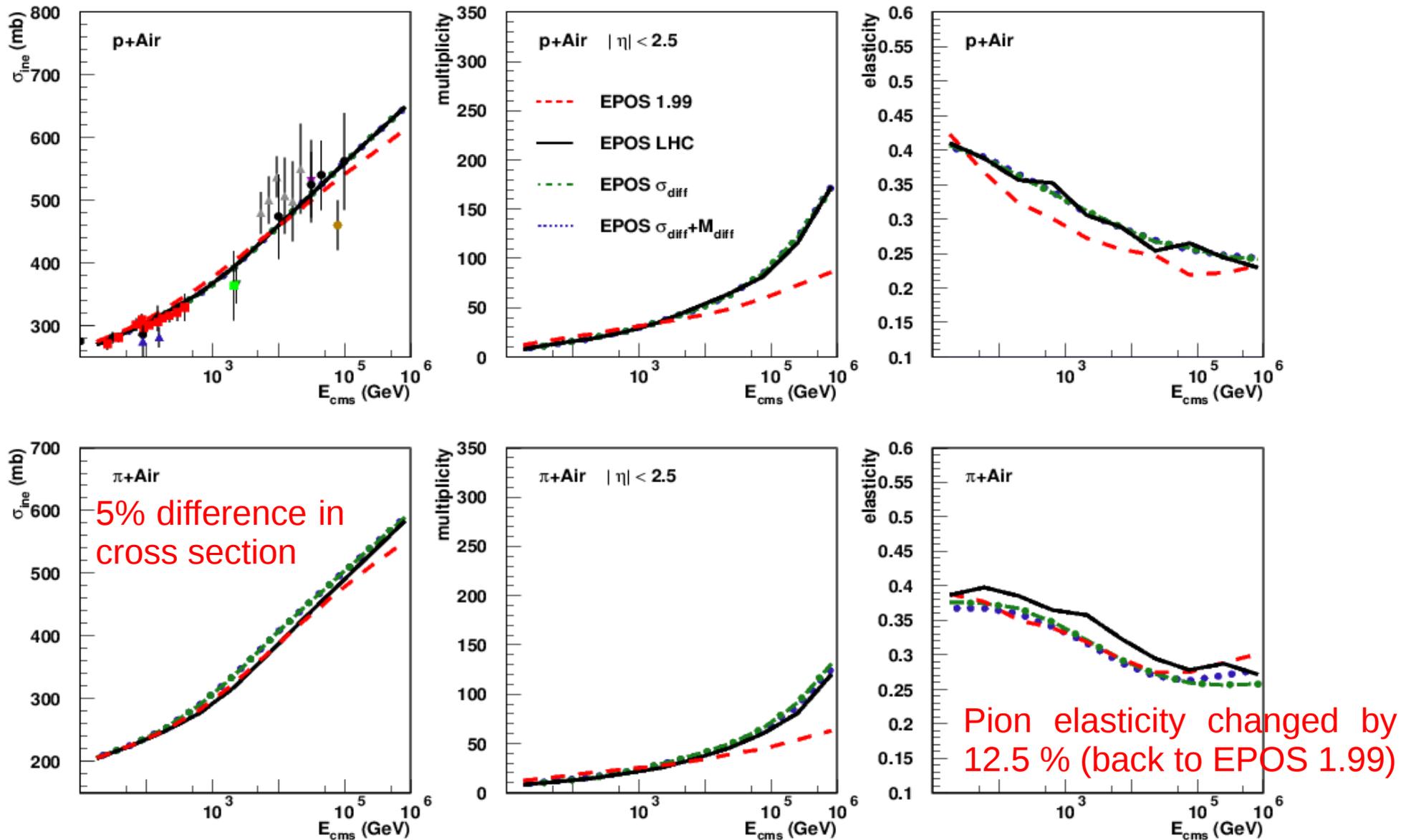
- **Rapidity gap measurement fixed by LHC**
 - ➔ should not change proton interactions
- **MPD driven by long chain of pion-Air interaction**
 - ➔ Modify in EPOS pion diffraction only
 - ➔ Test cross-section and diffractive mass distribution
 - ➔ first check existing pion data to tune parameter to REDUCE pion diffraction and INCREASE diffractive mass

