

I

Black hole degeneracies, Topological strings & quantum attractors

B. Pioline

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$$S_{\text{Thermo}} \stackrel{?}{\sim} \log \Omega_{\text{micro}}$$

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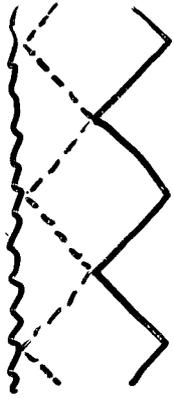
$$A_{\text{H}} / 4G + \dots$$

Try going beyond the thermodynamical limit, and understand the microscopic origin of subleading corrections to the Bekenstein - Hawking formula.

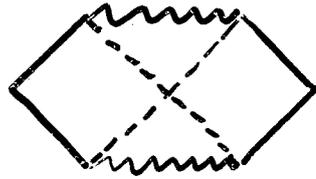
Ultimately, compute $\Omega_{\text{micro}}(Q)$ at finite charge.

Alas, not for realistic black holes yet...

Ext. Reissner-Nordström vs Schwarzschild



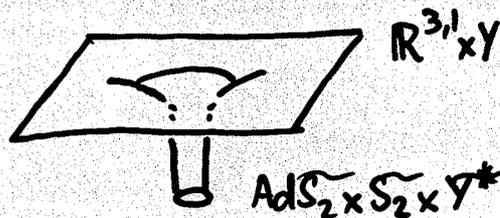
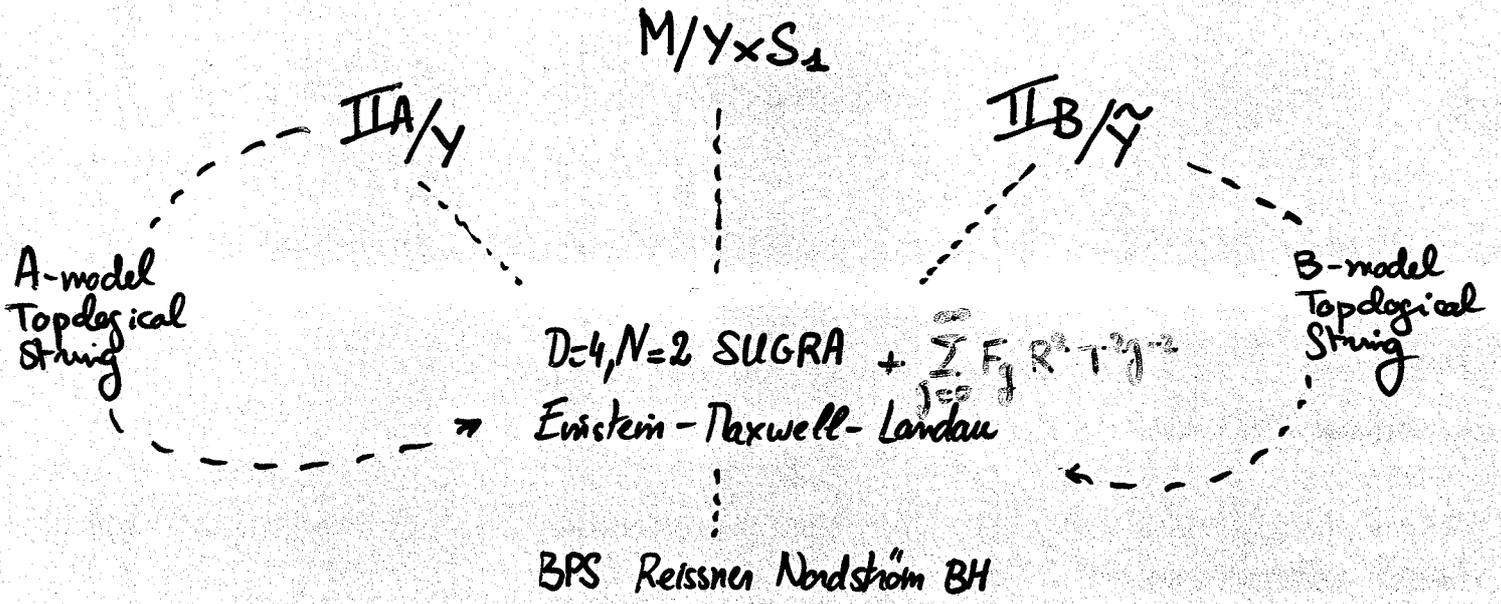
- time-like singularity
- $T=0$: stable
- near horizon $AdS_2 \times S^2$
- preserves $\frac{1}{2}$ SUSY in $N=2$



- space-like singularity
- $T > 0$: evaporates
- near horizon "Rindler₂ × S²"
- breaks all SUSY

$$Y = \begin{cases} \text{CY 3-fold} & \Rightarrow N=2 \\ K3 \times T^2 & \Rightarrow N=4 \\ T^6 & \Rightarrow N=8 \end{cases}$$

FLOW CHART



deformed Attractor flow

$$S_{BH, W}(P, q_i) = \frac{A_H}{4G} + a_1 \log \frac{A_H}{G} + b + c \left(\frac{G}{\hbar}\right) + \dots$$

SPECIFIC THERMODYNAMICAL ENSEMBLE

RADENACHER-CARDY: $\ln \Omega = 2\pi \sqrt{\frac{c}{6} N} + \tilde{a}_1 \log Q + \tilde{b}_1 + \tilde{c}_1/Q + \dots$

$D=1+1$ "black string CFT"

D2-D6-N5-mom

D1-D5-KK11-mom

M2-M5-momentum

Plan of the lectures

- [1. Special geometry and attractor flow
- [2. Very special geometry and black hole entropy in "homogeneous" SUGRA
- [3. Basics of topological string theory
- [4. Higher derivative corrections and the Ooguri-Vasiliev-Vafa conjecture
- [5. Precision counting of small black holes
- [6. The quantum attractor flow

Special geometry

* Recall that $N=2$ supersymmetry in 4 dimensions allows for two kinds of matter multiplets (besides the gravity multiplet):

• vector multiplets (A_μ, λ^i, z)

$z \in \mathcal{M}_V$, a Kähler-Hodge manifold of dimension $2n_V$
 $i=1,2$ R -symmetry index.

• hypermultiplets $(\psi, \begin{matrix} H \\ \tilde{H}^+ \end{matrix}, \tilde{\psi})$

$H, \tilde{H} \in \mathcal{M}_H$, a quaternionic-Kähler space of dimension $4n_H$

At two derivative order, the two interact only gravitationally.

The physics of BPS black holes is controlled by the vector multiplets, hypers are spectators.

* The couplings of vector multiplets in $N=2$ SUGRA are conveniently described by means of an holomorphic section $\Omega = (X^I(z), F_I(z))$ of a $Sp(2n_V+2)$ bundle over \mathcal{M}_V , defined up to a scaling $(X^I(z), F_I(z)) \approx (e^f(z) X^I(z), e^f F_I(z))$. This section is furthermore constrained to live in a lagrangian subspace of \mathbb{C}^{2n_V+2} , i.e. : LAGRANGIAN CONE

$$dX^I \wedge dF_I = 0 \quad \text{i.e.} \quad \langle d\Omega, d\Omega \rangle = 0$$

Thus, at generic points one may choose to parametrize \mathcal{M}_V by the "special" homogeneous coordinates X^I , $I=0 \dots n_V$, or by the inhomogeneous coordinates X^i/X^0 , $i=1 \dots n_V$. The remaining part of the section is then obtained by a "prepotential" $F(X^I)$ such that

$$F_I = \frac{\partial F}{\partial X^I}, \quad F(e^f X) = e^{2f} F(X)$$

(Note however that this choice obscures symplectic invariance)

* The metric on the manifold cb_V may be obtained from the Kähler potential

$$K(z, \bar{z}) = -\log(i(\bar{X}^I F_I - X^I \bar{F}_I)) = -\log i\langle \Omega, \bar{\Omega} \rangle$$

$$g_{i\bar{j}} = \partial_i \partial_{\bar{j}} K$$

This is well defined since, under rescalings $X^I \rightarrow e^{f(z)} X^I$, $e^{K/2} D_i \Omega$

$$K \rightarrow K - f(z) - f(\bar{z})$$

It will be convenient to define another section $U_i = (f_i^I, h_{iI})$

$$f_i^I = e^{-K/2} \partial_i (e^K X^I) = e^{K/2} (\partial_i X^I + \partial_i K X^I)$$

$$h_{iI} = e^{-K/2} \partial_i (e^K F_I) = e^{K/2} (\partial_i F_I + \partial_i K F_I)$$

(By assumption, $\partial_{\bar{i}} X^I = \partial_{\bar{i}} F_I = 0$)

Then the metric may be reexpressed as

$$g_{i\bar{j}} = -i \langle U_i, U_{\bar{j}} \rangle = i (f_i^I \bar{h}_{\bar{j}I} - h_{iI} \bar{f}_{\bar{j}}^I)$$

while the Riemann tensor of cb_V is

$$(*) \quad R_{i\bar{j}k\bar{l}} = g_{i\bar{j}} g_{k\bar{l}} + g_{i\bar{l}} g_{k\bar{j}} - C_{ikm} C_{\bar{j}\bar{l}\bar{m}} g^{m\bar{m}}$$

where the tensor C_{ijk} is defined as

$$C_{ijk} = \langle D_i U_j, U_k \rangle = C_{(ijk)}$$

In fact, (*) above is the definition of special geometry.

* The gauge kinetic terms may be written as

$$\mathcal{L} = -\frac{i}{2} \mathcal{N}_{IJ} F_{\mu\nu}^{-I} F^{-J\mu\nu} + cc$$

where

$$\mathcal{N}_{IJ} = \bar{\tau}_{IJ} + 2i \frac{(\text{Im} \tau X)_I (\text{Im} \tau X)_J}{X \cdot (\text{Im} \tau) X}$$

NOTE :

$$F_I = \mathcal{N}_{IJ} X^J$$

$$h_{iI} = \mathcal{N}_{IJ} f_i^J$$

$$\tau_{IJ} := \partial_I \partial_J F$$

In particular, the gauge couplings are $\frac{1}{g_{IJ}^2} = -\text{Im} \mathcal{N}_{IJ}$: positive definite

The symplectic symmetries may be made manifest by defining

$$g_{\mu\nu I} = \bar{N}_{IJ} F_{\mu}^J$$

The conserved electric and magnetic charges are then

$$\begin{pmatrix} p^I \\ q_I \end{pmatrix} = \int_{S_2} \begin{pmatrix} F^I \\ g_I \end{pmatrix}$$

The graviphoton is the linear combination

$$T_{\mu\nu} = -2i (X \cdot \text{Im} N \cdot F_{\mu\nu}) \cdot e^{K/2}$$

Its associated charge

$$Z = \int_{S_2} T = e^{K/2} (q_I X^I - p^I F_I)$$

is the central charge of the $N=2$ superalgebra.