- **2 “tunes”**

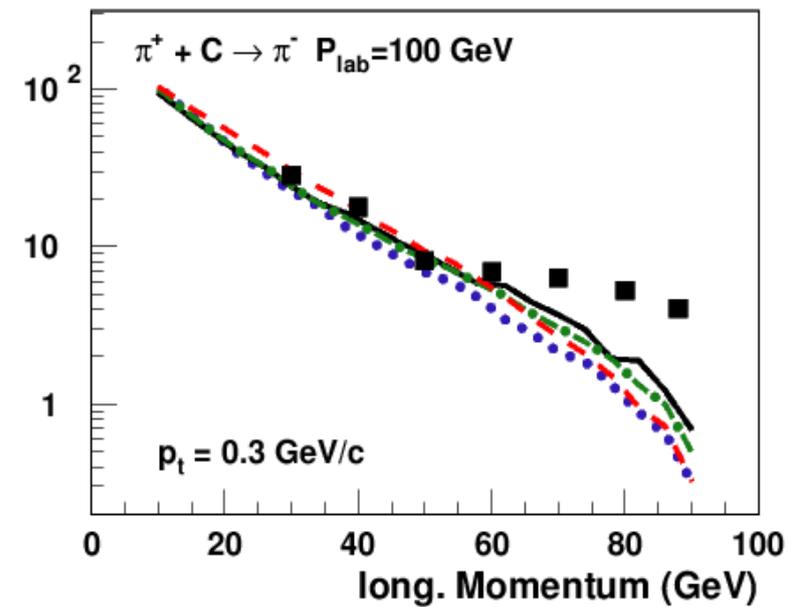
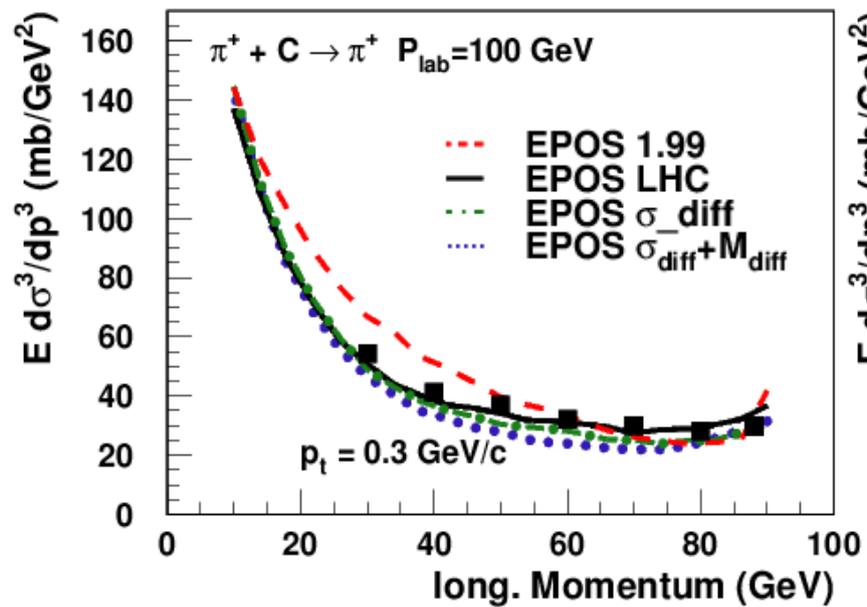
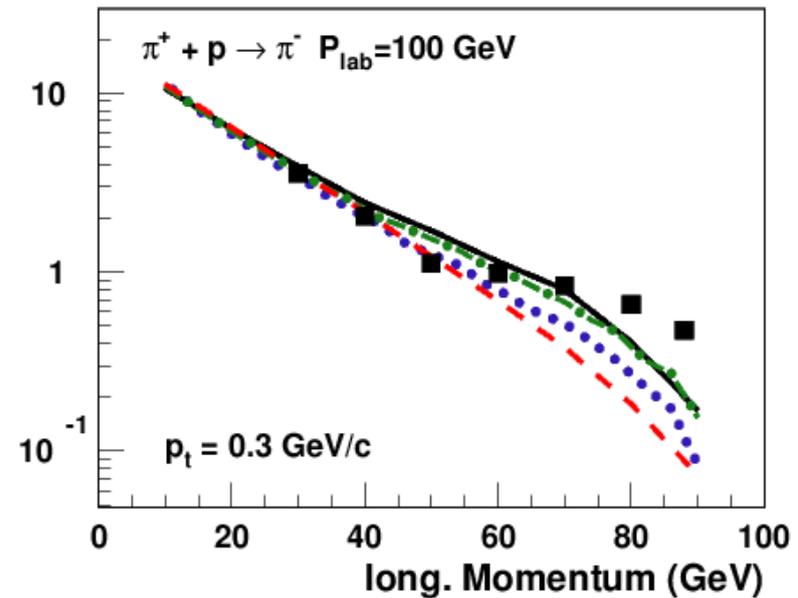
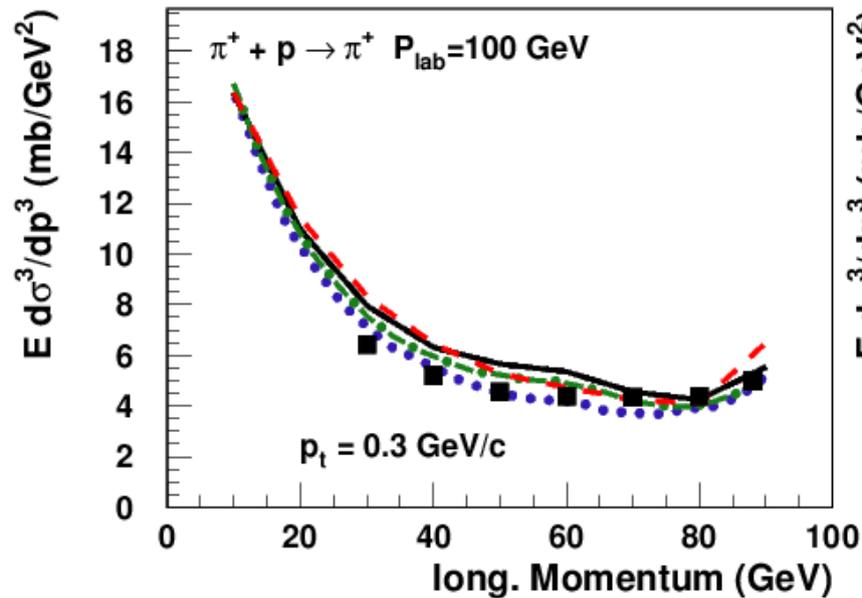
- ➔ EPOS (LHC) σ_{diff} :
diffractive cross section reduced
- ➔ EPOS (LHC) $\sigma_{\text{diff}} + M_{\text{diff}}$:
diffractive cross-section reduced and mass increased



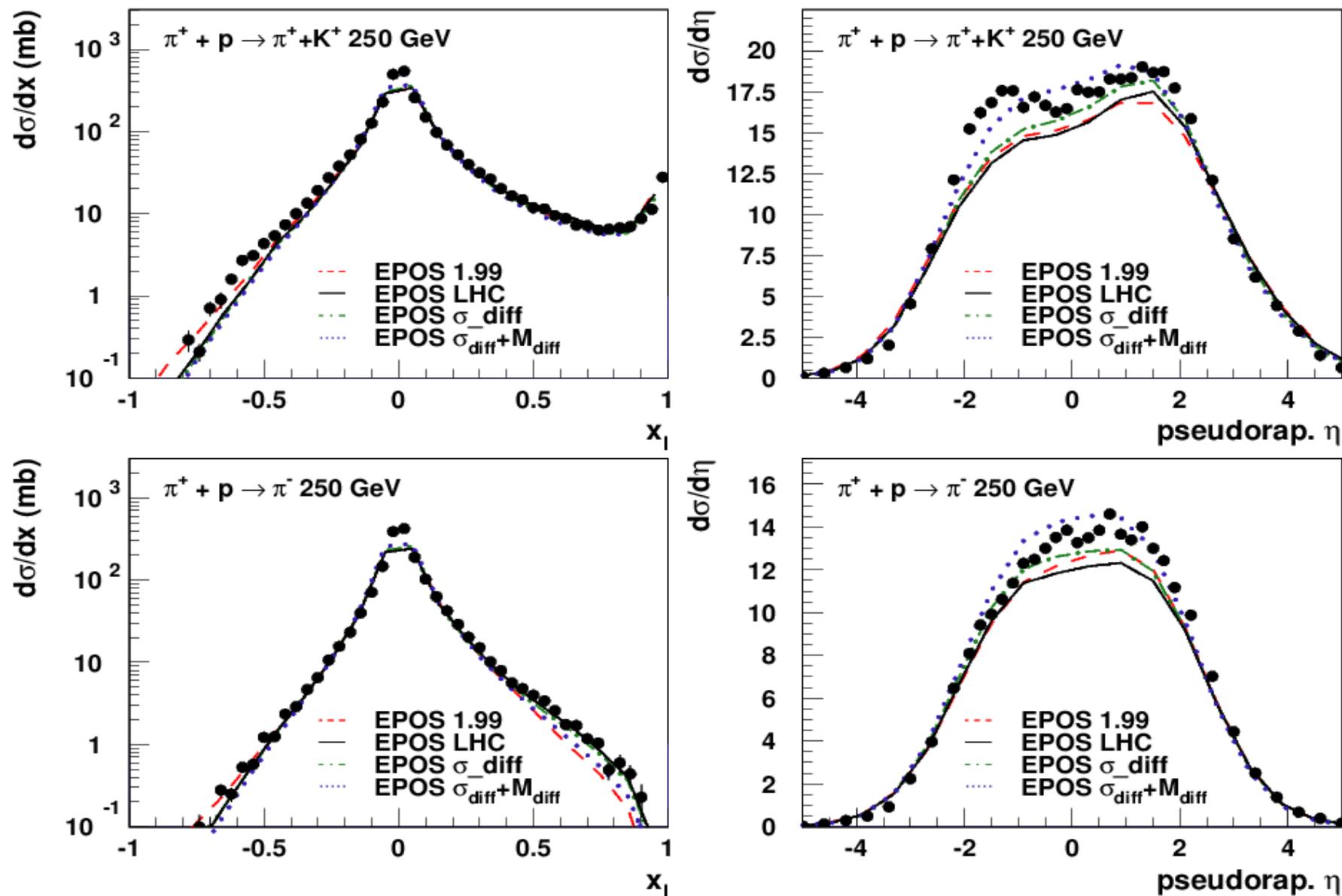
Extrapolation to CR interactions



Test with accelerator data

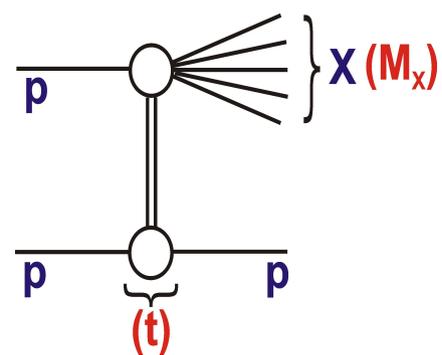


Test with accelerator data

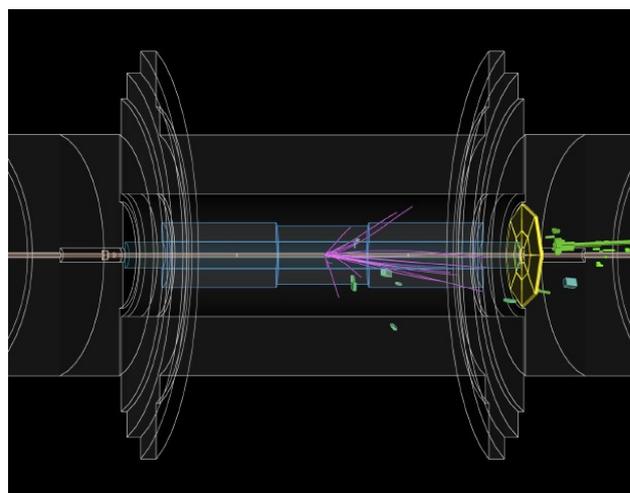


Rapidity Gap and (In)elasticity

diffractive process



ATLAS detector



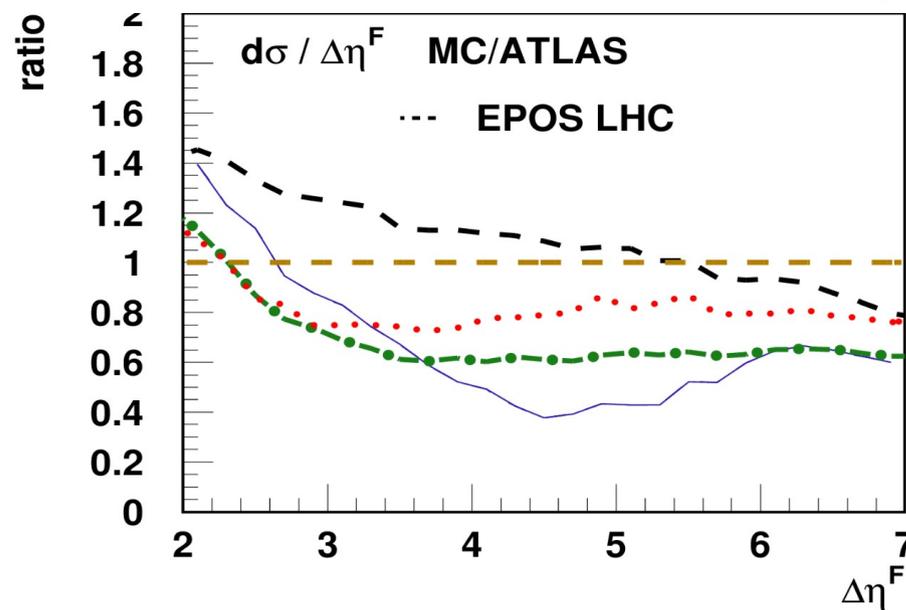
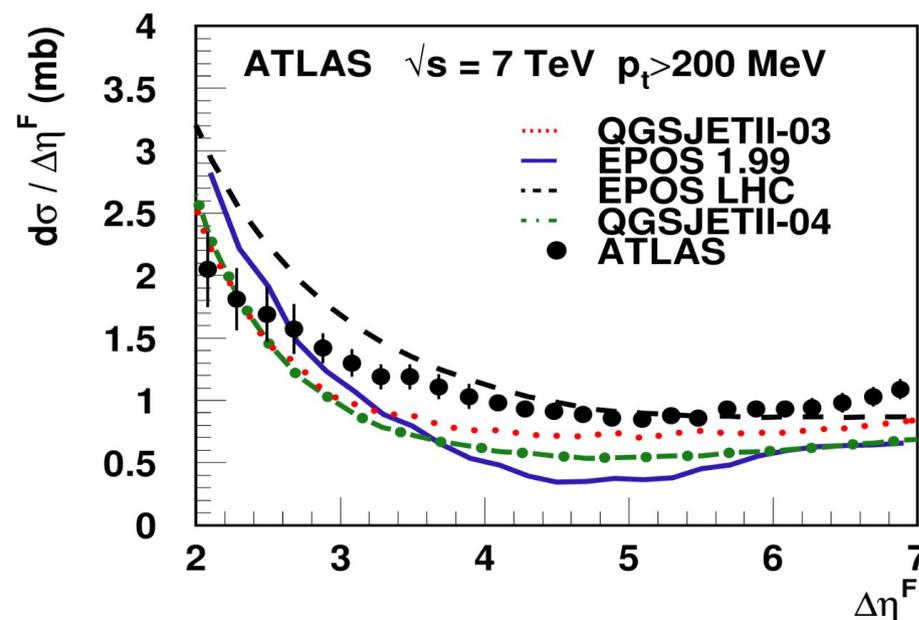
ATLAS Collaboration