BPS states satisfy $M_{\text{BPS}}^2 = |Z|^2$

* Electric-magnetic duality acts by symplectic rotations

$$\Omega \rightarrow S \Omega \quad \text{where} \quad S = \begin{pmatrix} A & B \\ C & D \end{pmatrix} \quad \begin{aligned} A^t D - C^t B &= 1 \\ A^t C &= C^t A \\ B^t D &= D^t B \end{aligned}$$

preserving the symplectic product.

Since $F(X) = \frac{1}{2} X^I F_I$ by homogeneity,

$$\begin{aligned} \hat{F}(\tilde{X}) &= \frac{1}{2} (X^t A^t + F^t B^t) (C X + D F) \\ &= F(X) + \frac{1}{2} X^t (C^t A) X + \frac{1}{2} F^t (D^t B) F + X^t (C^t B) F \end{aligned}$$

where $\tilde{X} = A X + B F$.

This is characteristic of the transformation of a generating function of a Lagrangian submanifold in symplectic geometry.

In particular, for $S = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}$, $\hat{F}(\tilde{X}) = F(X) - X^t \tilde{X}$ is the Legendre transform of $F(X)$. And in fact, by homogeneity, $\tilde{F}(\tilde{X}) = -F(X)$.

Type IIA compactified on a CY:

$h = h_{1,1} + 1$ vector multiplets, coming from reducing the RR 3 forms on $H_2(Y)$. The graviphoton comes from the RR 1-form in 10D.

To leading order,

$$F = -\frac{1}{6} C_{ABC} \frac{X^A X^B X^C}{X^0}; K = -\log V$$

The complex scalars are $\frac{X^A}{X^0} = \int_{\gamma^A} J$, $\frac{F_A}{X^0} = i \int_{\gamma^A} J \wedge J$, complexified Kähler structures $J = B_{25} + i\omega_{K1/8}$

where C_{ABC} is the intersection form on $H_2 \otimes H_2 \otimes H_2$: $\gamma^A \in H_2$, $\gamma_A \in H_4$

In addition, there are worldsheet instanton corrections known as Gromov-Witten instantons, indexed by rational curves in Y

$$F_{inst} = X^0{}^2 \sum_{\beta \in H_2(Y, \mathbb{Z})} N_{g,\beta} e^{2\pi i X^0 \beta / X^0}$$

$$Z = e^{K/2} X^0 \left(q_0 + q_A \int_{\gamma_A} J - p^A \int_{\gamma_A} J \wedge J - p^0 \int_{\gamma_A} J \wedge J \wedge J \right)$$

Type IIB compactified on CY Y :

$h = h_{2,1}$ vector multiplets, coming from the reduction of the RR 4 form (with self dual 5-form field strength) on $H_3(Y)$.

$$X^I = \int_{A^I} \Omega, F_I = \int_{B_I} \Omega$$

where (A^I, B_I) is a symplectic basis of $H_3(Y)$

$$K = -\log \int \Omega \wedge \bar{\Omega}$$

$$= -\log i(\bar{X}^I F_I - X^I \bar{F}_I) \text{ by Riemann's bilinear relations}$$

$$Z = \frac{\int_Y \Omega}{\sqrt{\int \Omega \wedge \bar{\Omega}}}$$

$$\gamma = q_I A^I - p^I B_I$$

gives the mass of D3 branes wrapped on calibrated (ie special Lagrangian) 3-cycles γ .

Rk: $m_{SFS} = |z| m_p = \left(\frac{V_{CY}}{l_s^6}\right)^{1/2} \frac{1}{g_s l_s} \cdot e^{K/2} (q_0 X^0 + \dots)$ so indeed $K = -\log(V_{CY}/l_s^6)$ in the gauge $X^0 = 1$

The attractor mechanism

Consider a general ansatz for a static, spherically symmetric black hole in $N=2$ sugra:

$$ds^2 = -e^{2u(r)} dt^2 + e^{-2u(r)} (dr^2 + r^2 d\Omega_2^2)$$

$$F^{I-} = \frac{1}{2} \left(p^I + i (\text{Im} \mathcal{N}_{JK}^{-1})^{IJ} (\text{Re} \mathcal{N}_{JK} p^K - q^J) \right) \cdot \left[\sin\theta d\theta d\varphi - \frac{i e^{2u}}{r^2} dt dr \right]$$

Assume that the solution preserves one susy:

then the metric factor $u(r)$ and scalars z are entirely determined by the 1st order equations:

$$\begin{cases} du/dr = |z| e^u / r^2 \\ dz^i/dr = -2 g^{i\bar{j}} e^u \partial_{\bar{j}} |z| / r^2 \end{cases}$$

where $Z = e^{K/2} (q_I X^I - p^I F_I)$ is the central charge.

Defining $\mu = e^{-u}$, the second equation becomes

$$\mu \frac{dz^i}{d\mu} = -g^{i\bar{j}} \partial_{\bar{j}} \log |z|^2$$

which is a gradient flow on \mathcal{M}_V with potential $\log |z|^2$.

* Near horizon : $r \rightarrow 0$

• $dz^i/dr = 0 \Rightarrow \partial_i |z| = \partial_{\bar{j}} |z| = 0$

• $e^{-u} du = |z| dr/r^2 \Rightarrow e^{-u} = \text{const} + \frac{|z|}{r}$

The metric becomes

$$ds^2 = - \frac{r^2}{|z|^2} dt^2 + \frac{|z|^2}{r^2} (dr^2 + r^2 d\Omega_2^2)$$

Define $u = |z|^2/r$

$$ds^2 = |z|^2 \left(- \frac{dt^2 + du^2}{u^2} + d\Omega_2^2 \right) : \text{AdS}_2 \times S_2$$

The black hole entropy is thus

$$S = \frac{1}{4} \cdot 4\pi |z|^2 = \pi |z|^2$$

where z is the central charge evaluated at the scalars which minimize i

* The attractor equations may be recast as follows:

$$Z = e^{K/2} (q_I X^I - p^I F_I)$$

$$f_i^I = e^{K/2} (\partial_i X^I + \partial_i K X^I)$$

hence

$$\partial_i Z = q_I (f_i^I - e^{K/2} \partial_i K X^I) - p^I (h_{iI} - e^{K/2} \partial_i K F_I) + \frac{1}{2} Z \partial_i K$$

$$\partial_{\bar{i}} \bar{Z} = \frac{1}{2} \partial_{\bar{i}} K \bar{Z}$$

$$\partial_{\bar{i}} Z = f_i^I (q_I - \bar{N}_{I\bar{J}} p^{\bar{J}}) - \frac{1}{2} Z \partial_{\bar{i}} K$$

$$\frac{\partial_i |z|}{|z|} = \frac{1}{2} \left(\frac{\partial_i Z}{Z} + \frac{\partial_{\bar{i}} \bar{Z}}{\bar{Z}} \right) = \frac{1}{2} \frac{f_i^I}{Z} (q_I - \bar{N}_{I\bar{J}} p^{\bar{J}})$$

The attractor point is thus determined by $q_I = \bar{N}_{I\bar{J}} p^{\bar{J}}$

$$\text{or } q_I = \left(\tau_{I\bar{J}} - 2i \frac{(\text{Im} \tau \bar{X})_I (\text{Im} \tau \bar{X})_{\bar{J}}}{\bar{X} \text{Im} \tau \bar{X}} \right) p^{\bar{J}}$$

This is rewritten

$$q_I - \tau_{I\bar{J}} p^{\bar{J}} = -i C (\text{Im} \tau)_{I\bar{J}} \bar{X}^{\bar{J}}$$

$$\text{where } C = +2 \frac{p \cdot \text{Im} \tau \cdot \bar{X}}{\bar{X} \cdot \text{Im} \tau \cdot X}$$

Combining this equation with its complex conjugate:

$$q - \tau p = -iC (\text{Im}\tau) \bar{X}$$

$$q - \bar{\tau} p = +i\bar{C} (\text{Im}\tau) X$$

$$\begin{aligned} &\Updownarrow \\ &\text{Re}(CX^I) = p^I \quad \Leftrightarrow \quad CX^I + \bar{C}\bar{X}^I = 2p^I \\ &\text{Re}(CF_I) = q_I \quad \Leftrightarrow \quad C\tau_{IJ}X^J + \bar{C}\bar{\tau}_{IJ}\bar{X}^J = 2q_I \end{aligned}$$

Using $\text{Im}\mathcal{N} = -\text{Im}\tau + \frac{\text{Im}\tau X \otimes \text{Im}\tau X}{X \text{Im}\tau X} - \frac{\text{Im}\tau \bar{X} \otimes \text{Im}\tau \bar{X}}{\bar{X} \text{Im}\tau \bar{X}}$

one easily concludes $\text{Im}\mathcal{N} \cdot X = -\frac{(X \text{Im}\tau \bar{X})}{\bar{X} \text{Im}\tau \bar{X}} \text{Im}\tau \bar{X}$

hence $C = -2 \frac{p \cdot \text{Im}\mathcal{N} \cdot X}{X \text{Im}\tau \bar{X}}$

Furthermore $K = -\log(-2 \bar{X} \text{Im}\tau \bar{X})$
 $Z = -2i e^{K/2} (p \cdot \text{Im}\mathcal{N} \cdot X)$

hence $C = 2i e^{K/2} Z$

We may fix the gauge by requiring $|C|=1$:

$$e^{-K/2} = \frac{|Z|}{2} \sim \frac{r e^{-u}}{2} \quad \text{near the horizon}$$

In this gauge, the area of the sphere is related to the Kähler potential via

$$A(r) = 4\pi r^2 e^{-2u} = \pi e^{-K}$$

The phase of C is related to the N=1 SUSY preserved ; for single centered black holes, there is no loss of generality in assuming that $C=1$.

The attractor equations are therefore most easily summarized as

$$\left[\begin{array}{l} \text{Re}(X^I) = p^I \\ \text{Re}(\bar{F}_I) = q_I \end{array} \right. \quad S_{\text{BH}} = \frac{\pi}{4} \cdot i(\bar{X}^I F_I - X^I \bar{F}_I)$$

Attractor mechanism and Legendre transform

The first equation $\text{Re}(X^I) = p^I$ may be solved by setting $X^I = p^I + i\varphi^I$
 $\bar{X}^I = p^I - i\varphi^I$

The second equation then reads

$$q_I = \frac{1}{2} \left(\frac{\partial F}{\partial X^I} + \frac{\partial \bar{F}}{\partial \bar{X}^I} \right) = \frac{1}{2i} \left(\frac{\partial F}{\partial \varphi^I} - \frac{\partial \bar{F}}{\partial \varphi^I} \right) = -\frac{1}{\pi} \frac{\partial \mathcal{F}}{\partial \varphi^I}$$

Where we define $\mathcal{F}(p, \varphi) = \pi \text{Im} [F(p + i\varphi)]$

The entropy is then

$$\begin{aligned} S_{\text{BH}} &= \frac{i\pi}{4} \left[(X^I - 2i\varphi^I) F_I - (\bar{X}^I + 2i\varphi^I) \bar{F}_I \right] \\ &= \frac{i\pi}{2} (F - \bar{F}) + \frac{\pi}{2} \varphi^I (F_I + \bar{F}_I) \\ &= \mathcal{F}(p, \varphi) + \pi \varphi^I q_I \end{aligned}$$

so that S_{BH} is the Legendre transform of the "free energy" $\mathcal{F}(p, \varphi)$ with respect to the "electric potential" φ^I !

We will see in lect 3 that this continues to hold when higher-derivative F-term interactions are included.

Ex: leading entropy for black holes with $p^0=0$

$$F = -\frac{1}{6} C_{ABC} \frac{X^A X^B X^C}{X^0}$$

$$\mathcal{F} = \frac{\pi}{6} \operatorname{Im} \frac{(p^A + i\varphi^A)(p^B + i\varphi^B)(p^C + i\varphi^C)}{p^0 + i\varphi^0} C_{ABC}$$

$$= -\frac{\pi}{6} \operatorname{Re} \frac{(p^A + i\varphi^A)(p^B + i\varphi^B)(p^C + i\varphi^C)}{\varphi^0} C_{ABC}$$

$$= -\frac{\pi}{6} \frac{C(p)}{\varphi^0} + \frac{\pi}{2} \frac{C_{AB}(p) \varphi^A \varphi^B}{\varphi^0}$$

where $C(p) = C_{ABC} p^A p^B p^C$

$C_{AB}(p) = C_{ABC} p^C$

$$S_{BH} = \langle \mathcal{F} + \pi(\varphi^0 q_0 + \varphi^A q_A) \rangle_{\varphi^0, \varphi^A}$$

The extremization wrt φ^A is Gaussian: $\varphi_*^A = -C^{AB}(p) q_B \varphi^0$

$$S_{BH} = \langle -\frac{\pi}{6} \frac{C(p)}{\varphi^0} + \frac{\pi}{2} \varphi^0 q_A C^{AB}(p) q_B + \pi \varphi^0 q_0 \rangle_{\varphi^0}$$

Define $\hat{q}_0 = q_0 + \frac{1}{2} q_A C^{AB}(p) q_B$:

$$S_{BH} = \langle -\frac{\pi}{6} \frac{C(p)}{\varphi^0} + \pi \varphi^0 \hat{q}_0 \rangle_{\varphi^0}$$

This is stationary at $\varphi_*^0 = \pm \sqrt{\frac{-\hat{C}(p)}{6 \hat{q}_0}}$ if $\hat{C}(p) > 0$
 $\hat{q}_0 < 0$

Choosing the + sign,

$$S_{BH} = 2\pi \sqrt{-\frac{\hat{C}(p)}{6} \hat{q}_0}$$

Choosing $q_A=0$, we get more simply

$$S_{BH} = 2\pi \sqrt{-C_{ABC} p^A p^B p^C q_0}$$

Very special supergravities

①

For our purpose to understand black hole entropy in $N > 2$ supergravity, it will be useful to consider a class of $N=2$ theories where the moduli space takes a particularly simple form, as a symmetric space G/H . Whether or not these theories exist as true quantum theories is still unknown.

The simplest way of constructing them is to start from 5 dimensions:

$$S = \int d^5x \sqrt{g} \left(R - a_{AB} F_{\mu\nu}^A F^{\mu\nu B} + g_{ij} \partial\phi^i \partial\phi^j \right) + \int C_{ABC} A^A \wedge F^B \wedge F^C$$

8 susy \Rightarrow C_{ABC} are constants

ϕ^i take values in hypersurface $\mathcal{N}=1$, where

$$\mathcal{N} = C_{ABC} \xi^A \xi^B \xi^C, \text{ with metric } g_{ij} \text{ induced from } \mathbb{R}^n$$

If one furthermore requires that the moduli space is a symmetric space, one finds that $\mathcal{N}(\xi)$ should be the norm of a Jordan algebra of degree 3:

$$\bullet j_1 \circ j_2 = j_2 \circ j_1$$

$$\bullet j_1 \circ (j_2 \circ j_2^2) = (j_1 \circ j_2) \circ j_1^2$$

Solutions were classified by Jordan, Von Neuman, Wigner '34:

1. \mathbb{R} , $\mathcal{N}(x) = x^3$

2. $\mathbb{R} \oplus \Gamma$, Γ Clifford algebra, $\mathcal{N}(x) = x_1 x^a Q_{ab} x^b$

3. 3×3 hermitean matrices over $\mathbb{R}, \mathbb{C}, \mathbb{H}, \mathbb{O}$ ("division algebras")

$$\bar{J} = \begin{pmatrix} \alpha_1 & x_3 & \bar{x}_2 \\ \bar{x}_3 & \alpha_2 & x_1 \\ x_2 & \bar{x}_1 & \alpha_3 \end{pmatrix} \quad \begin{array}{l} \alpha_i \in \mathbb{R} \\ x_i \in \mathbb{K} \end{array}$$

$$\bar{J}_1 \circ \bar{J}_2 = \bar{J}_1 \cdot \bar{J}_2 + \bar{J}_2 \bar{J}_1$$

$$* \quad \mathcal{N}(J) = \alpha_1 \alpha_2 \alpha_3 - \alpha_1 (x_1 \bar{x}_1) - \alpha_2 (x_2 \bar{x}_2) - \alpha_3 (x_3 \bar{x}_3) + 2 \operatorname{Re}(x_1 x_2 x_3)$$

One may then show that the moduli space is

$$\mathcal{M}_J = \operatorname{Str}_0(J) / \operatorname{Aut}(J)$$

where $\operatorname{Str}_0(J)$ is the "reduced structure group" of J ,
 is the group preserving \mathcal{N}
 $\operatorname{Aut}(J)$ is the automorphism group of J .

*: More explicitly:

J	\mathcal{N}	$\operatorname{Str}_0(J)$
$J_3^{\mathbb{R}}$	$\det(3 \otimes 3)$	$SL(3, \mathbb{R})$
$J_3^{\mathbb{C}}$	$\det(3 \otimes 3)$	$SL(3, \mathbb{C})$
$J_3^{\mathbb{H}}$	$\operatorname{Pf}(6 \wedge 6)$	$SU^*(6)$
$J_3^{\mathbb{O}}$	$(27 \otimes 27 \otimes 27) _5$	$E_6(-26)$

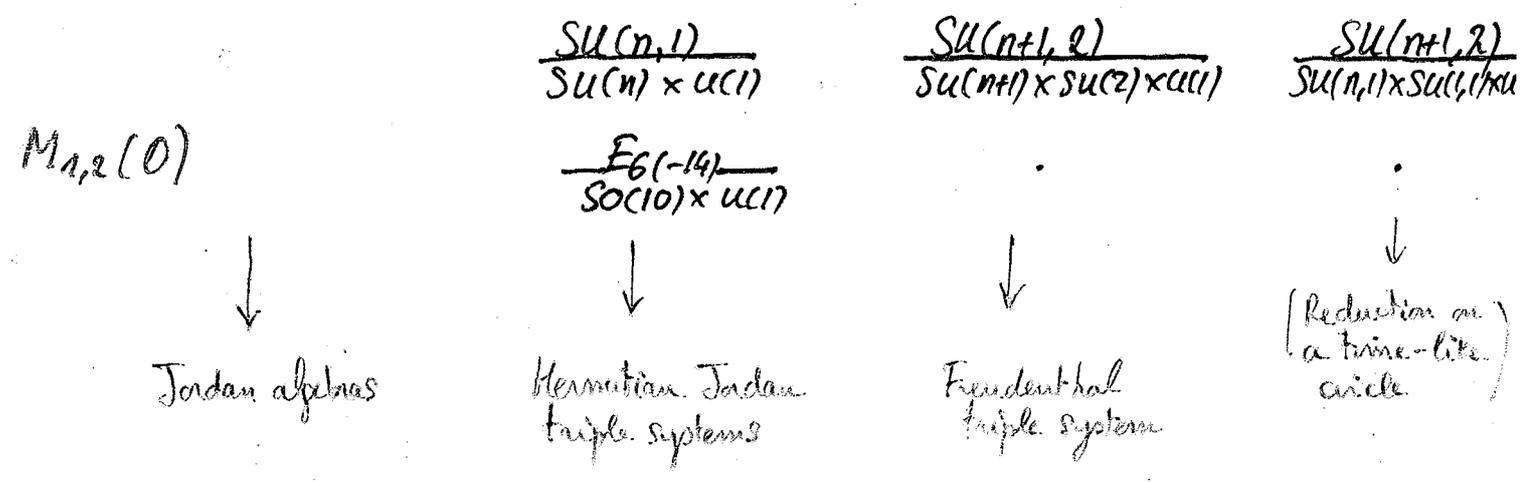
Very special geometry and Jordan algebras