- **Rapidity gap closely related to diffraction**

- ➔ diffractive cross-section
- ➔ AND diffractive mass distribution

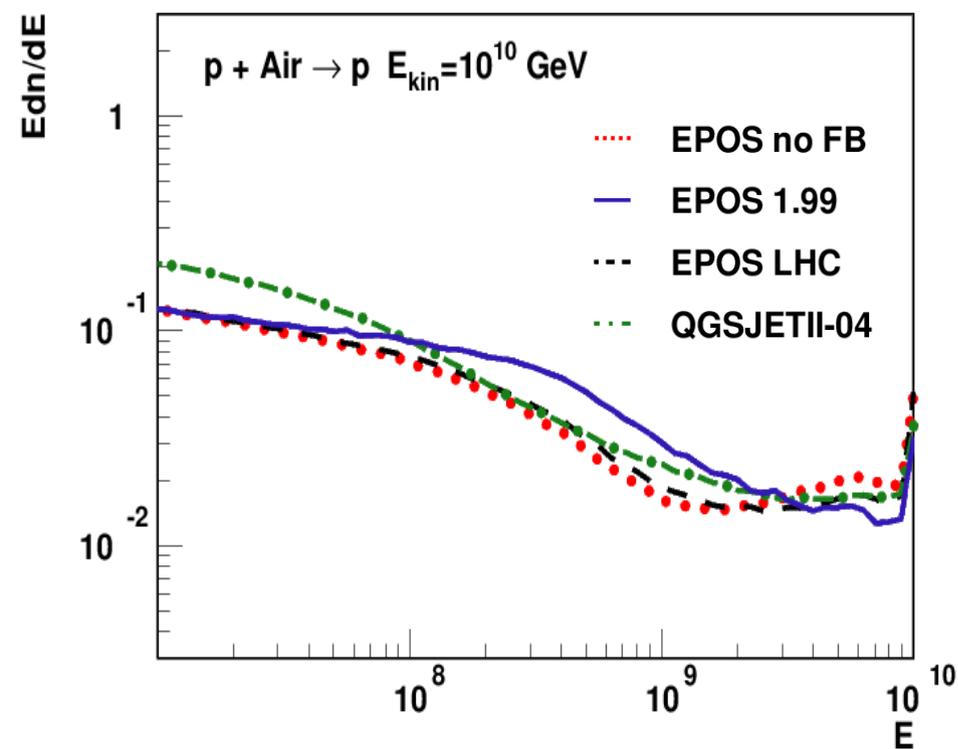
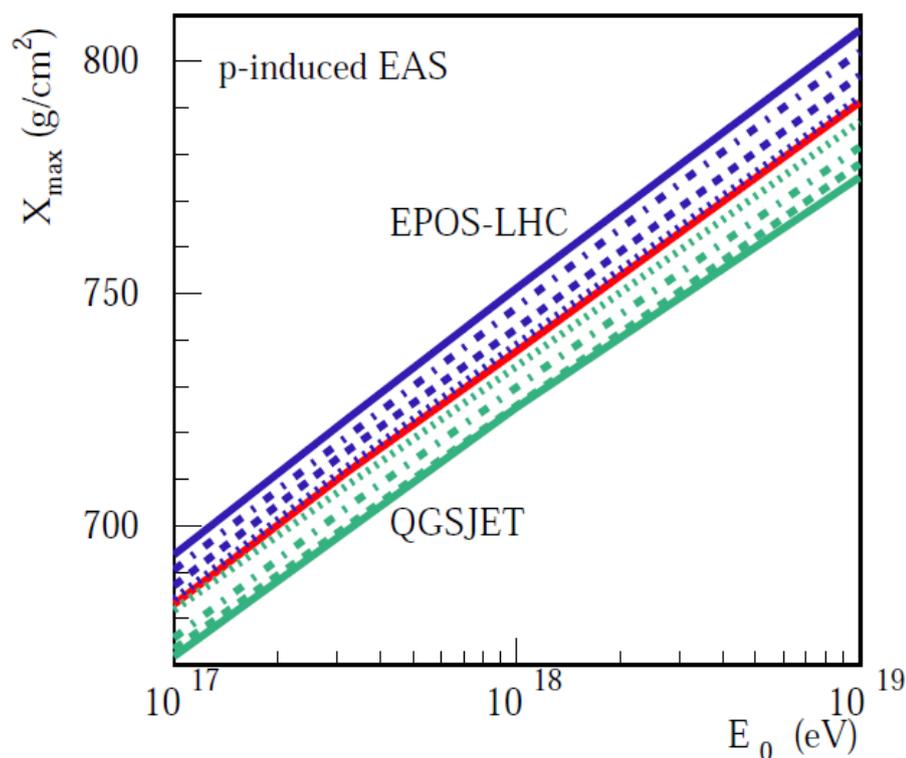