$(N=2)$	D=5	D=4	D=3	D=3*
$\Gamma_{n-1,1}$	$\frac{SO(n-1,1) \times SO(1,1)}{SO(n-1)}$	$\frac{SO(n,2)}{SO(n) \times SO(2)} \times \frac{SU(2)}{U(1)}$	$\frac{SO(n+2,4)}{SO(n+2) \times SO(4)}$	$\frac{SO(n+2,4)}{SO(n,2) \times SO(2,2)}$
\mathbb{R}	$SO(1,1)$	$\frac{SU(2)}{U(1)}$	$\frac{G_2(2)}{SO(4)}$	$\frac{G_2(2)}{SO(2,2)}$
$J_3^{\mathbb{R}}$	$\frac{SU(3)}{SO(3)}$	$\frac{Sp(6)}{SU(3) \times U(1)}$	$\frac{F_4(4)}{USp(6) \times SU(2)}$	$\frac{F_4(4)}{Sp(6) \times SU(1,1)}$
$J_3^{\mathbb{C}}$	$\frac{SU(3, \mathbb{C})}{SU(3)}$	$\frac{SU(3,3)}{SU(3) \times SU(3) \times U(1)}$	$\frac{E_6(2)}{SU(6) \times SU(2)}$	$\frac{E_6(2)}{SU(3,3) \times SU(1,1)}$
$J_3^{\mathbb{H}}$	$\frac{SU^*(6)}{USp(6)}$	$\frac{SO^*(12)}{SU(6) \times U(1)}$	$\frac{E_7(-25)}{SO(12) \times SU(2)}$	$\frac{E_7(-25)}{SO^*(12) \times SU(1,1)}$
$J_3^{\mathbb{O}}$	$\frac{E_6(-26)}{F_4}$	$\frac{E_7(-25)}{E_6 \times U(1)}$	$\frac{E_8(-24)}{E_7 \times SU(2)}$	$\frac{E_8(-24)}{E_7(-25) \times SU(1,1)}$

$(N=8)$				
$J_3^{\mathbb{O}_8}$	$\frac{E_6(6)}{USp(8)}$	$\frac{E_7(7)}{SU(8)}$	$\frac{E_8(8)}{SO(16)}$	$\frac{E_8(8)}{SO^*(16)}$

$(N=4)$				
$\Gamma_{n-5,5}$	$\frac{SO(n-5,5) \times SO(1,1)}{SO(n-5) \times SO(5)}$	$\frac{SO(n-4,6)}{SO(n-4) \times SO(6)} \times \frac{SU(2)}{U(1)}$	$\frac{SO(n-2,8)}{SO(n-2) \times SO(8)}$	$\frac{SO(n-2,8)}{SO(n-4,2) \times SO(6)}$

In addition, 4D models which cannot be obtained from SD:



Reduction 4D \rightarrow 3D:

$$e_{SD} = \begin{pmatrix} e^\sigma & e^\sigma B_\mu \\ 0 & e^{-\sigma/2} e_{\mu\nu} \end{pmatrix} \quad A_{SD}^A = (a^A, A_\mu^I) \\ B_\mu := A_\mu^0$$

Identify $C_{ABC} \xi^A \xi^B \xi^C = e^{3\sigma}$

Complexify $t^A = a^A + i \xi^A$

One finds that the scalar fields t^A, \bar{t}^A are inhomogeneous coordinates $t^A = z^A/z^0$ for a special Kähler manifold with prepotential

$$F = \frac{C_{IJK} z^I z^J z^K}{z^0}$$

The moduli space is a symmetric space

$$\mathcal{M}_4 = \frac{\text{Conf}(\mathcal{J})}{\widehat{\text{St}}_0(\mathcal{J}) \times U(1)}$$

where $\text{Conf}(\mathcal{J})$ is the group leaving the "light cone" $\mathcal{N}(\xi) = 0$ invariant, and $\widehat{\text{St}}_0(\mathcal{J})$ is its max compact subgroup (in fact, the compact version of $\text{St}_0(\mathcal{J})$)

Note that $\text{Conf}(\mathcal{J})$ acts linearly on the electric-magnetic charges p^0, p^A, q_A, q_0

Its action preserves the Lagrangian cone $\{F_I^I = \partial F / \partial x^I\}$

Therefore, it acts on \mathcal{M}_4 by linear fractional transformations.

A remark on Legendre invariance

- An important property following from the “adjoint identity” of Jordan algebras

$$X^{\#\#} = N(X)X, \quad X_A^\# := C_{ABC}X^B X^C$$

is that F is invariant under Legendre transform in all variables:

$$\langle N(X)/X^0 + X^0 Y_0 + X^A Y_A \rangle_{XI} = -N(Y)/Y^0$$

Proof: saddle point at $Y_A = X_A^\# / X^0$, $Y_0 = -N(X)/(X^0)^2$, hence

$$\begin{aligned} N(X)X^A &= (X^0 Y_A)^\# = (X^0)^2 (Y^A)^\# \Rightarrow X^A = -Y_A^\# / Y^0 \\ N(Y)Y_A &= (-X^A Y_0)^\# \Rightarrow X^0 = N(Y)/(Y_0)^2 \end{aligned}$$

In fact, $(X^0)^\alpha N(X)^\beta e^{iN(X)/X^0}$ is invariant under Fourier transform, for some choice of α, β !

BH entropy in very special SUGRA

- As an illustration of the OSV fact, let us compute the tree-level entropy of a black hole with arbitrary charges in very special SUGRA. The free energy is

$$\mathcal{F}(p, \phi) = \frac{\pi}{(p^0)^2 + (\phi^0)^2} \left\{ p^0 \left[\phi^A p_A^\# - I_3(\phi) \right] + \phi^0 \left[p^A \phi_A^\# - I_3(p) \right] \right\}$$

- In order to eliminate the quadratic term in ϕ^A , change variables to

$$x^A = \phi^A - \frac{\phi^0}{p^0} p^A, \quad x^0 = [(p^0)^2 + (\phi^0)^2] / p^0$$

and, so as to eliminate the square root in $q_0 \phi^0$, introduce an auxiliary variable t ,

$$S = \pi \left\langle -\frac{I_3(x)}{x^0} + \frac{p_A^\# + p^0 q_A}{p^0} x - \frac{t}{4} \left(\frac{x^0}{p^0} - 1 \right) - \frac{(2I_3(p) + p^0 p^I q_I)^2}{t (p^0)^2} \right\rangle_{\{x^I, t\}}$$

BK entropy and 4D/5D lift

(6)

Using Legendre invariance, we find

$$S = \pi \left\langle 4 \frac{I_3(p_A^\# + p^0 q_A) - [2I_3(p) + p^0 p^I q_I]^2 - \frac{t}{4}}{(p^0)^2 t} \right\rangle_t$$

$$= \frac{\pi}{|p^0|} \sqrt{4 I_3[p_A^\# + p^0 q_A] - [2I_3(p) + p^0 p^I q_I]^2}$$

This equation has a simple interpretation:

NOTE THAT
SD U-duality $Str_0(J)$
IS MANIFEST

$$S_{4D} = \frac{1}{|p^0|} S_{5D}(Q_A, J)$$

where Q_A, J are the electric charges and angular momentum of a 5D black hole obtained by lifting IIA/CY to M-theory/CY at strong coupling:

$$Q_A = p^0 q_A + C_{ABC} p^B p^C$$

$$2J = (p^0)^2 q_0 + p^0 p^A q_A + 2I_3(p)$$

Conversely, the 4D black hole may be obtained by "compactifying" $M / \mathbb{R}_t \times TN_{p^0} \times CY$ with a 5D black hole at the origin of TN

The TN space has a circle of fixed size at ∞ : translations along this circle correspond to $U(1) \subset SU(2)_L$ rotations around the center.

The charge $p^0 TN$ is obtained by quotienting by $\mathbb{Z}/p^0\mathbb{Z}$: this changes the radius to R/p^0 , and the horizon to S^3/p^0 :

$$q_0 = \frac{2J}{(p^0)^2} \quad ; \quad q_A = Q_A/p^0$$

The above formula suggests that this construction works even with $p^A \neq 0$. $I_3(p) \neq 0$ implies new contributions to the angular momentum J .

Continuing to use the special properties of $I_3(X)$, one may Taylor expand

$$\begin{aligned}
 I_3(p_A^\# + p^0 q_A) &= I_3(p_A^\#) + p^0 q_A \partial^A I_3(p_A^\#) \\
 &\quad + (p^0)^2 p_A^\# \partial_A I_3(p_A^\#) \\
 &\quad + (p^0)^3 I_3(p_A^\#) \\
 &= [I_3(p^A)]^2 + p^0 I_3(p^A) p^A q_A \\
 &\quad + (p^0)^2 p_A^\# q_A^\# + (p^0)^3 I_3(q_A)
 \end{aligned}$$

This allows to rewrite the entropy as

$$S = \pi \sqrt{I_4(p, q)}$$

$$\begin{aligned}
 \text{where } I_4(p, q) &= 4 p^0 I_3(q) - 4 q_0 I_3(p) \\
 &\quad + 4 p_A^\# q_A^\# - (p^0 q_0 + p^A q_A)^2
 \end{aligned}$$

This is the quartic invariant of $\text{Conf}(J)$ which follows from Freudenthal's triple system construction

$$\begin{pmatrix} p^0 & p^I \\ q_I & q_0 \end{pmatrix} \quad p^0, q_0 \in \mathbb{R}, \quad p^I, q_I \in J$$

4D u-duality is thus "manifest".

For $J = J_3^{\mathbb{Q}}$:	$J = \mathbb{R} \oplus \Pi :$
$I_3 =$ cubic invariant of $E_6(-26)$	$I_3 =$ cubic inv of $SO(1, n+4)$
$I_4 =$ quartic invariant of $E_7(-25)$	$I_4 =$ quartic $sl(2) \times SO(2, n+2)$

but the same algebra applies for $N=8$ SUGRA :	$N=4$ SUGRA
$I_3 =$ cubic invariant of $E_6(6)$	$SO(5, n-3)$
$I_4 =$ quartic invariant of $E_7(7)$	$sl(2) \times SO(6, n-2)$

Basic topological string theory

1. Standard N=1 σ -model

$$L = 2t \int d^2z \left(\frac{1}{2} g_{I\bar{J}}(\phi) \partial_z \phi^I \partial_{\bar{z}} \phi^{\bar{J}} + \frac{i}{2} g_{I\bar{J}} \psi_-^I D_z \psi_-^{\bar{J}} + \frac{\nu}{2} g_{I\bar{J}} \psi_+^I D_{\bar{z}} \psi_+^{\bar{J}} + \frac{1}{4} R_{I\bar{J}K\bar{L}} \psi_+^I \psi_+^{\bar{J}} \psi_-^K \psi_-^{\bar{L}} \right)$$

$\phi: \Sigma \rightarrow X$ metric $g_{I\bar{J}}$

$TX =$ complexified tang. bund

$\psi_+^I: \text{section of } K^{1/2} \otimes \Phi^*(TX)$

$\epsilon_-: \text{hd section of } K^{-1/2}$

$\psi_-^{\bar{I}}: \bar{K}^{1/2} \otimes \Phi^*(TX)$

$\epsilon_+ \bar{K}^{-1/2}$

$K =$ canonical bundle on $\Sigma =$ bundle of $(1,0)$ forms
 $\bar{K} =$ anti " $(0,1)$

$$D_{\bar{z}} \psi_+^I = \frac{\partial}{\partial \bar{z}} \psi_+^I + \frac{\partial \phi^{\bar{J}}}{\partial \bar{z}} \Gamma_{\bar{J}K}^I \psi_+^K \quad \text{is the } \bar{\partial} \text{ operator}$$

invariant under $\delta \Phi^I = i(\epsilon_- \psi_+^I + \epsilon_+ \psi_-^{\bar{I}})$

$$\delta \psi_+^I = -\epsilon_- \partial_z \phi^I - i \epsilon_+ \psi_-^{\bar{K}} \Gamma_{\bar{K}M}^I \psi_+^M$$

$$\delta \psi_-^{\bar{I}} = -\epsilon_+ \partial_{\bar{z}} \phi^{\bar{I}} - i \epsilon_- \psi_+^{\bar{K}} \Gamma_{\bar{K}N}^{\bar{I}} \psi_-^{\bar{N}}$$

yields N=1 SCFT : T, G

$$T(z) T(w) = \frac{c}{2} \frac{1}{(z-w)^4} + \frac{2T(w)}{(z-w)^2} + \frac{\partial T(w)}{z-w} + \dots$$

$$T(z) G(w) = \frac{3}{2} \frac{G(w)}{(z-w)^2} + \frac{\partial G(w)}{z-w} + \dots$$

$$G(z) G(w) = \frac{2c \cdot 13}{(z-w)^3} + 2 \frac{T(w)}{z-w} + \dots$$

In the Ramond sector, G has a zero mode G_0 :

$$\{G_0, G_0\} = 2 \left(L_0 - \frac{c}{24} \right) \Rightarrow \Delta \geq c/24$$

2. $N=2$ \mathcal{N} model

Assume X is Kähler and choose local complex coordinates $\phi^i, \phi^{\bar{i}}$

Decompose $TX = T^{(0,1)}X \oplus T^{(1,0)}X$

$$L = 2t \int d^2z \left(\frac{1}{2} g_{I\bar{J}} \partial_z \phi^I \partial_{\bar{z}} \phi^{\bar{J}} + i \psi_-^{\bar{i}} D_z \psi_-^i g_{i\bar{i}} \right. \\ \left. + i \psi_+^{\bar{i}} D_{\bar{z}} \psi_+^i g_{i\bar{i}} + R_{i\bar{i}j\bar{j}} \psi_+^i \psi_+^{\bar{i}} \psi_-^j \psi_-^{\bar{j}} \right)$$

invariant under $\delta \phi^i = i(\alpha_- \psi_+^i + \alpha_+ \psi_-^i)$
 $\delta \phi^{\bar{i}} = i(\tilde{\alpha}_- \psi_+^{\bar{i}} + \tilde{\alpha}_+ \psi_-^{\bar{i}})$

where $\alpha_-, \tilde{\alpha}_-$ are sections of $K^{-1/2}$
 $\alpha_+, \tilde{\alpha}_+$ $\bar{K}^{-1/2}$

This realizes the $N=2$ SCFT: T, G^\pm, J
 $\Delta = 2, 3/2, 1$
 $Q = 0, \pm 1, 0$

$$G^+ G^- = \text{reg} = G^- G^+$$

$$G^+(z) G^-(w) = \frac{2c}{3} \frac{1}{(z-w)^3} + \left(\frac{2J}{(z-w)^2} + \frac{\partial J}{z-w} \right) + \frac{2T}{z-w} + \dots$$

$$J(z) J(w) = \frac{c/3}{(z-w)^2} + \dots$$

In the Ramond sector,

$$(G_0^+)^2 = (G_0^-)^2 = 0$$

$$\{G_0^+, G_0^-\} = 2 \left(L_0^{R^\pm} - \frac{c}{24} \right) = 2 \left(L_0^{NS} - \frac{1}{2} J_0^{NS} \right) \text{ by } (*)$$

The spectral flow relates NS and R sectors: (\Rightarrow) spacetime SUSY

$$J_0^{R^\pm} = J_0^{NS} \mp \frac{c}{6}$$

$$L_0^{R^\pm} = L_0^{NS} - \frac{1}{2} J_0^{NS} + \frac{c}{24} \quad (*)$$

Hence $L_0^{R^\pm} - \frac{c}{24} \geq 0$ implies $L_0^{NS} - \frac{1}{2} J_0^{NS} \geq 0 : \Delta \geq |q|/2$

Such chiral states form the chiral ring

Iterating the spectral flow, we find a map $NS \rightarrow NS$

$$J'_0 = J_0 \mp \frac{c}{3}$$

$$L'_0 = L_0 - J_0 + \frac{c}{6}$$

In particular, the NS ground state maps to a state with $(h, q) = (\frac{c}{6}, \mp \frac{c}{3})$. This is in fact the NS state of maximal charge $|q|$ in the chiral ring.

For a Calabi-Yau 3-fold, $c=9$ so we have

$$(0, 0)_{NS} ; \underbrace{(\frac{3}{8}, \pm \frac{3}{2})_R}_{\text{covariantly w/ spinor fields}} ; \underbrace{(\frac{3}{2}, \pm 3)_{NS}}_{\text{associated to } (3,0) \text{ form on CY}}$$

The spectral flow above is in fact known as a topological twist: the new spin in the R^\pm sector is $L_0^{NS} - \frac{1}{2} J_0^{NS}$

In particular ψ_+^i now transform as $\Phi^*(T^{1,0}X)$
 ψ_-^i $K \otimes \Phi^*(T^{0,1}X)$

If we perform the same twist on the right movers,

$$\begin{matrix} \psi_-^i \\ \psi_-^{\bar{i}} \end{matrix} \text{ transform as } \begin{matrix} \bar{K} \otimes \Phi^*(T^{1,0}X) \\ \Phi^*(T^{0,1}X) \end{matrix}$$

This is known as the A-twist.

If we perform a different twist, e.g

$$\begin{matrix} \psi_+^i & K \otimes T^{1,0} \\ \psi_+^{\bar{i}} & T^{0,1} \\ \psi_-^i & \bar{K} \otimes T^{1,0} \\ \psi_-^{\bar{i}} & T^{0,1} \end{matrix}$$

They agree with the origin model provided $K^{1/2}$ and $\bar{K}^{1/2}$ are trivial.

we obtain the B-twist.

These two twists correspond to different topological versions of the $N=2$ σ model.

Topological A-model

(Pavino 0406005)

Here ψ_+^i and $\psi_-^{\bar{i}}$ are ^{scalars} in $T^{1,0}X \oplus T^{0,1}X$, combined into $\chi \in TX$

$\psi_+^{\bar{i}}$ is a (1,0) form: $\psi_z^{\bar{i}}$

ψ_-^i is a (0,1) form: $\psi_{\bar{z}}^i$

The lagrangian becomes

$$L = 2t \int d^2z \left(\frac{1}{2} g_{I\bar{J}} \partial_z \phi^I \partial_{\bar{z}} \phi^{\bar{J}} + i \psi_z^{\bar{i}} D_{\bar{z}} \chi^i g_{i\bar{j}} + i \psi_{\bar{z}}^i D_z \chi^{\bar{i}} g_{i\bar{j}} - Ric_{j\bar{k}} \psi_z^i \psi_{\bar{z}}^{\bar{j}} \chi^j \chi^{\bar{k}} \right)$$

Invariant under

$$\{Q, \phi^I\} = \chi^I$$

$$\{Q, \chi^I\} = 0$$

$$\{Q, \psi_z^{\bar{i}}\} = i \partial_{\bar{z}} \phi^i - \chi^j \Gamma_{j\bar{k}}^i \psi_z^{\bar{k}}$$

$$\{Q, \psi_{\bar{z}}^i\} = i \partial_z \phi^{\bar{i}} - \chi^{\bar{j}} \Gamma_{\bar{j}k}^{\bar{i}} \psi_{\bar{z}}^k$$

$Q^2 = 0$; ghost number ϕ, χ, ψ, Q is classically conserved
0, 1, -1, 1

The lagrangian can be written as

$$\mathcal{L} = -i \{Q, V\} - t \int_{\Sigma_g} \Phi^*(J)$$

Where $\Phi^*(J) = i g_{i\bar{j}} d\phi^i \wedge d\phi^{\bar{j}}$ is the pull back of the Kähler class

$$V = t \int_{\Sigma_g} d^2z g_{i\bar{j}} \left(\psi_z^i \partial_z \phi^{\bar{j}} + \partial_z \phi^i \psi_{\bar{z}}^{\bar{j}} \right)$$

is the "gauge fermion"

It is natural to complexify the Kähler class by adding $-it \int \tilde{\Phi}^*(B)$