- **Hard constraint for CR**

- ➔ change (in)elasticity



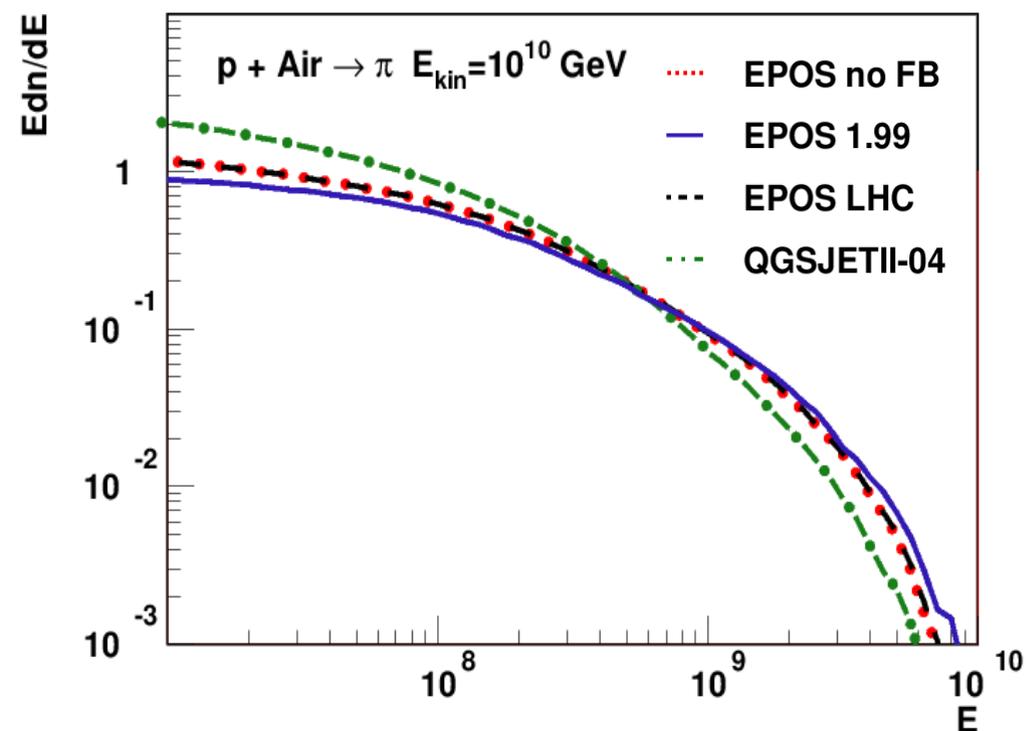
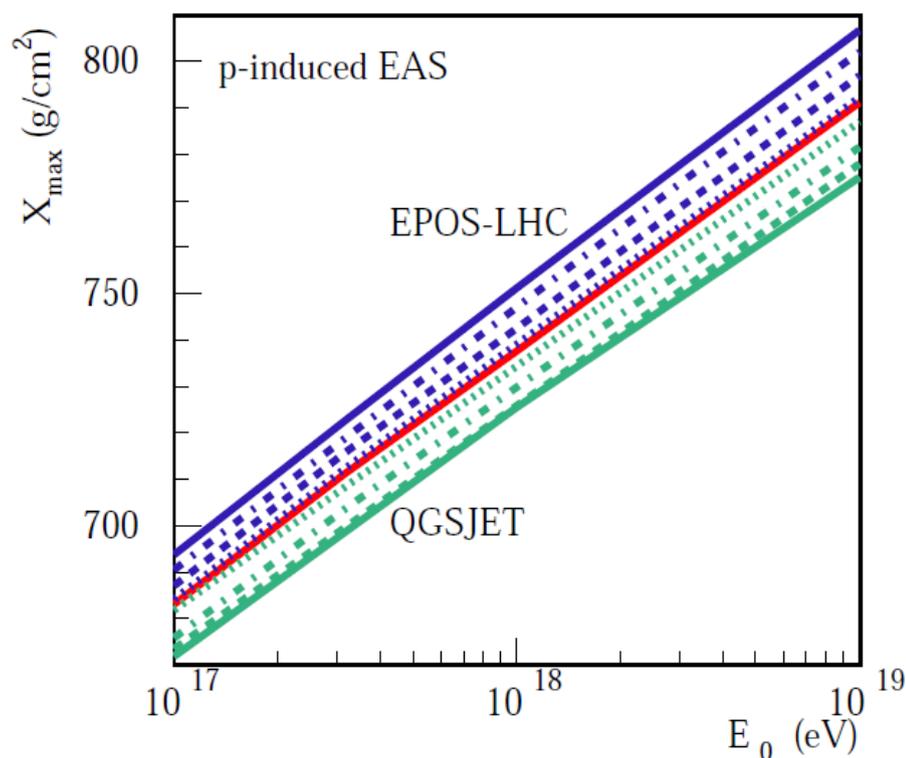
Mix models to identify source of differences

- Different mixing to extract useful information on X_{\max}
 - ➔ QII only for cross-section and nucleon spectra of 1st int. : dot-dashed



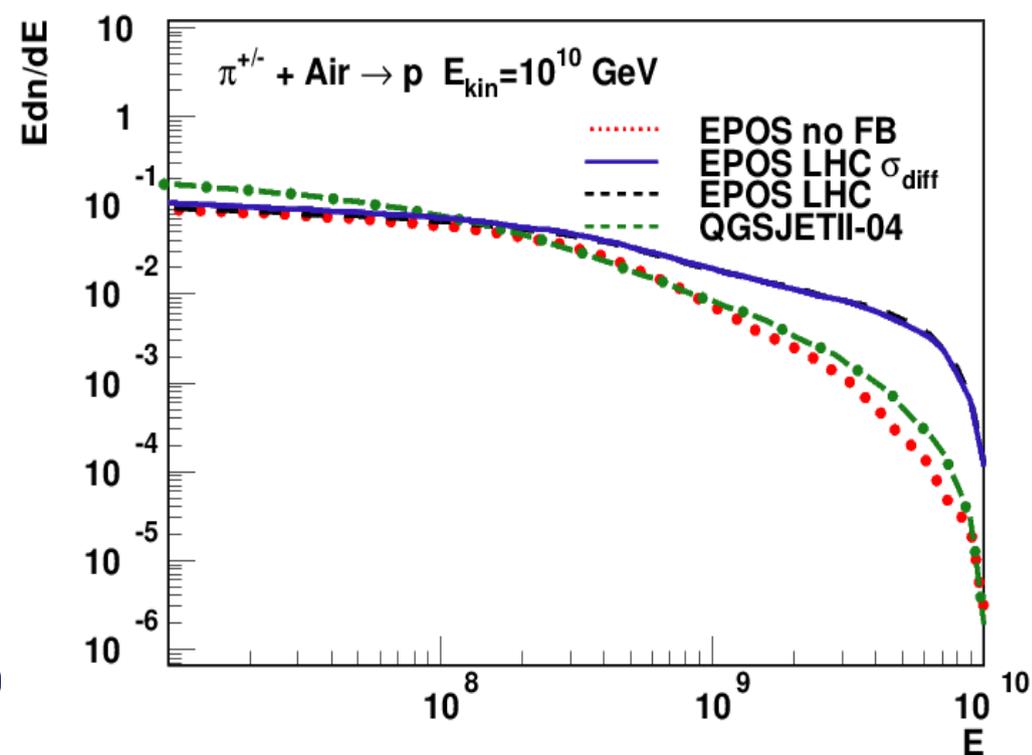
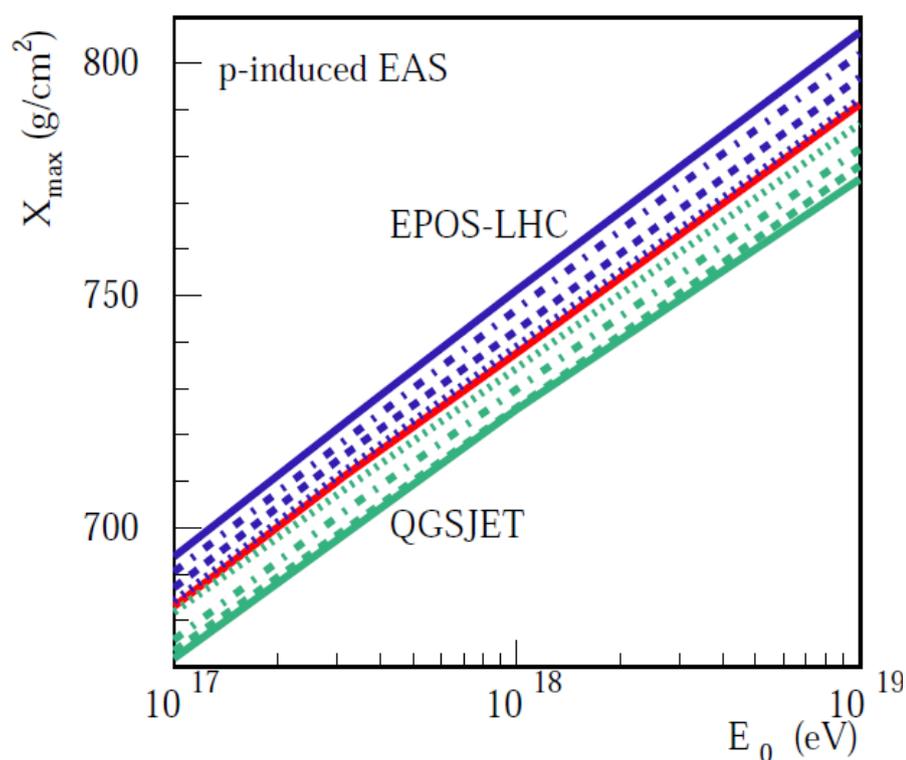
Mix models to identify source of differences