Remark that

- the energy momentum tensor on the world sheet is Q-exact

$$T_{\alpha\beta} = \{Q, b_{\alpha\beta}\} \quad b_{\alpha\beta} = \frac{\delta V}{\delta \bar{g}^{\alpha\beta}}, \quad \text{ghost\#} = -1$$

- Aside from the topological term $+it \int \phi^*(iJ + B)$, the partition function is independent of t , and of the α structure of X . In particular, the semi-classical approximation around $t \rightarrow \infty$ is exact: classical solutions are holomorphic maps

$$\partial_{\bar{z}} \phi^i = 0$$

$$\partial_z \bar{\phi}^{\bar{i}} = 0$$

↑ the fixed points of Q , really

- BRST symmetry ensures that the ratio of bos and ferm determinants is ± 1

The observables of the A-model are in 1-1 correspondence with the de Rham cohomology of X :

$$O_W = W_{I_1 \dots I_n} \chi^{I_1} \dots \chi^{I_n} \quad \text{ghost\#} = n$$

Indeed, $\{Q, O_W\} = -O_{dW}$

Due to an anomaly in the $U(1)$ ghost current, correlators vanish unless

$$\sum_{k=1}^g \text{deg}(\phi_k) = 2d(1-g) + 2 \underbrace{\int_{\Sigma} \phi^*(c_1(X))}_{0 \text{ for } CY}$$

Eg: for 03 folds,

- $g=0$: the only non-zero correlator is

$$\langle O_{W_1} O_{W_2} O_{W_3} \rangle = \int W_1 \wedge W_2 \wedge W_3 + \sum_{\beta \in [n; (S,)]} \int_{\beta} W_1 \int_{\beta} W_2 \int_{\beta} W_3 e^{-\sum t_i m_i}$$

- $g=1$: partition function is the elliptic genus

Topological B-model (requires $c_1(X) = 0$ for anomaly cancellation) ⑥

Parino 04/10/165

This time, ψ_+^i and ψ_-^i are scalars in $T^{0,1}$

while ψ_+^i and ψ_-^i are sections of $(K \oplus \bar{K}) \otimes T^{1,0}$

Define $\eta^{\bar{i}} = \psi_+^{\bar{i}} + \psi_-^{\bar{i}}$
 $\theta_i = g_{i\bar{i}} (\psi_+^{\bar{i}} - \psi_-^{\bar{i}})$

1-form ρ

The Lagrangian is now

$$L = t \int d^2z \left(g_{z\bar{z}} \partial_z \Phi^{\bar{i}} \partial_{\bar{z}} \Phi^i + i \eta^{\bar{i}} (D_z e_{\bar{z}}^i + D_{\bar{z}} e_z^i) g_{i\bar{i}} \right. \\ \left. + i \theta_i (D_{\bar{z}} e_z^i - D_z e_{\bar{z}}^i) \right. \\ \left. + R_{i\bar{i}j\bar{j}} e_z^i e_{\bar{z}}^{\bar{j}} \eta^{\bar{i}} \theta_{\bar{k}} g^{k\bar{j}} \right) \\ = it \int \{Q, V\} + tW$$

where the gauge fermion $V = g_{i\bar{i}} (e_{\bar{z}}^i \partial_z \Phi^{\bar{i}} + e_z^{\bar{i}} \partial_{\bar{z}} \Phi^i)$

and $W = \int -\theta_i \underset{\substack{\uparrow \\ \text{exterior derivative}}}{D} e^i - \frac{i}{2} R_{i\bar{i}j\bar{j}} e^i e^{\bar{j}} \eta^{\bar{i}} \theta_{\bar{k}} g^{k\bar{j}}$

not obvious to
 \uparrow true

Thus B is independent of the ω -structure of Σ and of the Kähler metric on X . It does however depend on the ω -structure of X

$$\begin{aligned} \{Q, \psi^i\} &= 0 \\ \{Q, \psi^{\bar{i}}\} &= \eta^{\bar{i}} \\ \{Q, \eta^{\bar{i}}\} &= 0 \\ \{Q, \theta_i\} &= 0 \\ \{Q, e^i\} &= i d\phi^i \end{aligned}$$

ghost #: $\eta \ \theta \ \rho \ \phi \ \ Q$

	η	θ	ρ	ϕ	Q
	1	1	-1	0	1

Again, the semi-classical approximation is exact: the t dependence of the tW term can be realised by $\Theta \rightarrow \Theta/t$, so the only t -dependence comes as an overall factor in correlation functions.

The classical solutions (fixed points of Q) are constant maps

$$d\phi^i = 0, \quad \eta^{\bar{i}} = 0$$

$$\eta^{\bar{i}} \Leftrightarrow \bar{\partial}$$

Observables: 1-1 correspondence with $\bigoplus_{p,q} H^p(X, \wedge^q T^{1,0}X)$

$$V = V_{\bar{i}_1 \dots \bar{i}_p}^{\bar{j}_1 \dots \bar{j}_q} d\bar{z}^{\bar{i}_1} \dots d\bar{z}^{\bar{i}_p} \frac{\partial}{\partial z^{\bar{j}_1}} \dots \frac{\partial}{\partial z^{\bar{j}_q}}$$

$$\text{Indeed, } \{Q, \theta_V\} = -Q\bar{\partial}V$$

so the BRST cohomology of Q consists of V such that $\bar{\partial}V = 0$
 $V \approx V + \bar{\partial}S$

Again, the anomalous ghost number conservation requires

$$\sum p_k = d(1-g)$$

$$\sum q_k = d(1-g)$$

The path integral reduces to an integral over X . For example, for CY 3-folds, $p=q=1$,

$H^1(X, T^{1,0}X) \approx H^{2,1}$: parametrises complex structures deformations

$$W_{\bar{j}}^{\bar{i}} \rightarrow \underbrace{\Omega_{i\bar{k}l}}_{\text{holomorphic 3-form}} W_{\bar{j}}^{\bar{i}} \in H^{2,1}$$

$$\langle \mathcal{O}_{W_1} \mathcal{O}_{W_2} \mathcal{O}_{W_3} \rangle = \int_X W_1^{\bar{i}_1}_{\bar{j}_1} W_2^{\bar{i}_2}_{\bar{j}_2} W_3^{\bar{i}_3}_{\bar{j}_3} \underbrace{\Omega_{i_1 i_2 i_3}}_{\wedge \Omega} d\bar{z}^{\bar{j}_1} d\bar{z}^{\bar{j}_2} d\bar{z}^{\bar{j}_3}$$

Special geometry gives this as $\frac{\partial^3 F_0}{\partial t_a \partial t_b \partial t_c}$

Coupling to topological gravity

Due to the issue of ghost number conservation, we have seen that only $\langle 000 \rangle$ sphere and $\langle 1 \rangle$ torus are non vanishing.

In order to compute higher genus correlators, it is necessary to couple the A or B topological matter to topological gravity. The answer is roughly as follows:

Recall that in bosonic string, genus g amplitudes are obtained by inserting the "antighost" b of reparametrization invariance, "folded" with Beltrami differentials:

$$F_g = \int_{\bar{\mathcal{M}}_g} \langle \prod_{k=1}^{6g-6} (b, \mu_k) \rangle \quad \mu_k \in H^1(\Sigma_g, T^1 \Sigma_g)$$

where $\bar{\mathcal{M}}_g$ is the "Deligne-Mumford" compactification of moduli space of complex structures of Σ_g , and

$$(b, \mu_k) = \int_{\Sigma_g} d^2z \left[b_{z\bar{z}} (\mu_k)_{\bar{z}}^z + b_{\bar{z}z} (\bar{\mu}_k)_z^{\bar{z}} \right]$$

Since b has ghost number -1 , this exactly compensates the background charge.

In topological string, G_- has spin 2 so can be identified with b :

$$\begin{array}{l} \text{bosonic string: } T_{\alpha\beta} = \{Q, b_{\alpha\beta}\} \\ \text{top string: } T = \{G_+, G_-\} \end{array} \quad ; \quad F_g^{A/B} = \int_{\bar{\mathcal{M}}_g} \prod_{a=1}^{3g-3} (\mu_a G_-) (\bar{\mu}_a \tilde{G}_\pm)$$

"Scattering" amplitudes at higher genus may be defined by including integrated vertex operators:

$$dO^{(0)} = \{Q, O^{(0)}\}$$

$$dO^{(1)} = \{Q, O^{(2)}\}$$

then $O^{(2)}$ has ghost number 0 if $O^{(0)}$ has # 2.

The target space theories are known as

- A: Idempotent Chern Simons
- B: Kodaira-Spencer

Gromov - Witten invariants & Gopakumar Vafa invariants

With this prescription for coupling to topological gravity, the vacuum amplitudes are now non trivial at arbitrary genus.

It is convenient to define a generating function a "string amplitude"

$$F_{\text{top}} = \sum_{g=0}^{+\infty} \lambda^{2g-2} F_g$$

where λ is known as the "topological string coupling". → $t = \text{fixed}, \bar{E} \rightarrow \infty$

For the A model on a CY three fold X , the (holomorphic) string amplitude has an asymptotic expansion

$$F_{\text{top}} = -i \frac{(2\pi)^3}{6\lambda^2} C_{ABC} t^A t^B t^C - \frac{i\pi}{12} C_{2A} t^A + F_{\text{GW}} \quad (*)$$

where t^A are the complexified Kähler moduli, on a basis^A of $H_{1,1}$

C_{ABC} are the triple intersection numbers of the dual 4-cycles

$C_{2A} = \int J_A C_2(T^{1,0}X)$ the second Chern classes

$$C_{ABC} = \int J_A \wedge J_B \wedge J_C \quad \downarrow \text{inv}'$$

The two first terms in (*) are perturbative, while F_{GW} contains the effect of worldsheet instantons at arbitrary genus:

$$F_{\text{GW}} = \sum_{g \geq 0; \beta} N_{g, \beta} e^{2\pi i \beta_A t^A} \lambda^{2g-2} \quad \beta = \text{effective curve}, \beta_A \geq 0$$

$$= \sum_{g \geq 0; \beta; d \geq 1} n_{g, \beta} \frac{1}{d} \left[2 \sin\left(\frac{d\lambda}{2}\right) \right]^{2g-2} e^{2\pi i d \beta_A t^A}$$

$N_{g, \beta}$ are rational-valued topological invariants of X known as Gromov-Witten

$n_{g, \beta}$ are the Gopakumar-Vafa invariants, believed to be integer: it counts the number of genus g holomorphic curves in the homology class $\beta_A \gamma^A$.

Note in particular that the leading term as $\lambda \rightarrow 0$ is

$$\frac{1}{\lambda^2} \sum_{\beta} N_{g,\beta} e^{2\pi i \beta_A t^A} = \sum_{\beta} n_{\beta}^0 \sum_{d=1}^{\infty} \frac{1}{d^3} \lambda^2 e^{2\pi i d \beta_A t^A}$$

which incorporates the effects of multiple coverings for an isolated genus 0 curve.

Note also that for $\beta=0$, the contributions are no longer exp. suppressed. They occur at genus 0 only, with $n_0^0 = -\frac{1}{2} \chi(X)$

They are given by the Mac-Nahan function

$$\begin{aligned} f(\lambda) &= \sum_{d=1}^{\infty} \frac{1}{d} \frac{1}{\left(\beta \sin^2 \frac{d\lambda}{2}\right)^2} \\ &= \sum_{d=1}^{\infty} \frac{1}{d} \left(\frac{i}{e^{i d \lambda / 2} - e^{-i d \lambda / 2}} \right)^2 = - \sum_{d=1}^{\infty} \frac{e^{i d \lambda}}{d (1 - e^{i d \lambda})^2} \\ &= \sum_{d=1}^{\infty} \sum_{n=1}^{\infty} \frac{n q^{nd}}{d} \quad \text{where } q = e^{i \lambda} \\ &= + \sum_n \log(1 - q^n) \end{aligned}$$

The latter form converges in the upper half plane $\text{Im} \lambda > 0$, and is suitable to study the large coupling limit $\lambda \rightarrow i\infty$.

In the weak coupling limit, the leading term is easy to extract,

$$f(\lambda) \sim \frac{1}{\lambda^2} \sum_{d=1}^{\infty} \frac{1}{d^3} = \frac{\zeta(3)}{\lambda^2} + \dots$$

This constant shift $-\frac{1}{2} \chi \zeta(3)$ of the tree level prepotential can be traced to the $\zeta(3) R^4$ tree-level term in the 10-dimensional action.

The subleading terms are more tricky to extract, but interesting too:

The Mac-Mahon function

$$f(z) = \sum_{n=1}^{\infty} n \log(1-q^n) \quad q = e^{i\pi z} = e^{-\pi t}$$

$$= - \sum_{n=1}^{\infty} \sum_{k=1}^{\infty} \frac{n e^{-nkt}}{k}$$

In order to analyze the behavior at $t \rightarrow 0$, we first consider its Mellin transform

$$Mf(s) = \int_0^{\infty} \frac{dt}{t^{1-s}} f(t)$$

$$= - \sum \frac{n}{k} (nk)^{-s} \Gamma(s)$$

$$= - \zeta(s-1) \zeta(s+1) \Gamma(s)$$

The original function may be obtained conversely by

$$f(t) = \frac{1}{2\pi i} \int_{\text{Re } s = \sigma} Mf(s) t^{-s}$$

where σ is chosen to lie to the right of any pole of $Mf(s)$.

Moving the contour to the left and crossing the poles generate the Laurent series expansion of $f(t)$.

Recall that $\Gamma(s)$ has poles at $s = -n$ $n=0, 1, \dots$, residue $\frac{(-1)^n}{n!}$

$\zeta(s)$ has a pole at $s=1$, $\zeta(s) \sim \frac{1}{s-1}$

and "trivial" zeros at $s = -2, -4, -6, \dots$

Thus:

-5	-4	-3	-2	-1	0	1	2	3	4	
										→ s
x	x	x	x	x	x					Γ(s)
		0		0			x			ζ(s-1)
		0			x					ζ(s+1)

x = pole
0 = zero

$$f(t) = \sum_{q=0}^{+\infty} \text{Residue}_{s=2-2q} \left[-\zeta(s-1) \zeta(s+1) \Gamma(s) t^{-s} \right]$$

At $s=2$: Residue = $-\zeta(3)/t^2$

$s=0$: double pole; residue = $-\zeta'(-1) - \frac{1}{12} \log t$

$s=2-2g$, $g \geq 2$: residue = $-\frac{\zeta(1-2g)\zeta(3-2g)}{(2g-2)!} t^{2g-2}$

use $\zeta(3-2g) = -\frac{B_{2g-2}}{2g-2}$ ($g \geq 2$)

to obtain the Laurent series expansion

$$f(t) = \frac{\zeta(3)}{t^2} + \frac{1}{12} \log(it/\lambda) - \zeta'(-1) + \sum_{g=2}^{\infty} \frac{t^{2g-2}}{(2g-2)!} \frac{B_{2g-2}}{2g-2} \cdot \frac{B_{2g}}{2g}$$

Impressively, this agrees with

$$\int \mathcal{M}_g c_{g-1}^3 = -\frac{B_{2g} B_{2g-2}}{2g (2g-2) \cdot (2g-2)!}$$

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Amusingly, the MacMahon function is also the partition function of three dimensional Young tableaux. This is the simplest occurrence of a general phenomenon which will not be covered here.

Holomorphic anomalies

Naively, a variation of the anti-holomorphic moduli \bar{E}^i corresponds to the insertion of a BRST trivial operator

$$\bar{\Phi}_i = \{G^+, [\bar{G}^+, \bar{\Phi}_i]\}$$

while a variation of the holomorphic moduli E^i would correspond to n -point functions:

$$C_{i_1 \dots i_n}^g = \partial_{i_1 \dots i_n}^n F_g$$

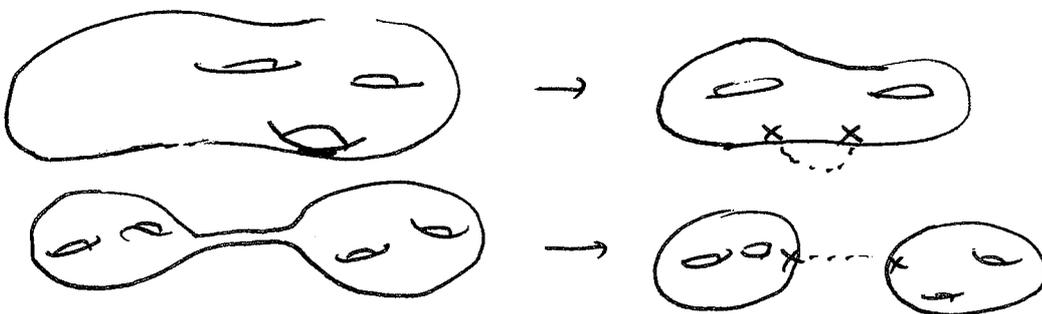
This however makes no sense, since F_g is a non-trivial section of \mathcal{L}^{2-2g} .

Both of these statements are corrected by taking into account boundary contributions in the integral over moduli of \mathcal{M}_g :

$$\begin{aligned} \frac{\partial}{\partial \bar{E}^i} F_g &= \int_{\mathcal{M}_g} \int d^2z \left\langle \oint_{C_z} G^+ \oint_{\bar{C}_z} \bar{G}^+ \bar{\Phi}_i(z) \prod_{a=1}^{3g-3} \int_{\bar{a}=1}^{3g-3} P_a G^- \prod_{\bar{a}=1}^{3g-3} \bar{P}_a \bar{G}^- \right\rangle \\ &= \sum_{b, \bar{b}=1}^{3g-3} \int_{\mathcal{M}_g} \int \bar{\Phi}_i \int 2\mu_b T \int 2\bar{\mu}_{\bar{b}} \bar{T} \prod_{\substack{a=1 \\ a \neq b}}^{3g-3} \left(\int P_a G^- \right) \prod_{\substack{\bar{a}=1 \\ \bar{a} \neq \bar{b}}}^{3g-3} \left(\int \bar{P}_a \bar{G}^- \right) \\ &= \sum_{b, \bar{b}=1}^{3g-3} \int_{\mathcal{M}_g} \frac{\partial}{\partial m_b} \frac{\partial}{\partial \bar{m}_{\bar{b}}} \left\langle \int \bar{\Phi}_i \prod_{\substack{a \neq b}} (\int P_a G^-) \prod_{\substack{\bar{a} \neq \bar{b}}} (\int \bar{P}_a \bar{G}^-) \right\rangle \end{aligned}$$

where we used $G^+(z)G^-(w) = \frac{2T}{z-w}$

There are 2 types of boundaries:



$$\bar{\partial}_i F_g = \frac{1}{2} \bar{C}_{ij\bar{k}} e^{2k} G^{j\bar{j}} G^{k\bar{k}} \left(D_j D_{\bar{k}} F_{g-1} + \sum_{r=1}^{g-1} (D_j F_r) (D_{\bar{k}} F_{g-r}) \right)$$

where $D_i = \partial_i - \frac{\Gamma_i}{\Gamma} - (2-2g)\partial_i K$

At genus 1, the holomorphic equation becomes 2nd order:

$$\partial_i \bar{\partial}_{\bar{j}} F_1 = \frac{1}{2} C_{ikl} C_{\bar{j}\bar{k}\bar{l}} e^{2K} G^{k\bar{l}} G^{l\bar{k}} - \left(\frac{\chi}{24} - 1\right) G_{i\bar{j}}$$

Similarly, one may derive anomaly equations for correlation functions.