- Different mixing to extract useful information on X_{\max}
 - ➔ QII only for cross-section and nucleon spectra of 1st int. : dot-dashed
 - ➔ QII complete 1st int : dashed



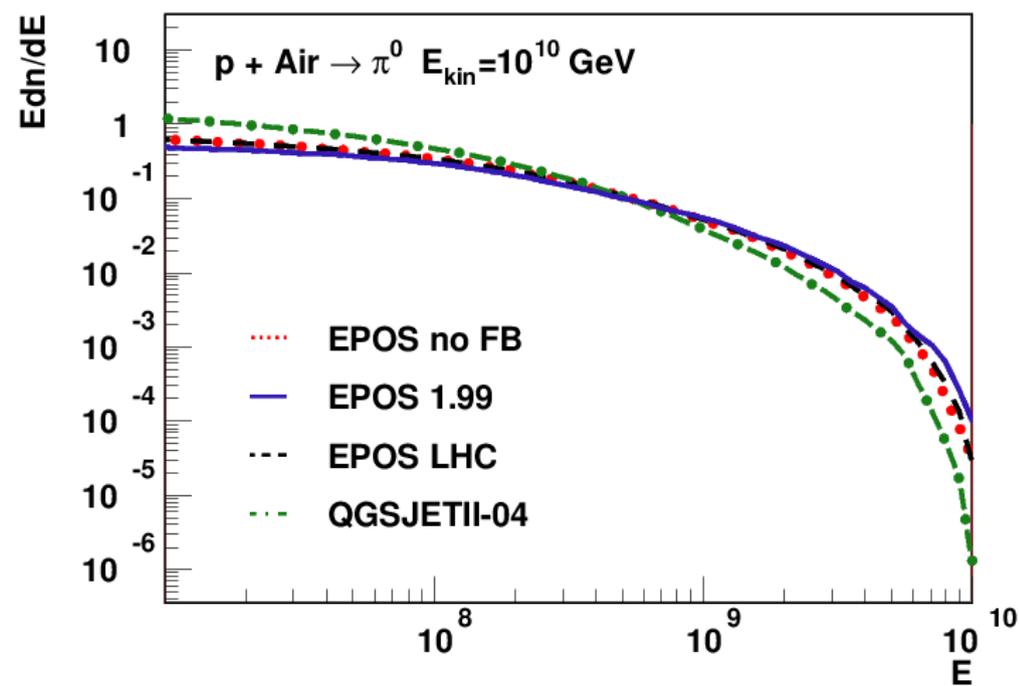
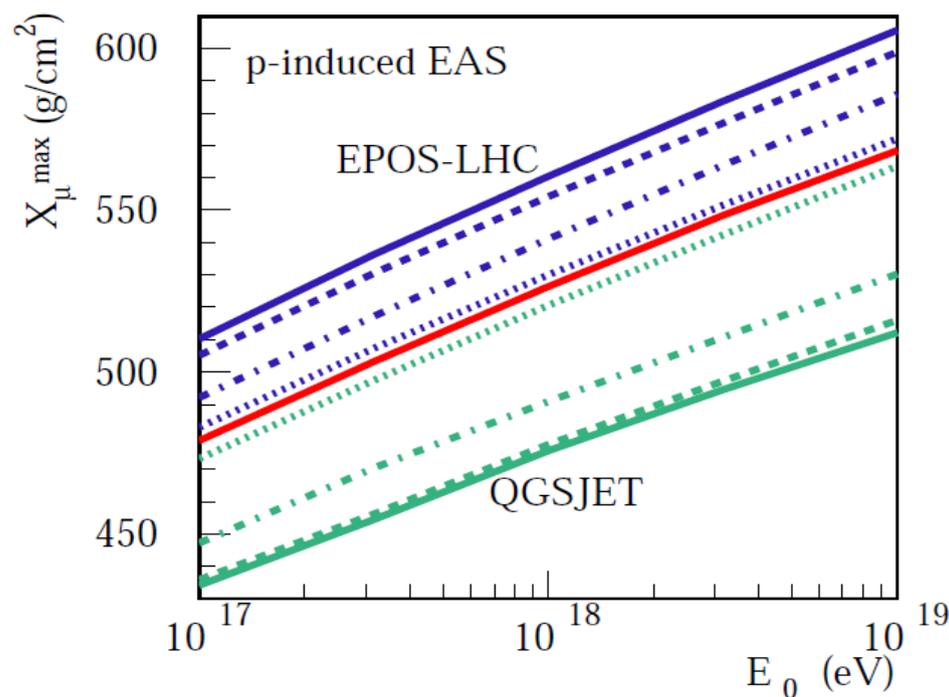
Mix models to identify source of differences