This is best summarized by defining the generating function

$$W(\lambda, x; t, \bar{t}) = \sum_{g=0}^{\infty} \sum_{n=0}^{\infty} \frac{1}{n!} \lambda^{2g-2+n} C_{i_1 \dots i_n}^{(g)} x^{i_1} \dots x^{i_n} + \left(\frac{\chi}{24} - 1\right) \log \lambda$$

differs from BCOV by $x \rightarrow \lambda x$
 $\frac{\partial}{\partial x} \rightarrow \frac{1}{\lambda} \frac{\partial}{\partial x}$, $\frac{\partial}{\partial \lambda} \rightarrow \frac{\partial}{\partial \lambda} - \frac{x}{\lambda} \frac{\partial}{\partial x}$

where $C_{i_1 \dots i_n}^{(0)} = D_{i_1} \dots D_{i_{n-3}} C_{i_{n-2} i_{n-1} i_n}$, $n \geq 3$

$C_{i_1 \dots i_n}^{(g)} = D_{i_1} \dots D_{i_n} F^{(g)}$ $g \geq 1$

$= 0$ if $2g-2+n \leq 0$

so $W = \frac{1}{3!} C_{ijk}^{(0)} x^i x^j x^k + \frac{1}{4!} C_{ijkl}^{(0)} x^i x^j x^k x^l + \dots$
 $+ \left(\lambda C_i^{(1)} x^i + \frac{\lambda^2}{2} C_{ij}^{(1)} x^i x^j + \dots \right)$
 $+ \left(\lambda^2 F_2 + \lambda^3 C_i^{(2)} x^i + \dots \right)$
 $+ \left(\lambda^4 F_3 + \dots \right) + \left(\frac{\chi}{24} - 1\right) \log \lambda$

This satisfies

$$\left(\frac{\partial}{\partial \bar{t}^i} - \frac{1}{2} C_{\bar{i}\bar{j}\bar{k}} e^{2K} G^{j\bar{l}} G^{l\bar{k}} \frac{\partial^2}{\partial x^j \partial x^k} - \lambda^2 G_{i\bar{j}} x^j \frac{\partial}{\partial \lambda} \right) e^W = 0$$

$$\left(\frac{\partial}{\partial t^i} + P_{ij}^k x^j \frac{\partial}{\partial x^k} + \partial_i K \left(\lambda \frac{\partial}{\partial \lambda} - x \frac{\partial}{\partial x} + 1 - \frac{K}{24} \right) + \partial_i F_1 - \frac{\partial}{\partial \lambda} x^i + \frac{1}{2} C_{ijk} x^j x^k \right) e^W = 0$$

e^W is a section of the line bundle $L^{\frac{\chi}{24} - 1}$

By noting that the general solution of F_1 's anomaly equation is

$$F_1 = \frac{1}{2} \log \det G + \left(\frac{h+1}{2} + 1 - \frac{\chi}{24} \right) K + f_1 + \bar{f}_1$$

where $h = h_{2,1}$ (IIB) and f_1 is holomorphic,
 $h_{1,1}$ (IIA)

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and multiplying W by $\exp f_1$, one may rewrite the system as

$$\left[\begin{aligned} \partial_i \Psi_{\text{top}} &= \left[\frac{e^{2K}}{2} \bar{c}_{i\bar{j}} \bar{c}_{\bar{j}i} G^{\bar{j}\bar{k}} G^{k\bar{i}} \frac{\partial^2}{\partial x^i \partial x^{\bar{j}}} + G_{ij} x^{\bar{j}} \frac{\partial}{\partial \lambda^{-1}} \right] \Psi_{\text{top}} \\ \left(\nabla_i + \Gamma_{ij}^k x^{\bar{j}} \frac{\partial}{\partial x^k} \right) \Psi_{\text{top}} &= \left[\frac{1}{\lambda} \frac{\partial}{\partial x^i} - \frac{1}{2} \partial_i \log \det G - \frac{1}{2} x^{\bar{j}} x^k C_{ijk} \right] \Psi_{\text{top}} \end{aligned} \right]$$

where $\Psi_{\text{top}} = \exp(f_1) W$ is now a section of $\mathcal{L}^{\frac{h+1}{2}}$.

This suggests that $\Psi_{\text{top}}(\lambda, x; t, \bar{t})$ should be viewed as a wave function, obtained by quantizing a finite dimensional phase space $(\lambda, x; \bar{\lambda}, \bar{x})$ in a "complex" polarization specified by (t, \bar{t}) . The anomaly equations indicate the canonical transformation experienced by the wave function as the polarization is varied. In particular, the norm

$$\langle \Psi_{\text{top}} | \Psi_{\text{top}} \rangle = \int \frac{d\lambda d\bar{\lambda} d^h x d^h \bar{x}}{|\lambda|^4} e^{\frac{h+1}{2} K} |G|^{1/4}$$

$$\exp \left[-e^{-K} G_{i\bar{j}} x^i \bar{x}^{\bar{j}} + \frac{e^{-K}}{\lambda \bar{\lambda}} \right] \Psi_{\text{top}}(\lambda, x)_{t\bar{t}} \Psi_{\text{top}}^*(\bar{\lambda}, \bar{x})_{\bar{t}\bar{t}}$$

is independent of (t, \bar{t}) .

The two anomaly equations are conjugate with respect to this norm.

This is most easily understood in the B-model: the space $H^3(M, \mathbb{R})$ has a symplectic structure (15)

$$Q = \int_{CY} \delta\gamma \wedge \delta'\gamma$$

If one parametrizes $H^3(M, \mathbb{R})$ using real periods

$$\int_{A^I} \gamma = p^I \quad \text{then } Q = dp^I \wedge dq_I$$

$$\int_{B_J} \gamma = q_J$$

However, one could instead use the Hodge decomposition

$$\gamma = \frac{1}{2} \left(\frac{\Omega}{\lambda} + x^i \nabla_i \Omega + \bar{x}^{\bar{i}} \bar{\nabla}_{\bar{i}} \bar{\Omega} + \frac{\bar{\Omega}}{\bar{\lambda}} \right)$$

$$\in H^{3,0} \oplus H^{2,1} \oplus H^{1,2} \oplus H^{0,3}$$

The relation between the two basis is

$$p^I = \text{Re} \left(\frac{X^I}{\lambda} + x^i \nabla_i X^I \right) \quad \text{since } X^I = \int_{A^I} \Omega$$

$$q_I = \text{Re} \left(\frac{F_I}{\bar{\lambda}} + \bar{x}^{\bar{i}} \bar{\nabla}_{\bar{i}} F_I \right) \quad F_I = \int_{B_I} \bar{\Omega}$$

In this basis, the symplectic form becomes

$$Q = i e^{-K} G_{i\bar{j}} dx^i d\bar{x}^{\bar{j}} + i e^{-K} d\bar{\lambda}^{-1} \wedge d\lambda$$

Under a variation of the complex structure $\Omega(t, E)$, the parameters $1/\lambda, x^i$ can be adjusted by

$$\bar{\partial}_{\bar{i}} 1/\lambda = -G_{\bar{i}j} x^j$$

$$\bar{\partial}_{\bar{i}} x^k = e^K \bar{C}_{\bar{i}j\bar{k}} G^{k\bar{k}}$$

so that γ itself remains constant.

Quantum mechanically, the coherent state which diagonalizes x, \bar{x}^{-1} experiences a Boglioubov transformation

$$\bar{\partial}_{\bar{i}} |x, \lambda\rangle = \left(\frac{e^{2K}}{2} C_{\bar{i}j\bar{k}} G^{k\bar{k}} G^{i\bar{j}} \frac{\partial}{\partial x^i} \frac{\partial}{\partial \bar{x}^{\bar{k}}} - G_{\bar{i}j} x^j \frac{\partial}{\partial \bar{\lambda}^{-1}} \right) |x, \lambda\rangle$$

This is precisely the holomorphic anomaly equation,

provided one identifies

$$\Psi_{\text{top}}(x, \lambda; t, \bar{t}) = \langle \psi_{\text{top}} | x, \lambda \rangle$$

where $|\psi_{\text{top}}\rangle$ is a state in an abstract Hilbert space,

independent of t, \bar{t} .

In particular, we may choose to represent it in the real polarization,

where $q_I = \frac{\partial}{\partial p^I}$:

Define the "big phase space" variables $x^I = \frac{X^I}{\lambda} + x^i \nabla_i X^I$

so $x^I \sim (\lambda, x^i)$

The relation with real basis is now $\begin{cases} \text{Re}(x^I) = p^I \\ \text{Re}(\tau_{IJ} x^J) = q_I \end{cases}$

which is solved by $x^I = -i (\text{Im} \tau^{-1})^{IJ} (q_J - \bar{t}_{JK} p^K)$

Quantum mechanically, $\Psi_{\text{top}}(x^I)$ is related to $\Psi_{\text{top}}(p)$ by a Bergmann kernel

$$\Psi_{\text{top}}(x^I; t, \bar{t}) = \int dp \Psi_{\text{top}}(p) \exp \left[\frac{i}{2} \left(p^I \bar{t}_{IJ} p^J + p^I (t - \bar{t}) X^J - \frac{1}{4} (t - \bar{t}) X^I X^J \right) \right]$$

There is an interesting dynamical system leading to a phase space $H_3(\Pi, \mathbb{R})$ the NS5-brane in IIA on CY - or equivalently the NS-brane in M theory.

Since it carries a self dual 3-form flux, its phase space is $H_3(\Pi, \mathbb{R})$.

Its partition function may be obtained by factorizing that of a general 3-form flux:

In this set up, p^I is viewed as the flux on the NS5 brane, and x^I are the background values of the 3 form. The full partition function appears to involve a theta series over p^I for the classical part, as the infinite series of topological amplitudes $\exp \left[\sum_{g \geq 1} F_g(t, \bar{t}) \lambda^{2g-2} \right]$ for the quantum part

As we shall discuss later, there is another natural set-up where the topological string amplitude seems to play the rôle of a wave function: radial quantization of ⁸⁸ black holes, or equivalently, the quantum attractor flow.



Before closing this chapter, one final observation: since $\Psi_{\text{top}}(x^{\pm}; t, \bar{E})$ is interpreted as a wave function, it follows that under electric magnetic duality $x^{\pm} \rightarrow F_{\pm}$, $F_{\pm} \rightarrow -x^{\pm}$, it transforms by Fourier under all variables x^{\pm} .

This is reminiscent of Kontsevich's "Very Wild Ours" Arbeitstagung, 91

$$\text{Fourier}(\exp F_0) \stackrel{?}{=} \exp(\text{Legendre } F_0)$$

This "Master equation" is at least verified by all the very special supergravities, where

Kazhdan, Etingov, Poinchelnik

$$F_0 = \frac{N(X)}{X^0}$$

$N(X)$ = norm of a degree 3 Jordan algebra