- Different mixing to extract useful information on X_{\max}
 - ➔ QII only for cross-section and nucleon spectra of 1st int. : dot-dashed
 - ➔ QII complete 1st int : dashed
 - ➔ QII complete 1st int and all nucleon prod. in the shower: dotted



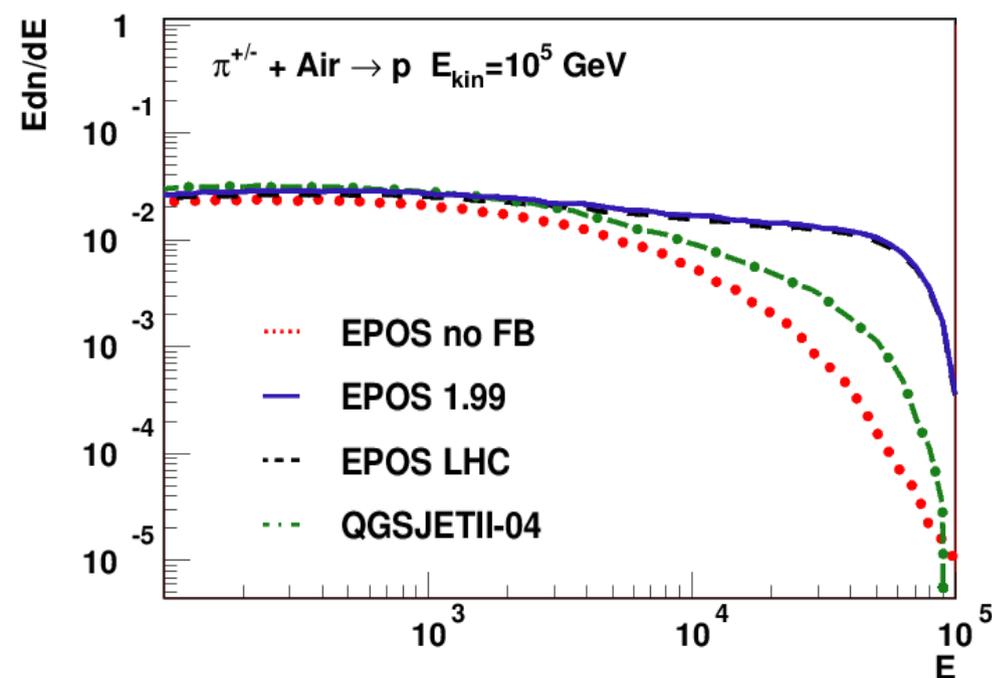
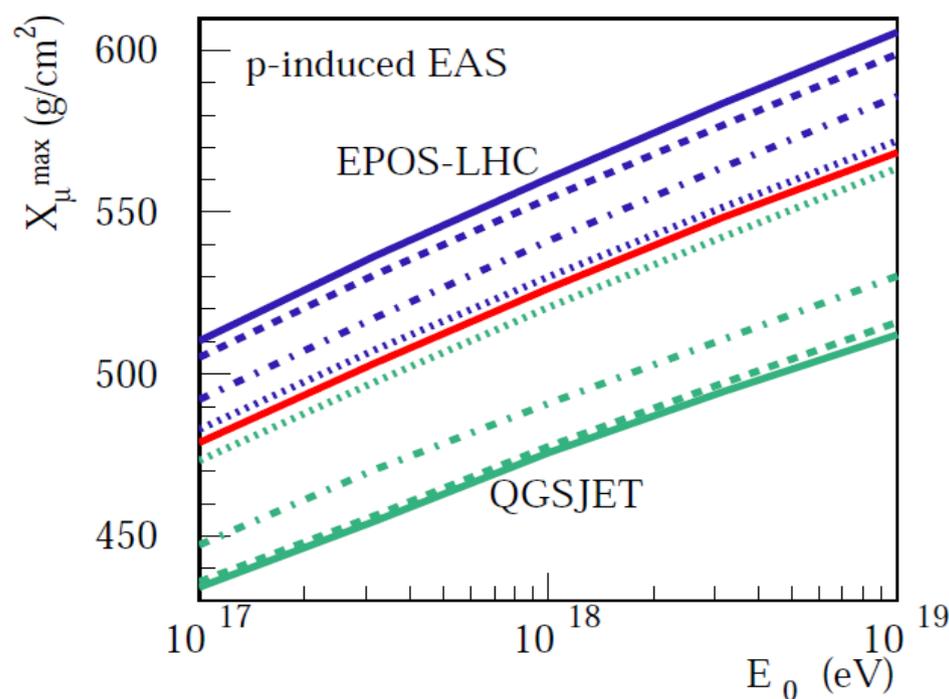
Mix models to identify source of differences

- Different mixing to extract useful information on X_{μ}^{\max}
 - ➔ QII complete 1st int. : dashed



Mix models to identify source of differences

- Different mixing to extract useful information on X_{μ}^{\max}
 - ➔ QII complete 1st int. : dashed
 - ➔ QII complete 1st int. and all nucleon prod. in the shower: dot-dashed



Mix models to identify source of differences

- Different mixing to extract useful information on X_{μ}^{\max}
 - ➔ QII complete 1st int. : dashed
 - ➔ QII complete 1st int. and all nucleon prod. in the shower: dot-dashed
 - ➔ QII complete 1st int. and hadron spectra in pion and kaon int.: dotted

